## STRESS ANALYSIS OF POROELASTIC SEABED

## SLIOPES UNDER WAVE LOADING USING THE

 BOUNDARY ELEMENT METHOD
## CENTRE FOR NEWFOUNDLAND STUDIES

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# STRESS ANALYSIS OF POROELASTIC SEABED SLOPES UNDER WAVE LOADING USING THE BOUNDARY ELEMENT METHOD 

## by

(C) Wayne W. Raman-Nair, B.Sc. (Math.), B. Sc. (Eng.), M. Eng.

> A. thesis submitted to the School of Graduate Studies in partial fulfillment of the requirements for the degree of Doctor of Philoso-. $y$

Faculty of Engineering and Applied Science Memorial University of Newfoundland October 1990

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#### Abstract

The effect of wave forces on the stability of a sloping seabed is investigated using linear wave theory and Biot's theory of poroelasticity. The wave forces on the slope, the wave-induced effective stresses and pore pressures are all computed by the boundary element method. Full derivations are presented for the boundary integral equations and fundamental solutions of the poroelastic theory for the case of sinusoidal loading conditions. The wave-induced stress field is combined with the in-situ stress field and the Mohr-Coulomb failure criterion is used to determine the zone of incipient failure. It is found that wave induced stresses must be analysed throughout an entire wave cycle in order to assess the likelihood of failure and also that the stability of a slope under waves is strongly dependent on the initial stress distribution.


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## List of Symbols

## Latin symbols

| $a$ | parameter dependent on soil properties and wave frequency (equation 5.81) |
| :---: | :---: |
| $c$ | parameter related to $a$ by a multiplicative constant (equation 6.15) |
| $D_{\alpha k j}$ | kernels for the evaluation of interior stresses |
| e | dilatation of solid skeleton |
| $\mathbf{e}_{\boldsymbol{i j}}$ | strain tensor of solid skeleton |
| $f$ | porosity of soil |
| $F_{i j}$ | sum of all inter-granular forces vector on an area $\delta A$ of soil normal to the $j$ axis |
| $g$ | acceleration due to gravity |
| $G$ | shear modulus of soil skeleton |
| $G_{i k}$ | part of the fundamental solution $\tilde{\bar{u}}_{i}^{* k}$ |
| $h_{0}$ | water depth at slope base |
| $h_{0}^{\prime}$ | water depth at slope top |
| H | wave height; an elastic modulus of soil (equation 5.52) |
| $H_{i k}$ | part of the fundamental solution $\tilde{\bar{u}}_{i}^{* *}$ |
| $i$ | $\sqrt{-1}$ |
| $k$ | soil parameter (equation 3.21); wave number |
| $k_{0}$ | permeability of soil in $\mathrm{ms}^{-1}$ |
| $K$ | conjugate stress ratio |
| $K_{0}$ | coefficient of lateral earth pressure at rest |
| $K_{0}, K_{1}, K_{2}$ | modified Bessel functions of the second kind of orders $0,1,2$ respectively |
| $K_{6}$ | bulk modulus of soil skeleton |
| $K_{f}$ | bulk modulus of water |
| $K_{r}$ | bulk modulus of soil grains |
| $l 0$ | soil depth at slope base |
| $l_{0}^{\prime}$ | soil depth at slope top |
| $L$ | wavelength |
| M | soil parameter (equation 3.2) |
| $n_{i}$ | unit outward normal to a curve or surface |
| $p$ | pore pressure |
| $\bar{p}$ | frequency independent part of pore pressure |
| $\bar{p}^{0} \cdot \vec{p}^{01}, \bar{p}^{*}$ | fundamental solutions |
| $\underline{p}, P$ | pesition vector of a point $P$ |


| $P_{0}$ | amplitude of wave pressure at the mudline |
| :---: | :---: |
| $P_{\text {ik }}$ | part of the fundamental solution $\tilde{\bar{T}}_{\mathbf{i}}^{* k}$ |
| $Q$ | soil parameter (equation 3.14) |
| $Q_{i k}$ | part of the fundamental solution $\overline{\bar{T}}_{\boldsymbol{i}}^{* k}$ |
| $r$ | distance from node $j$ to an arbitrary point |
| $R$ | soil parameter (equation 3.14) |
| $S_{\alpha k j}$ | kernels for the evaluation of interior stresses |
| $t$ | time |
| $T_{i}$ for $i=1,2$ | surface tractions |
| $T_{3}$ | $=-p$ |
| $\bar{T}_{i}$ for $i=1,2,3$ | frequency independent part of $T_{i}$ |
| $\tilde{T}_{i}(i=1,2,3)$ | dimensionless form of $\bar{T}_{i}$ |
| $\bar{T}_{i}^{e}, \tilde{T}_{i}^{e}$ | nodal values of $\bar{T}_{i}, \tilde{\bar{T}}_{i}$ |
| $\bar{T}_{i}^{0}, \bar{T}_{i}^{01}, \bar{T}_{i}^{*}, \tilde{T}_{i}^{*}$ | fundamental solutions |
| $\underline{u}, u_{i}(i=1,2)$ | displacement vector of the soil skeleton |
|  | $=f\left(U_{i}-u_{i}\right)$ |
| $\bar{u}_{i}$ for $i=1,2,3$ | frequency independent part of $u_{i}$ |
| $\tilde{\bar{u}}_{i}(i=1,2,3)$ | dimensionless form of $\bar{u}_{i}$ |
| $\bar{u}_{i}^{e}, \tilde{\bar{u}}_{i}^{e}$ | nodal values of $\bar{u}_{i}, \tilde{\bar{u}}_{i}$ |
| $\bar{u}_{i}^{0}, \bar{u}_{i}^{01}, \bar{u}_{i}^{*}, \bar{u}_{i}^{* k}$ | fundamental solutions |
| $\underline{U}, U_{i}(i=1,2)$ | displacement vector of the pore water |
| $\bar{U}_{i}$ for $i=1,2$ | frequency independent part of $U_{i}$ |
| $\bar{U}_{i}^{0}, \widehat{U}_{i}^{01}, \bar{U}_{i}^{*}, \tilde{U}_{i}^{* k}$ | fundamental solutions |
| $\bar{v}_{n}$ | $=\bar{u}_{3}$ |
| $\bar{v}_{n}^{01}, \bar{v}_{n}^{* k}$ | fundamental solutions |
| $X, X_{i}$ | body force per unit mass |
| $Z_{i k}$ | part of the fundamental solution $\tilde{\bar{T}}_{\mathbf{i}}{ }^{*}$ |

## Greek symbols

soil parameter (equation 3.2)
slope angle in degrees; function of $\mu$ and $\nu$ (equation 5.63)
dry unit weight of soil
buoyant unit weight of soil
$\Gamma$
$\Gamma_{e}$
$\Delta$
$\delta_{i j}$
$\varepsilon$
$\eta_{0}$
$\theta$
$\lambda$
$\lambda_{c}$
$\mu$
$\nu$
$\xi$
$\rho$
$\rho_{s}$
$\rho_{f}$
$\sigma=-f p$
$\sigma_{i j}$
$\sigma_{i j}^{k}, \sigma_{0}$
$\sigma_{x}^{\prime}, \sigma_{z}^{\prime}$
$\tau_{i j}^{\prime}$
$\tau_{i j}$
$\tau_{x z}$
$\Phi$
$\phi$
$\phi^{\prime \prime}$
$\phi^{\prime}$

$\omega$
$\Omega$
boundary of soil region
boundary segment of soil region
Dirac delta function
Kronecker delta
fluid dilatation
incident wave amplitude
stress angle (degrees)
Lame constant of soil skeleton
soil parameter (equation 3.2)
shear modulus of soil skeleton
Poisson's ratio of soil skeleton
local coordinate on boundary element
density of soil (bulk)
density of soil grains
density of pore water
fluid part of total stress tensor (equation 3.1.1)
solid part of total stress tensor (equation 3.11)
fundamental solutions
$x, z$ components of effective stress
total stress tensor
effective stress tensor
(effective) shear stress
wave velocity potential
frequency independent part of wave velocity potential
fundamental solution (potential problem)
angle of internal friction of soil in degrees
(effective stress parameter)
wave frequency (circular)
a region of soil or water

## Chapter 1

## INTRODUCTION

Engineering interest in the stability of the seafloor has been motivated by the presence of petroleum beneath the submarine soil. The necessity of locating bottommounted structures, pipelines and cables on or in the vicinity of seabed slopes has made the analysis of submarine slope stability a matter of great importance. One of the factors contributing to submarine slope failures is the effect of the loading due to water waves. A well known example is the failure of two Shell jacket pile platforms in the Gulf of Mexico during Hurricane Camille (1969). Other areas susceptible to wave-induced slope instabilities are the Gulf of Alaska and the Nova Scotian shelf. Although many researchers (including the author) have presented stability analyses of a flat seabed under wave loading, few have attempted comprehensive theoretical studies of seabed slopes. It has been long recognised that there is a need for a stress analysis approach to problems of slope stability to supplement existing limit equilibrium analyses. In performing such a study the important first step is the choice of a soil model. Unfortunately, there is at present no soil model
which by itself completely describes the complex behaviour of soil. In choosing a model one must therefore identify those aspects of soil behaviour which are most pertinent to the problem being addressed. Several researchers (Yamamoto et al., 1978-1985; Mynett and Mei, 1982; Zienkiewicz et al., 1984; Cheng and Liu, 1986) have pointed out that in determining the seabed response to wave loading, the coupling between the soil skeleton and pore water is of primary significance. In light of this fact, Biot's poroelastic model will be used as the basis for the theoretical analysis. This model has been experimentally verified for the wave loading of sand beds (Yamamoto, 1978 ; Cheng and Liu, 1986). In order to determine the wave forces on the slope we shall use linear (Airy) wave theory and the boundary element method. The problem of wave propagation over a slope has been studied analytically by Stoker (1957) for certain values of the slope angle. For our purposes it is necessary to determine the wave forces in an arbitrary slope, and to this end the boundary element method furnishes a quick and efficient procedure. The boundary conditions on the surface of the slope are thus provided in a natural way.

We distinguish between two features of soil response under wave loading. In the terminology of Finn et al., (1983) these are due to "transient" and "residual" stresses respectively. The "transient" stresses and pore pressures are the instantaneous soil response to the passing wave as determined by the constitutive laws and equations of motion. This response vanishes when the wave passes. The "residual" stresses and pore pressures are due to the cumulative effects of the passage of several waves. These induce permanent strains and "residual" pore pressures which depend on the relative rates of generation and dissipation. In this thesis only the problem of "transient" instability will be addressed. We would emphasise that
although the term "transient" is used, we actually determine the steady-state solution to the governing equations under sinusoidal loading. Biot's po: welastic model is thus adopted in its basic form i.e. with constant soil properties. Even with these simplifications the problem is intractable by analytic methods and we employ the boundary element method (BEM) to compute the wave-induced effective stresses and pore pressures. These are combined with the in-situ stresses and the MohrCoulomb failure criterion is used to determine the incipient zone of failure. The manner in which such failure progresses is a matter which requires further research and the methods presented herein may be used as a starting point in this regard.

## Chapter 2

## LITERATURE SURVEY

We have already mentioned that in the analysis of seabed stability it is important to choose an appropriate soil model. We review some of the approaches taken by various authors for modelling the saturated submarine sediments.

### 2.1 Soil Models

Some of the early attempts at soil modelling (Putnam, 1949; Sleath, 1970; Moshagen and Torum, 1975) neglected soil deformation under loading and thus do not provide a complete model. Oner and Janbu (1975) presented an analysis of the seabed under offshore storage tanks, in which the soil is modelled by springs and dashpots. Bell et al. (1976) performed both stochastic and deterministic analyses assuming homogeneous, isotropic, and elastic soil. Moshagen and Monkmeyer (1979) have used potential theory to describe the flow within the seabed (as did Putnam and Sleath). The soil is assumed to be homogeneous, isotropic, coarse-
grained and rigid. They have also computed the wave-induced pressure distribution on the surface of a vertical embedded cylinder, as well as the total horizontal force and overturning moment generated by the hydrodynamic seepage pressure. O'Donnell (1982) has determined the wave-induced pore pressure fields around a variety of structures (buried pipelines, breakwaters, seawalls, sheetpile bulkhead, offshore foundation) using a potential theory and the Boundary Integral Equation Method. Potential theory has also been used by Lai et al. (1975) and Macpherson (1978) to determine the wave-induced pore pressure around a buried pipeline, and by Liu (1985) to compute the wave-induced pore pressure under a gravity structure. The latter considers the soil skeleton and pore water to be incompressible and he has employed the Riemann-Hilbert technique to compute the pore pressures. Rahman, Seed and Booker (1977) have computed the wave-induced pore pressure history under an axisymmetric tank assuming Darcy's Law and by incorporating a mechanism for pore-pressure generation and dissipation. Munro et al. (1985) have used an elasto-plastic soil model that includes foundation stratigraphy and embedment of an offshore structure. They have developed equivalent spring stiffnesses of soil for vertical, horizontal and moment loading. An elasto-plastic soil model has also been suggested by Prevost et al. (1980). This model incorporates both drained and undrained behavior. It describes the anisotropic, elastoplastic, path dependent, non-linear stress-strain-strength properties of inviscid saturated soils. It is assumed that the elasticity of the material is linear and isotropic and that non-linearity and anisotropy result from its plasticity.

Potts and Windle (1985) did a numerical study of the foundation behavior of a gravity platform using concepts from critical state soil mechanics (Drammen clay
model). Meimon and Lassoudiere (1985) have proposed an elastoplastic model with multiple yield surfaces and kinematical hardening, and involving eleven material constants. Zienkiewicz et al. (1982) have also adopted an elastoplastic model for describing soil behavior. Bouckovalas et al. (1984) have used an empirical model describing the cumulative effects of cyclic loading of sands. He draws upon an analogy between accumulation of strain during cyclic loading and viscoelastic creep, and uses equations and rules derived from the behavior of actual sands. A viscoelastic model was adopted by Schapery and Dunlap (1978) who, in addition, considered non-linear soil properties and the variation of soil properties with depth. Using the linear dynamic theory by Biot (1965), Schapery and Dunlap solved the elastic-viscous problem of wave-seabed interaction. Kraft et al. (1985) have compared this model with field data and have observed good agreement.

While the above approaches have their individual merits, they do not account for an important feature of wave-soil interaction, viz. the coupled responses of the soil 'skeleton' and pore water. Such behaviour must also be considered in problems of soil consolidation. The earliest attempt at modelling soil consolidation is due to Terzaghi. In this theory the effective stresses (in the soil skeleton) and the pore pressure are regarded as autonomous stress systems linked only through the general equations of equilibrium (Zaretskii, 1972). It has been generally recognised that a more accurate picture of the coupled interaction between the solid and fluid phases is provided by the Biot theory of three dimensional consolidation, first presented by M.A. Biot in 1941 for quasi-static phenomena and later extended by him in a series of papers to include soil anisotropy and visco-elasticity (Biot, 1955, 1956, 1963). A general poroelastic theory including dynamical terms was presented in

1962 (Biot,1962). In Biot's formulation the interaction between the soil skeleton and pore water is represented through body forces. The coupling between soil skeleton and pore water is also considered in a 'mixture' theory by Katsube and Carroll (1987). This theory reduces to Biot's when certain terms involving velocity gradient are suppressed. These terms account for part of the shear interaction between the solid and fluid constituents. However, as pointed out by Katsube and Carroll (1987), these terms are unimportant for most practical applications requiring the solution of steady state boundary value problems, and Biot's model does in fact account for the most significant shear interaction effects (which are due to fluid viscosity).

### 2.2 The Poroelastic Model of Biot

This theory has been used by a number of athors for analysing seabed stability. The first appears to be Yamamoto et al. (1978) and Madsen (1978) who developed an analytical solution for wave-induced effective stresses and pore pressures in homogeneous isotropic seabeds. They have analysed beds of infinite depth, while Yamamoto (1978) has considered beds of finite depth. Yamamoto et al. (1978) have substantiated their theoretical findings for sand beds by extensive laboratory experiments. Subsequently, however, Clukey et al. (1984) and Davies (1985) have indicated discrepancies between their experimental results and Yamamoto's theoretical findings. They suggest the incorporation of additional features in the Biot model, among them hydraulic anisotropy (as done by Madsen, 1978), layering and damping. Yamamoto et al. $(1983,1985)$ introduced into the Biot model the effects
of Coulomb damping and non-linear shear modulus via the Hardin-Drnevich (1972) formulas and have obtained good correlation with experiments.

Several other authors have adopted the Biot poroelastic model. Spierenburg (1985) has used the analytic solutions of Yamamoto to determine the wave-induced forces on a buried pipeline. He has also used these solutions together with a one dimensional version of Biot's pore pressure equation to develop a model for pore pressure generation. Silvestri et al. (1985) have included the effect of soil anisotropy and found that the soil is closer to failure than when considered isotropic. Simon, Zienkiewicz and Paul (1984) have developed an analytical solution to Biot's equations for the transient response of a one-dimensional column of fluid-saturated poroelastic solid. Mei (1982) has solved Biot's equations ising a boundary layer approximation method, his argument being that relative motion between pore water and soil skeleton is significant only within a boundary layer at the mudline. He has considered the problem of waves propagating uver a horizontal seabed, waves passing over a pipe laid on the seabed, waves normally incident on a long caisson, and the lifting of a large object from the seafloor. Kokkinowrachos (1985) has determined the wave-induced pore pressure below an offshore structure resting directly on the seabed or embedded in the soil. The two-dimensional problem is considered and both the cross-section of the structure and the contour of the seabottom can be of arbitrary shape. He has used the macroelement approach to solve Biot's equations. In this method the cross-section of the structure and the seabed contour are approximated by step curves, and the flow field around the structure and in the soil is subdivided into macro-elements. The method can be extended for stratified and inelastic soil. Mynett and Mei (1982) have considered the wave-
induced stresses in a saturated, homogeneous, isotropic poro-elastic seabed beneath a rectangular caisson. They have analyzed the two-dimensional problem and have used Mei's boundary layer approximation in conjunction with Biot's equations. A complex variable technique is used to achieve an analytic solution for stresses and pore pressure. Barends and Calle (1985) have used the solution of Yamenoto et al. (1978) to determine the response of the seabed to wave loading. They then establish a procedure for the assessment of cyclic pore pressure build-up and liquetaction of seabeds under random wave loading using a one-dimensional equation for pore pressure. Finn et al. (1983) have used Biot's equations for the transient response of the seabed, and the Seed-Rahman model of pore pressure generation and dissipation (Seed and Rahman, 1978) to assess residual pore pressures. Ishihara and Yamazaki (1984) have used the Yamamoto-Madsen solution of Biot's equations and have done a liquefaction analysis in terms of cyclic stress :atio. Okusa (1985) has used Biot's equations to analyze the wave-induced stresses in unsaturated soils. Mei and McTigue (1984) have performed an analysis of a submarine ridge and canyon under wave loading using the Biot theory. Sabin (1989) has presented analytic formulae for effective stresses and pore pressures in a poroelastic seabed under normal and shear loads. The author's M.Eng thesis (1985) deals with the analytic solution of Biot's equations for wave loading for both the dynamic (i.e. inertia terms included) and quasi-static cases. A homogeneous, isotropic seabed is assumed. It is shown that the response of sandbeds is essentially quasi-static, i.e. the inertia terms in the governing equations may be ignored.

For many problems an analytic approach is either too difficult or impossible and it is necessary to resort to numerical methods. The finite element method (FEM)
has been used by a number of authors to solve Biot's equations for analyzing various problems of soil consolidation and flow through porous media (Sandhu and Wilson,1969 ; Yokoo et al.,1971 ; Ghaboussi and Wilson, 1973 ; Smith and Hobbs,1976). Zienkiewicz and Shiomi (1984) have used the FEM to solve Biot's equations and have introduced various approximations for 'fast' and 'slow' phenomena and for the cases of compressible and incompressible fluid. In recent years the boundary element method (BEM) has emerged as a significant challenger to the finite element method and has been successfully applied in many areas of solid and fluid mechanics. One of the advantages of the BEM is that the field equations in the region of interest are transformed into integral equations on the boundary of the region, thus reducing the dimension of the problem by one. Only the boundary of the region needs to be discretised and the user has control over the number of interior points at which the solution of the governing equations are found. This contrasts with the FEM which requires domain discretisation and necessarily produces solutions at all interior nodes. Further, for many problems, the BEM gives greater accuracy than the FEM for the same level of discretisation (Mukherjee et al.,1984). The BEM is still undergoing development and it is certain that further research will extend the power and range of applicability of the method. Kuroki et al.(1982) and Aramaki et al.(1985) have used the BEM in conjunction with Biot's theory but only after uncoupling the soil deformation and pore pressure equations. The fully coupled equations were solved using the BEM by Cheng and Liggett (1984). They have assumed homogeneous, isotropic soil and constant soil properties and they have used Laplace transformed space to deal with time dependent problems. Cheng and Liu (1986) have used the boundary element method and the Biot theory to deter-
mine wave induced seepage forces on a buried pipeline. They have obtained good agreement with experimental results. They have also investigated the sensitivity of the pore pressure response to several soil and fluid parameters. A boundary element formulation for Biot's equations including viscoelastic behaviour of the soil skeleton has been presented by Predeleanu (1981).

### 2.3 Slope Stability under Waves

The stability of underwater slopes was discussed by Terzaghi (1956) who pointed out that slope failures occur when the average shearing stress on the potential surface of sliding becomes equal to the average shearing resistance along this surface. One of the first analyses of slope stability under waves is due to Henkel (1970). He presented a total stress analysis (i.e. neglecting pore pressures) based on the principle of limiting equilibrium and the assumption of a circular failure surface. He used a standing wave as the loading on the slope. Henkel used his model to show that the overturning moment produced by large waves on a sloping bed in the Gulf of Mexico could exceed the resisting moment provided by the soft sediments. Bea (1971) used the same model to investigate the failures of the Shell platforms in the Gulf of Mexico during Hurricane Camille in 1969. Rahman et al. (1985) have developed a probabilistic analysis for slope stability in which the waves are considered as a random process and the undrained shear strength of sediments is treated as a random variable. In this paper Henkel's approach is extended to develop a method for evaluating the probability of soil failure. Mitchell et al. (1972) proposed that slope failures were caused by strength reductions due to wave
remoulding, rather than additional driving moments or shear stresses. An effective stress method of slices for the stability of seabed slopes has been presented by Finn and Lee (1979), which includes both wave and earth.quake loading. Wright (1976) developed a finite element model which included gravity stresses and utilized a hyperbolic stress-strain relationship. Wave forces were estimated from linear wave theory using a constant water depth. The major limitation of the above studies is that the interaction between soil skeleton and pore water is not taken into account. We therefore choose the poroelastic model for our present study. As mentioned previously, Mei and McTigue (1984) have used Biot's poroelastic model to analyse the effects of wave loading on a gentle slope. They use an analytic solution valid for flat beds with a modified wave pressure which approximates the slow variation in the wave forces over a gentle siope. For certain values of the slope angle the analytic techniques of Stoker (1957) may be used to determine the wave forces on the slope, but for an arbitrary slope numerical methods are required. Alliney (1981) has used linear wave theory and the BEM to determine the wave velocity potential over an arbitrary coastal planar slope. In this thesis we determine (via the BEM ) the wave forces on an arbitrary planar slope in the open ocean. The major difference between these two problems is that the latter has an additional boundary for which appropriate conditions must be specified. We have defined the side boundary conditions using the techniques illustrated by Alliney. The waveinduced stress field in the sloping bed is also computed by the BEM using Biot's poroelastic model. To provide boundary conditions on the side boundaries of the poroelastic medium we use an analytic solution for flat beds.

## Chapter 3

## Theoretical Background

The theory of poroelasticity presented by M.A. Biot in 1962 will be used to model the soil medium. Here we present the final equations with an explanation of the notation.

### 3.1 The Biot Soil Model (Theory of Poroelasticity)

The saturated soil is modelled as a solid, porous, deformable skeleton, the pores being completely filled with water. We assume thet the soil is homogeneous and isotropic. The Cartesian coordinates are denoted by $x_{1}, x_{2}$ and $x_{3}$. Subscripts 1, 2, and 3 refer to the Cartesian coordinate directions, unless otherwise specified. A repeated suffix indicates summation with respect to that suffix over the range $1,2,3$ unless otherwise specified. A comma followed by a suffix indicates partial
differentation with respect to the appropriate coordinate direction, e.g. $A_{i} \equiv \frac{\partial A}{\partial x_{i}}$. We let $u=\left(u_{1}, u_{2}\right.$, , $\left.u_{3}\right)$ be the displacement of the solid matrix at any point; $U=$ ( $U_{1}, U_{2}, U_{3}$ ) be the displacement of the fluid at any point; and $f$ be the porosity of the soil. We denote by $\tau_{i j}$ the total stress components on the bulk material; $\tau_{i j}$ is a $j$-direction stress acting on a plane normal to the $i$-axis. The pore pressure is denoted by $p$. For the present, we use the usual sign convention adopted in the theory of elasticity, i.e.;
(a) tensile stresses are positive and,
(b) a shear stress is reckoned positive when acting in a positive coordinate direction on a plane whose outward normal points in a positive coordinate direction; or when acting in a negative coordinate direction on a plane whose outward normal points in a negative coordinate direction. We defer until a later section the use of the conventional soil mechanics sign convention, which is exactly opposite to that described above.

We now present the constitutive laws as (Biot, 1962)

$$
\begin{align*}
& \tau_{i j}=2 \mu e_{i j}+\left(\lambda_{\mathrm{c}} e-\alpha M \zeta\right) \delta_{i j}  \tag{a}\\
& p=-\alpha M e+M \zeta \tag{b}
\end{align*}
$$

where

$$
\begin{aligned}
& e_{i j}=\frac{1}{2}\left(u_{i, j}+u_{j, i}\right) \\
& e=e_{k k} \\
& \zeta=-f\left(U_{k}-u_{k}\right)_{, k}
\end{aligned}
$$

$\delta_{i j}$ is the Kronecker delta.
The constant $\mu$ is the shear modulus of the soil skeleton. The other constants $\lambda_{c}, \alpha$
and $M$ appearing in (3.1) may be expressed in terms of the following familiar elastic moduli :
$K_{r}=$ bulk modulus of soil grain
$K_{b}=$ bulk modulus of soil skeleton
$K_{f}=$ bulk modulus of water
It can be shown that (Biot and Willis, 1957; Stoll,1974)

$$
\begin{align*}
\lambda_{c} & =\frac{\left(K_{r}-K_{b}\right)^{2}}{D-K_{b}}+K_{b}-\frac{2 \mu}{3} \\
\alpha & =1-\frac{K_{b}}{K_{r}}  \tag{3.2}\\
M & =\frac{K_{r}^{2}}{D-K_{b}}
\end{align*}
$$

where

$$
D=K_{r}\left[1+f\left(\frac{K_{r}}{K_{f}}-1\right)\right]
$$

It is of interest to examine the relationship between (3.1) and Hooke's Law. From equation (3.1) (b) we have

$$
\zeta=\frac{1}{M}(p+\alpha M e)
$$

Substituting this into (3.1) (a) gives

$$
\begin{equation*}
\tau_{i j}=2 \mu e_{i j}+\delta_{i j}\left[\left(\lambda_{c}-\alpha^{2} M\right) e-\alpha p\right] \tag{3.3}
\end{equation*}
$$

From (3.2) we find that

$$
\begin{equation*}
\lambda_{c}-\alpha^{2} M=K_{b}-\frac{2 \mu}{3} \equiv \lambda \tag{3.4}
\end{equation*}
$$

Here, $\lambda$ is the standard Lamé constant of the theory of elasticity. Equation (3.3) may now be re-written as:

$$
\begin{equation*}
\tau_{i j}=2 \mu e_{i j}+\delta_{i j} \lambda e-\delta_{i j} \alpha p \tag{3.5}
\end{equation*}
$$

We shall show that this reduces to Hooke's Law when written in terms of effective stress. To define the effective stress tensor, we consider a plane area $\delta A$ of saturated soil normal to the $i$ axis, shown schematically in Figure (3.1).

We denote by $F_{i j}$ the sum of all inter-granular forces acting on th.e area $\delta A$ in the $j$ direction. The pore pressure $p$ exerts a force $p \delta A$ in opposition to the normal inter-granular forces. Since $\tau_{i j}$ is the net total stress in the $j$ direction we have the following balance of forces on the area $\delta A$ :

$$
\tau_{i j} \delta A=F_{i j}-p \delta A \delta_{i j}
$$

We define the effective stress components $\tau_{i j}^{\prime}$ by

$$
\begin{equation*}
\tau_{i j}^{\prime}=\frac{F_{i j}}{\delta A} \tag{3.6}
\end{equation*}
$$

Hence the above equation becomes

$$
\begin{equation*}
\tau_{i j}=\tau_{i j}^{\prime}-p \delta_{i j} \tag{3.7}
\end{equation*}
$$

Writing (3.5) in terms of effective stress gives

$$
\begin{equation*}
\tau_{i j}^{\prime}=2 \mu e_{i j}+\delta_{i j} \lambda e+\delta_{i j}(1-\alpha) p \tag{3.8}
\end{equation*}
$$

For most soils the ratio $\frac{K_{b}}{K_{\mathrm{r}}}$ is negligible (of the order $10^{-4}$ ) so that $\alpha$ is effectively equal to unity (equation 3.2). Thus (3.8) reduces to

$$
\begin{equation*}
\tau_{i j}^{\prime}=2 \mu e_{i j}+\delta_{i j} \lambda e \tag{3.9}
\end{equation*}
$$

which is Hooke's Law for the soil skeleton.
We now re-write the constitutive laws (3.1) in a form suitable for deriving the boundary integral equations. It is first necessary to write the total stress tensor in terms of stresses on the solid and liquid portions of an element of soil :


Fig.(3.1) Plane area $\delta A$ normal to i-axis

Referring again to Figure (3.1) we define $\sigma_{i j}$ as the net $j$-direction stress on the solid area averaged over the entire area $\delta A$ i.e. $\sigma_{i j} \delta A$ is the $j$-direction resultant of intergranular and pore pressure forces on the solid area. With this definition we find

$$
\sigma_{i j} \delta A=F_{i j}-\delta_{i j}(1-f) \delta A p
$$

which simplifies to (using 3.6)

$$
\begin{equation*}
\sigma_{i j}=\tau_{i j}^{\prime}-\delta_{i j}(1-f) p \tag{3.10}
\end{equation*}
$$

Eliminating $\tau_{i j}^{\prime}$ between (3.7) and (3.10) gives

$$
\begin{equation*}
\tau_{i j}=\sigma_{i j}+\delta_{i j} \sigma \tag{3.11}
\end{equation*}
$$

where

$$
\begin{equation*}
\sigma=-f p \tag{3.12}
\end{equation*}
$$

We proceed to re-write the constitutive equations (3.1) in terms of $\sigma_{i j}$ and $\sigma$ : Writing $\varepsilon=U_{k, k}$ we note that

$$
\zeta=-f(\varepsilon-e)
$$

Multiplying equation (3.1) (b) by $-f$ gives

$$
\begin{equation*}
\sigma=Q e+R \varepsilon \tag{3.13}
\end{equation*}
$$

where

$$
\begin{equation*}
Q=(\alpha-f) f M \quad ; \quad R=f^{2} M \tag{3.14}
\end{equation*}
$$

Substituting equation (3.1) into (3.11) gives, using (3.4)

$$
\begin{align*}
\sigma_{i j} & =2 \mu e_{i j}+\delta_{i j}\left[\left\{\lambda+M(\alpha-f)^{2}\right\} e+f M(\alpha-f) \varepsilon\right] \\
& =2 \mu e_{i j}+\left(\lambda+\frac{Q^{2}}{R}\right) e \delta_{i j}+Q \varepsilon \delta_{i j} \tag{3.15}
\end{align*}
$$

Since the soil grain is practically incompressible we may make the assumption $K_{r} \rightarrow \infty$, and thus $\alpha \rightarrow 1$. Then, substituting for $M$ (from 3.2) in (3.14) we have

$$
\begin{equation*}
R=f K_{f} ; \quad \frac{Q}{R}=\frac{1-f}{f} \tag{3.16}
\end{equation*}
$$

To complete the model we present the equatiuns of motion as (Biot, 1962)

$$
\begin{align*}
\tau_{i j, i}+\rho X_{i} & =\rho \ddot{u}_{i}+\rho_{f} \ddot{w}_{i}  \tag{3.17}\\
-p_{i}+\rho_{f} X_{i} & =\rho_{f} \ddot{u}_{i}+m \tilde{w}_{i}+\frac{\rho_{f} g}{k_{0}} \dot{w}_{i} \tag{3.18}
\end{align*}
$$

where

$$
\begin{aligned}
& w_{i}=f\left(U_{i}-u_{i}\right) \\
& \rho=(1-f) \rho_{s}+f \rho_{f} \\
& \rho_{s}=\text { density of soil grains } \\
& \rho_{f}=\text { density of pore water } \\
& f=\text { porosity } \\
& k_{0}=\text { soil permeability in } m s^{-1} \\
& m=\text { added mass parameter }
\end{aligned}
$$

The "added mass" parameter $m$ is due to the coupling of soil grains and fluid in relative motion. In (3.17) and (3.18) we have included the body force per unit mass $X_{i}$. Differentiation with respect to time is denoted by dots. As demonstrated in the author's M. Eng. thesis, the terms in $\bar{u}_{i}$ and $\bar{w}_{i}$ may be deleted from the equations of motion for problems of water wave loading on sand beds. This is due to the low frequency of the loading and the relative stiffness of sand beds (as compared to clay beds, for example). Then using (3.11) and (3.12) we write the
governing equations (3.17) and (3.18) as

$$
\begin{align*}
\sigma_{i j, j}+\sigma_{, i} & =-\rho X_{i}  \tag{3.19}\\
\sigma_{, i}+f \rho_{f} X_{i} & =f k\left(\dot{U}_{i}-\dot{u}_{i}\right) \tag{3.20}
\end{align*}
$$

where

$$
\begin{equation*}
k=\frac{\rho_{f} g f}{k_{0}} \tag{3.21}
\end{equation*}
$$

For convenienre we record again the constitutive laws from (3.15) and (3.13) :

$$
\begin{align*}
\sigma_{i j} & =2 \mu e_{i j}+\left(\lambda+\frac{Q^{2}}{R}\right) e \delta_{i j}+Q \varepsilon \delta_{i j}  \tag{3.22}\\
\sigma & =Q e+R \varepsilon \tag{3.23}
\end{align*}
$$

where

$$
\begin{align*}
& R=f K_{f}  \tag{3.24}\\
& Q=(1-f) K_{f}
\end{align*}
$$

and

$$
\mu=\text { shear modulus of soil skeleton } ; \lambda=\text { Lamé constant of soil skeleton }
$$

The relationships between $\sigma_{i} \cdot \sigma$ and effective stresses $\tau_{i j}^{\prime}$ and pore pressure $p$ are given by (3.10) and (3.12).

### 3.2 The Biot Model in Conventional Notation

It is nstructive to write the basic equations of the Biot soil model in the notation used in conventional soil mechanics. The Cartesian coordinate axes are such that the $x$ and $y$ axes are horizontal and $z$ axis points vertically downward. Normal
stresses in the $x, y$ and $z$ directions are denoted by $\sigma_{x}, \sigma_{y}$ and $\sigma_{z}$ respectively, and shear stresses are denoted by $\tau_{x y}, \tau_{y z}$ and $\tau_{z x}$. The sign convention is exactly opposite to that used in classical elasticity, i.e.
(1) for normal stresses, compression is positive
(2) a shear stress is considered positive if (a) it points in a positive coordinate direction when acting on a plane whose outward normal points in a negative coordinate direction, or (b) it points in a negative coordinate direction when acting on a plane whose outward normal points in a positive coordinate direction.

Effective stresses are denoted by $\sigma_{x}^{\prime}, \sigma_{y}^{\prime}, \sigma_{z}^{\prime}, \tau_{x y}^{\prime}, \tau_{y z}^{\prime}$, and $\tau_{z x}^{\prime}$. In equation (3.7) we identify subscripts $1,2,3$ with the $x, y, z$ coordinates axes respectively. Then, allowing for the change in sign convention (3.7) becomes

$$
\begin{array}{r}
\sigma_{x}^{\prime}=\sigma_{x}-p, \sigma_{y}^{\prime}=\sigma_{y}-p, \sigma_{z}^{\prime}=o_{z}-p \\
\tau_{x y}^{\prime}=\tau_{x y}, \tau_{y z}^{\prime}=\tau_{y z}, \tau_{z x}^{\prime}=\tau_{z x} \tag{3.25}
\end{array}
$$

We have already shown that the constitutive law (3.1) (a) or (3.22), when written in terms of effective stress, is equivalent to Hooke's law for the soil skeleton, equation (3.9). Writing

$$
G \equiv \mu, \lambda=\frac{2 G \nu}{1-2 \nu}
$$

this becomes

$$
\begin{aligned}
\sigma_{x}^{\prime} & =-2 G e_{x}-\left(\frac{2 G \nu}{1-2 \nu}\right) e \\
\sigma_{y}^{\prime} & =-2 G e_{y}-\left(\frac{2 G \nu}{1-2 \nu}\right) e \\
\sigma_{z}^{\prime} & =-2 G e_{x}-\left(\frac{2 G \nu}{1-2 \nu}\right) e
\end{aligned}
$$

$$
\begin{aligned}
& \tau_{x y}^{\prime}=-2 \mu e_{x y} \\
& \tau_{y z}^{\prime}=-2 \mu e_{y z} \\
& \tau_{z x}^{\prime}=-2 \mu e_{z x}
\end{aligned}
$$

where

$$
\begin{aligned}
e_{x}=\frac{\partial u}{\partial x}, e_{y} & =\frac{\partial}{\partial y}, \iota_{z}=\frac{\partial w}{\partial z} \\
e_{x y} & =\frac{1}{2}\left(\frac{\partial u}{\partial y}+\frac{\partial v}{\partial x}\right) \\
e_{y z} & =\frac{1}{2}\left(\frac{\partial v}{\partial z}+\frac{\partial w}{\partial y}\right) \\
e_{z x} & =\frac{1}{2}\left(\frac{\partial w}{\partial x}+\frac{\partial u}{\partial z}\right)
\end{aligned}
$$

$u, v, w$ are displacements $u_{1}, u_{2}, u_{3}$ of the soil skeleton in the $x, y, z$ directions respectively.

The parameter $G$ is the shear modulus of the soil skeleton, and the Lamé constant $\lambda$ is expressed in terms of the more familiar Poisson's ratio $\nu$. The constitutive law for pore pressure is written from (3.23) and (3.24) as

$$
\begin{equation*}
p=-\frac{K_{f}}{f}[(1-f) e+f \varepsilon] \tag{3.26}
\end{equation*}
$$

where

$$
\begin{array}{r}
e=\frac{\partial u}{\partial x}+\frac{\partial v}{\partial y}+\frac{\partial w}{\partial z} \\
\varepsilon=\frac{\partial U}{\partial x}+\frac{\partial V}{\partial y}+\frac{\partial W}{\partial z}
\end{array}
$$

$(U, V, W)=\left(U_{1}, U_{2}, U_{3}\right)$, the displacements of the pore water.

The governing equations (3.17) and (3.18) become, neglecting the acceleration terms and using (3.25) :

$$
\begin{gather*}
\frac{\partial \sigma_{x}^{\prime}}{\partial x}+\frac{\partial \tau_{x y}^{\prime}}{\partial y}+\frac{\partial \tau_{x z}^{\prime}}{\partial z}-\rho X=-\frac{\partial p}{\partial x} \\
\frac{\partial \tau_{x y}^{\prime}}{\partial y}+\frac{\partial \sigma_{y}^{\prime}}{\partial y}+\frac{\partial \tau_{y z}^{\prime}}{\partial y}-\rho Y=-\frac{\partial p}{\partial y}  \tag{3.27}\\
\frac{\partial \tau_{x z}^{\prime}}{\partial z}+\frac{\partial \tau_{y z}^{\prime}}{\partial z}+\frac{\partial \sigma_{z}^{\prime}}{\partial z}-\rho Z=-\frac{\partial p}{\partial z} \\
-\underline{\nabla} p+\rho_{f} X=\frac{\rho_{f} g f}{k_{0}} \frac{\partial}{\partial t}(\underline{U}-\underline{u}) \tag{3.28}
\end{gather*}
$$

where $\underline{X}=(X, Y, Z)$ is the body force per unit mass; $\underline{U}=(U, V, W) ; \underline{u}=(u, v, w)$. Equation (3.28) is recognised as Darcy's Law. It reduces to a more familiar form by considering no body force $X$, zero displacement of the soil skeleton ( $\underline{u}=\underline{0}$ ) and by defining the fluid velocity $\underline{U}^{*}=f \underline{U}$.

It is possible to define a coefficient of consolidation analogous to that defined in Terzaghi's consolidation theory. It has been shown by the author (RamanNair,1985) that the pore pressure $p$ can be written as the sum of an harmonic function $r$ and a function $q$ which satisfies the "heat equation" form. Specifically,

$$
\begin{equation*}
p(x, z, t)=q(x, z, t)+r(x, z, t) \tag{3.29}
\end{equation*}
$$

where

$$
\begin{align*}
\nabla^{2} r & =0  \tag{3.30}\\
c_{v} \nabla^{2} q & =\frac{\partial q}{\partial t} \tag{3.31}
\end{align*}
$$

with

$$
\begin{equation*}
c_{v}=\frac{k_{0}}{\rho_{f} g}\left(\frac{f}{K_{f}}+\frac{1-2 \nu}{2 \mu(1-\nu)}\right)^{-1} \tag{3.32}
\end{equation*}
$$

Equation (3.31) is of the same form as Terzaghi's consolidation equation and the coefficient $c$, may thus be defined as a consolidation coefficient. To the author's knowledge, equations (3.29) to (3.31) have not previously appeared in the literature. The formula for $c_{v}$ has been presented without derivation by Cheng and Liu (1986), and in different notation using a different analysis by Yamamoto (1978).

It will not be convenient to use the notation of this section in the following development. We shall, however, return to this notation when we consider the failure analysis.

## Chapter 4

## Wave Forces on a Sloping Bed

### 4.1 Problem Formulation and Boundary Condi- <br> tions

According to Sleath (1984) viscous effects are unimportant when considering the wave loading of the seafloor in the absence of strong currents. We shall therefore assume that the flow is essentially irrotational. Further, we may ignore shear stresses at the mudline since these are due primarily to the fluid viscosity. For the free surface, we employ the assumptions of linear wave theory i.e. that the wave amplitude is small relative to the wavelength so that boundary conditions may be applied at the still water level. It is worth noting that for the purpose of determining the wave forces on the seabed, the error involved in using linear wave theory versus higher order wave theories is negligible. For example, in evaluating the dynamic
wave pressure on the seabed under a wave crest, we find that for a wave of length 300 m ., height 24 m . in 80 m . of water, the linear theory estimate is only 1.3 per cent higher than the estimate using Stokes' second order theory.

We shall assume that there is negligible flow into the seabed i.e. the fluid velocity normal to the bed is essentially zero. There have been some attempts in the literature to account for flow into the seabed. For example, using Darcy's law Dean and Dalrymple (1984) write

$$
-\frac{\partial \Phi}{\partial y}=-\frac{K}{\mu} \frac{\partial p}{\partial y} \text { at seabed }
$$

where

$$
\begin{aligned}
\frac{K}{\mu} & =\frac{k_{0}}{\rho_{f} g} \\
\Phi & =\text { wave velocity potential } \\
k_{0} & =\text { soil permeability } \\
\mu & =\text { shear modulus of soil } \\
\rho_{f} & =\text { density of water } \\
y & =\text { vertical coordinate }
\end{aligned}
$$

However, the pore pressure $p$ on the right hand side is not known until the soil equations are solved. Dean and Dalrymple obtain an approximate picture of the pore pressure by employing a simplified model, viz., the assumption that $p$ obeys Laplace's equation in the soil and decays exponentially with soil depth. They deduce that for a wave of amplitude $\eta_{0}$, wave number $k$ and frequency $\omega$ in water
of depth $h$, the wave pressure $P$ at the mudline is given by

$$
P=\frac{\rho_{f} g \eta_{0}}{\cosh k h\left[1-\frac{i \omega k_{0}}{g} \tanh k h\right]} e^{i(k x-\omega t)}
$$

We note that for 10 second waves in 80 m . of water propagating over sandy beds the imaginary term in the denominator is of the order $10^{-4}$ to $10^{-6}$ so that the above equation may be approximated by

$$
\begin{equation*}
P=\frac{\rho_{f} g \eta_{0}}{\cosh k h} e^{i(k x-\omega t)} \tag{4.1}
\end{equation*}
$$

which is the same as would be obtained by assuming negligible flow into the seabed. Further support for this assumption comes from the fact that good agreement with experimentally measured wave-induced pore pressure was obtained by Yamamoto et al. (1978) and Cheng et. al. (1986). They used (4.1) to provide boundary conditions at the mudline for the poroelastic soil model.

It is reasonable to assume that the problem is two dimensional i.e. the slope is planar and conditions are uniform in the direction normal to the direction of propagation of the wave. The problem domain $O A B C$ and incident wave direction are shown in Figure (4.1). The condition of irrotationality guarantees the existence of a velocity potential $\Phi(x, y, t)$, where $x$ and $y$ are spatial coordiuates as shown (Figure (4.1)), and $t$ is time. The wave forces on the seabed will be computed from the function $\Phi$. The flow above the seabed may be considered incompressible and the continuity equation then requires that

$$
\begin{equation*}
\nabla^{2} \Phi=0 \tag{4.2}
\end{equation*}
$$

where

$$
\nabla^{2} \equiv \frac{\partial^{2}}{\partial x^{2}}+\frac{\partial^{2}}{\partial y^{2}}
$$



The boundary conditions are

$$
\begin{align*}
\frac{\partial \Phi}{\partial n} & =0 \quad \text { on } \mathrm{BC}  \tag{4.3}\\
\frac{\partial^{2} \Phi}{\partial t^{2}}+g \frac{\partial \Phi}{\partial y} & =0 \quad \text { on } \mathrm{OA} \tag{4.4}
\end{align*}
$$

where we have made the assumptions of classical linear wave theory (see Appendix A); $\frac{\partial \Phi}{\partial n}$ is the normal derivative on $B C$.

Since the wave frequency $\omega$ coes not change with water depth, the function $\Phi$ may be written

$$
\begin{equation*}
\Phi(x, y, t)=\phi(x, y) e^{-i \omega t} \tag{4.5}
\end{equation*}
$$

where $\phi(x, y)$ is a complex function and it is understood that the physical situation is represented by the real part of equation (4.5). Thus, equation (4.2) with boundary conditions ( 4.3) and (4.4) may be re-written:

$$
\begin{array}{rlrl}
\nabla^{2} \phi & =0 & \text { in } O A B C \\
\frac{\partial \phi}{\partial n} & =0 & & \text { on } B C \\
\frac{\partial \phi}{\partial y}-\frac{\omega^{2} \phi}{g} & =0 & & \text { on } O A \tag{4.8}
\end{array}
$$

The boundary $O C$ (Fig. 4.1) experiences waves progressing from the right (incident waves) as well as reflected waves from the slope coming from the left. The boundary $A B$,however, experiences waves travelling in one direction only i.e. towards the left. To obtain the boundary conditions on the interfaces $O C$ and $A B$ we follow the approach of Alliney (1981). In the problem considered by Alliney the boundary $A B$ (Fig. 4.1) is absent since he considered a coastal situation (i.e. $h_{0}^{\prime}=0$ ). Here we need to determine the boundary condition on $A B$. First, we obtain the general solution $\phi_{a}$ of equation (4.6) in an open-ended two-dimensional domain of constant
depth $h$. The boundary conditions are:

$$
\begin{array}{rlr}
\frac{\partial \phi_{a}}{\partial n} & =0 & \text { at } y=-h \\
\frac{\partial \phi_{a}}{\partial y}-\frac{\omega^{2} \phi_{a}}{g} & =0 & \text { at } y=0 \tag{4.10}
\end{array}
$$

Using the technique of separation of variables we assume that

$$
\phi_{a}(x, y)=X(x) Y(y)
$$

Substitution into (4.6) gives, for sinusoidal behaviour in $X$,

$$
\frac{d^{2} X}{d x^{2}}+k^{2} X=0 \text { and } \frac{d^{2} Y}{d y^{2}}-k^{2} Y=0
$$

where $k$ is a complex constant. The general solutions of these equations are

$$
\begin{gathered}
X=a e^{i k x}+b e^{-i k x} \\
Y=C \cosh (k y+\alpha)
\end{gathered}
$$

where $a, b, C$ and $\alpha$ are complex constants. From boundary condition (4.9) we find that $\alpha=k h$ and hence the solution is of the form

$$
\begin{equation*}
\phi_{a}=\cosh (k y+k h)\left[c e^{i k x}+d e^{-i k x}\right] \tag{4.11}
\end{equation*}
$$

Using (4.10) we have the dispersion relation

$$
\begin{equation*}
\frac{\omega^{2}}{g}=k \tanh (k h) \tag{4.12}
\end{equation*}
$$

This equation has two real roots $\pm k_{0}$ and an infinity of purely imaginary roots $\pm i k_{1}, \pm i k_{2}, \ldots$, where $k_{i} \geq 0, i=0,1,2 \ldots$ (see Appendix B). Since (4.6) is a
linear equation we obtain its general solution from (4.11) as

$$
\begin{aligned}
\phi_{a}= & \cosh \left(k_{0} y+k_{0} h\right)\left[c_{0} e^{i k_{0} x}+d_{0} e^{-i k_{0} x}\right] \\
& +\cosh \left(-k_{0} y-k_{0} h\right)\left[c_{0}^{\prime} e^{-i k_{0} x}+d_{0}^{\prime} e^{i k_{0} x}\right] \\
& +\sum_{n=1}^{\infty} \cosh \left[i k_{n}(y+h)\right]\left[c_{n} e^{-k_{n} x}+d_{n} e^{k_{n} x}\right] \\
& +\sum_{n=1}^{\infty} \cosh \left[-i k_{n}(y+h)\right]\left[c_{n}^{\prime} e^{k_{n} x}+d_{n}^{\prime} e^{-k_{n} x}\right]
\end{aligned}
$$

This may be written more concisely as

$$
\begin{align*}
\phi_{a} & =A \cosh \left[k_{0}(y+h)\right]\left[a_{0} e^{i k_{0} x}+e^{-i k_{0} x}\right] \\
& +\sum_{n=1}^{\infty} \cos \left[k_{n}(y+h)\right]\left[a_{n} e^{-k_{n} x}+b_{n} e^{k_{n} x}\right] \tag{4.13}
\end{align*}
$$

The constants $A, a_{0}, a_{n}, b_{n}$ appearing in (4.13) must be interpreted for the external regions $E$ and $E^{\prime}$ which lie outside the problem domain, as illustrated in Figure (4.1).

In the region $E$ to the right of $O C$ in Figure (4.1) we consider an incident wave of amplitude $\eta_{0}$, wave number $k_{0}$ and frequency $\omega$ propagating in the negative $x$ direction. It is easily shown (see Appendix A) that the velocity potential corresponding to the incident wave is:

$$
\begin{equation*}
\phi_{I}(x, y)=-\frac{i g \eta_{0}}{\omega} \frac{\cosh \left[k_{0}\left(y+h_{0}\right)\right]}{\cosh \left(k_{0} h_{0}\right)} e^{-i k_{0} x} \tag{4.14}
\end{equation*}
$$

By putting $h=h_{0}$ in (4.13) we see that the term $A \cosh \left[k_{0}\left(y+h_{0}\right)\right] e^{-i k_{0} x}$ represents the incident wave, so that for the region $E$,

$$
A=\frac{-i g \eta_{0}}{\omega \cosh \left(k_{0} h_{0}\right)}
$$

The term $a_{0} A \cosh \left[k_{0}\left(y+h_{0}\right)\right] e^{i k_{0} x}$ represents the right running reflected wave, and $a_{0}$ can be interpreted as the reflection coefficient. In order that the velocity potential
remain finite as $x \rightarrow \infty$ we must have $b_{n}=0$ for all $n$. Hence from (4.13) we have the velocity potential $\phi_{E}$ in region $E$ :

$$
\begin{align*}
\phi_{E}(x, y) & =-\frac{i g \eta_{0}}{\omega} \frac{\cosh \left[k_{0}\left(y+h_{0}\right)\right]}{\cosh \left(k_{0} h_{0}\right)}\left[a_{0} e^{i k_{0} x}+e^{-i k_{0} x}\right] \\
& +\sum_{n=1}^{\infty} a_{n} \cos \left[k_{n}\left(y+h_{0}\right)\right] e^{-k_{n} x} \tag{4.15}
\end{align*}
$$

For the region $E^{\prime}$ to the left of $A B$ we put $h=h_{0}^{\prime}$ in equation (4.12) and denote the roots by $\pm k_{0}^{\prime}, \pm i k_{n}^{\prime}, n=1,2 \cdots$ where $k_{0}^{\prime}, k_{n}^{\prime} \geq 0$. Since there is no right running wave in this region, the coefficient $a_{0}$ in (4.13) must be zero. Also, in order that the velocity potential be finite as $x \rightarrow-\infty$ we must have $a_{n}=0$ for all $n$. Thus, from (4.13), we have the velocity potential $\phi_{E^{\prime}}$ in region $E^{\prime}$ as

$$
\begin{equation*}
\phi_{E^{\prime}}(x, y)=a_{0}^{\prime} \cosh \left[k_{0}^{\prime}\left(y+h_{0}^{\prime}\right)\right] e^{-i k_{0}^{\prime} x}+\sum_{n=1}^{\infty} a_{n}^{\prime} e^{\prime_{n}^{\prime} x} \cos \left[k_{n}^{\prime}\left(y+h_{0}^{\prime}\right)\right] \tag{4.16}
\end{equation*}
$$

We can now obtain expressions for $\phi$ and its normal derivative on the lateral boundaries $O C$ and $A B$. For $O C$, we put $x=0$ in (4.15) and its derivative with respect to $x$ :

$$
\begin{align*}
\phi_{l o c} & =f_{0}(y)+\sum_{n=1}^{N+1} a_{n} f_{n}(y), y \in O C  \tag{4.17}\\
\left.\frac{\partial \phi}{\partial n}\right|_{O C} & =-g_{0}(y)+\sum_{n=1}^{N+1} a_{n} g_{n}(y), y \in O C \tag{4.18}
\end{align*}
$$

where we have used $N$ terms of the infinite series and :

$$
\begin{align*}
f_{0}(y) & =\frac{-i g \eta_{0}}{\omega} \frac{\cosh \left[k_{0}\left(y+h_{0}\right)\right]}{\cosh \left(k_{0} h_{0}\right)} \\
f_{n}(y) & =\cos \left[k_{n}\left(y+h_{0}\right)\right], n=1,2, \cdots, N \\
g_{0}(y) & =\frac{g \eta_{0} k_{0}}{\omega} \frac{\left[k_{0}\left(y+h_{0}\right)\right]}{\cosh \left(k_{0} h_{0}\right)}  \tag{4.19}\\
g_{n}(y) & =-k_{n} \cos \left[k_{n}\left(y+h_{0}\right)\right], n=1,2, \cdots, N \\
f_{N+1}(y) & =f_{0}(y), g_{N+1}(y)=g_{0}(y), a_{N+1}=a_{0} \tag{4.20}
\end{align*}
$$

For $A B$ we have

$$
\begin{aligned}
\left.\phi\right|_{A B} & =\phi_{E^{\prime}}(-d, y) \\
& =a_{0}^{\prime} e^{i k_{0}^{\prime} d} \cosh \left[k_{0}^{\prime}\left(y+h_{0}^{\prime}\right)\right]+\sum_{n=1}^{N} a_{n}^{\prime} e^{-k_{n}^{\prime} d} \cos \left[k_{n}^{\prime}\left(y+h_{0}^{\prime}\right)\right] \\
\left.\frac{\partial \phi}{\partial n}\right|_{A B} & =-\left.\frac{\partial \phi}{\partial x}\right|_{A B} \\
& =i k_{0}^{\prime} a_{0}^{\prime} e^{i k_{0}^{\prime} d} \cosh \left[k_{0}^{\prime}\left(y+h_{0}^{\prime}\right)\right]+\sum_{n=1}^{N^{\prime}}-a_{n}^{\prime} k_{n}^{\prime} e^{-k_{n}^{\prime} d} \cos \left[k_{n}^{\prime}\left(y+h_{0}^{\prime}\right)\right]
\end{aligned}
$$

Since the horizontal distance $d$ is large, the terms in $e^{-k_{n}^{\prime} d}$ are effectively zero and these equations reduce to

$$
\begin{aligned}
\left.\phi\right|_{A B} & =a_{0}^{\prime} e^{i k_{0}^{\prime} d} \cosh \left[k_{0}^{\prime}\left(y+h_{0}^{\prime}\right)\right] \\
\left.\frac{\partial \phi}{\partial n}\right|_{A B} & =i k_{0}^{\prime} a_{0}^{\prime} e^{i k_{0}^{\prime} d} \cosh \left[k_{0}^{\prime}\left(y+h_{0}^{\prime}\right)\right]
\end{aligned}
$$

Thus we have the relation

$$
\begin{equation*}
\left.\frac{\partial \phi}{\partial n}\right|_{A B}=\left.i k_{0}^{\prime} \phi\right|_{A B} \tag{4.21}
\end{equation*}
$$

which is a well known radiation condition.

### 4.2 The Bouidary Integral Equation

The problem defined by (4.6) with boundary conditions (4.7), (4.8), (4.17), (4.18) and (4.21), may be solved by the boundary element method with a view to determining $\phi$, and hence the wave pressure, on the seabed $B C$. The following derivation of the boundary integral equation is standard (Brebbia et al. 1984) but is presented here for completeness. For an arbitrary function $\phi^{*}(x, y)$ we may write,
from (4.6):

$$
\int_{\Omega} \phi^{*} \nabla^{2} \phi d \Omega=0
$$

where $\Omega$ denotes the region $O A B C$.
Assuming that $\phi, \phi^{*}$ are continuous and ha:e centinuous first and second partial derivatives in $\Omega$ we use Green's second identity (Sokolnikoff et al., 1966) to get

$$
\begin{equation*}
\int_{\Omega} \phi \nabla^{2} \phi^{*} d \Omega+\int_{j \Gamma}\left(\phi^{*} \frac{\partial \phi}{\partial n}-\phi \frac{\partial \phi^{*}}{\partial n}\right) d \Gamma=0 \tag{4.22}
\end{equation*}
$$

where $\Gamma$ denotes the boundary $O A B C$. We choose $\phi^{*}$ to be the solution of the equation

$$
\begin{equation*}
\nabla^{2} \phi^{*}+\Delta\left(\underline{x}, \underline{x}_{p}\right)=0 \tag{4.23}
\end{equation*}
$$

where $\Delta$ denotes the Dirac delta function, $x_{p}$ is the position vector of the point $P$ at which the delta function is non-zero, and $\underline{x}$ is the position vector of an arbitrary point. The solution of equation (4.23) for the case of two dimensions is

$$
\begin{equation*}
\phi^{*}=-\frac{1}{2 \pi} \ln r \tag{4.24}
\end{equation*}
$$

where $r=\left|\underline{x}-\underline{x}_{p}\right|$ (see Appendix D). Using (4.23) we have

$$
\begin{equation*}
\int_{\Omega} \phi \nabla^{2} \phi^{*} d \Omega=-\int_{\Omega} \phi \Delta\left(\underline{x}, x_{p}\right) d \Omega=-\phi\left(\underline{x}_{p}\right) \tag{4.25}
\end{equation*}
$$

by the sifting property of the delta function. Substituting (4.25) into (4.22) gives

$$
\begin{equation*}
-\phi\left(x_{p}\right)+\int_{\Gamma}\left(\phi^{*} \frac{\partial \phi}{\partial n}-\phi \frac{\partial \phi^{*}}{\partial n}\right) d \Gamma=0 \tag{4.26}
\end{equation*}
$$

We now take the point $P$ (position vector $x_{p}$ ) to the boundary $\Gamma$. Since $\phi^{*}$ is singular at $P$, integration along the section of $\Gamma$ containing $P$ requires special consideration. For generality, we consider $P$ at a corner point on the boundary and draw a circular


Fig.(4.2) Boundary Point P.
$\operatorname{arc} S_{\epsilon}$ of radius $\epsilon$, centre $P$ as shown in Figure (4.2). We now write equation (4.26) for a boundary point $x_{p}$ :

$$
\begin{align*}
-\phi\left(\underline{x}_{p}\right) & +\lim _{c \rightarrow 0} \int_{\Gamma-S_{e}}\left(\phi^{*} \frac{\partial \phi}{\partial n}-\phi \frac{\partial \phi^{*}}{\partial n}\right) d \Gamma \\
& +\lim _{\varepsilon \rightarrow 0} \int_{S_{\varepsilon}}\left(\phi^{*} \frac{\partial \phi}{\partial n}-\phi \frac{\partial \phi^{*}}{\partial n}\right) d \Gamma=0 \tag{4.27}
\end{align*}
$$

Now,

$$
\begin{align*}
\lim _{\epsilon \rightarrow 0} \int_{S_{e}} \phi^{*} \frac{\partial \phi}{\partial n} d \Gamma & =\lim _{\epsilon \rightarrow 0} \int_{S_{e}}-\frac{1}{2 \pi} \ln r \frac{\partial \phi}{\partial n} d \Gamma \\
& =-\left.\frac{1}{2 \pi} \frac{\partial \phi}{\partial n}\right|_{S_{\varepsilon}} \lim _{\epsilon \rightarrow 0}\left[\ln \epsilon \int_{S_{e}} d \Gamma\right] \\
& =-\left.\frac{1}{2 \pi} \frac{\partial \phi}{\partial n}\right|_{s_{e}} \lim _{\epsilon \rightarrow 0}[(\pi+\alpha) \epsilon \ln \epsilon]=0 \tag{4.28}
\end{align*}
$$

where we have assumed that $\frac{\partial \phi}{\partial n}$ is approximately constant on the small arc $S_{c}$. Also,

$$
\begin{align*}
\lim _{\epsilon \rightarrow 0} \int_{S_{e}} \phi \frac{\partial \phi^{*}}{\partial n} d \Gamma & =\lim _{\epsilon \rightarrow 0} \int_{S_{e}} \phi \frac{\partial}{\partial r}\left(-\frac{1}{2 \pi} \ln r\right) d \Gamma \\
& =-\frac{\phi\left(\underline{x}_{p}\right)}{2 \pi} \lim _{e \rightarrow 0} \int_{S_{e}} \frac{1}{r} d \Gamma \\
& =\frac{-\phi\left(\underline{x}_{p}\right)}{2 \pi} \lim _{e \rightarrow 0}\left[\frac{1}{\epsilon} \int_{S_{e}} d \Gamma\right] \\
& =\frac{-\phi\left(\underline{x}_{p}\right)}{2 \pi}(\pi+\alpha) \tag{4.29}
\end{align*}
$$

where $\alpha$ is the angle illustrated in Figure (4.2). Substituting (4.28) and (4.29) into (4.27) we have for a boundary point $P$ :

$$
\begin{equation*}
\phi\left(\underline{x}_{p}\right)\left(\frac{1}{2}-\frac{\alpha}{2 \pi}\right)+\int_{\Gamma}\left(\phi \frac{\partial \phi^{*}}{\partial n}-\phi^{*} \frac{\partial \phi}{\partial n}\right) d \Gamma=0 \tag{4.30}
\end{equation*}
$$

where it is understood that the domain of integration does not include the point $x_{p}$.

### 4.3 Boundary Element Procedure

The boundary $\Gamma$ is discretised into $M$ straight line segments or "elements" and a node is located at the mid-point of each element. We shall assume that $\phi$ and $\frac{\partial \phi}{\partial n}$ are constant and take their nodal values on each element. The use of these "constant" elements gives very good accuracy, when the boundary does not exhibit a significant degree of curvature. The "constant" element is the simplest type of discontinous element and avoids the ambiguities present in the value of $\frac{\partial \phi}{\partial n}$ at corner points.

The boundary element technique involves writing the boundary integral equation (4.30) for each of the $M$ nodes, i.e. putting $\underline{x}_{p}$ equal to each node in turn. This produces $M$ equations in $2 M$ variables, these being the nodal values of $\phi$ and $\frac{\partial \phi}{\partial n}$. In a well posed problem we know either $\phi, \frac{\partial \phi}{\partial n}$ or a relationship between them at each node, which thus reduces the number of unknowns to $M$.

We let $\phi_{c}$ and $\phi_{c}^{\prime}$ denote the values of $\phi$ and $\frac{\partial \phi}{\partial n}$ respectively at node $e$, which lies at the mid-point of element $e$ as noted above. In equation (4.30) we let the point $x_{p}$ lie at the $j$ th node, so that $\phi\left(x_{p}\right)=\phi_{j}$ and $r=\left|\underline{x}-x_{j}\right|$, where $\underline{x}_{j}$ is the position vector of the $j$ th node and $\underline{x}$ is the position vector of an arbitrary point on the boundary.

Since $x_{p}$ is not a corner point, $\alpha=0$ and the discretisation of (4.30) leads to

$$
\begin{equation*}
\frac{1}{2} \phi_{j}+\sum_{e=1}^{M} \int_{\Gamma_{e}}\left(\phi_{e} \frac{\partial \phi^{*}}{\partial n}-\phi_{e}^{\prime} \phi^{*}\right) d \Gamma_{e}=0 \tag{4.31}
\end{equation*}
$$

where $\Gamma_{e}$ denotes element $e$ and we recall that $\phi^{*}$ is given by (4.24). Since $\phi_{e}$ and $\phi_{e}^{\prime}$ are constant on $\Gamma_{e}$ we can take them outside the integral sign and (4.31) takes
the form :

$$
\begin{equation*}
\sum_{e=1}^{M}\left(A_{e} \phi_{e}-B_{e} \phi_{e}^{\prime}\right)=0 \tag{4.32}
\end{equation*}
$$

where

$$
\begin{align*}
& A_{e}= \begin{cases}\int_{\Gamma_{e}} \frac{\partial \phi^{*}}{\partial n} d \Gamma_{e} & e \neq j \\
\int_{\Gamma_{j}} \frac{\partial \phi^{*}}{\partial n} d \Gamma_{j}+\frac{1}{2} & e=j\end{cases}  \tag{4.33}\\
& B_{e}=\int_{\Gamma_{e}} \phi^{*} d \Gamma_{c}
\end{align*}
$$

The evaluation of the integrals $A_{\mathrm{e}}$ and $B_{\mathrm{e}}$ is described in Appendix C.
We choose $M_{1}, M_{2}, M_{3}, M_{4}$ elements on $O A, A B, B C$ and $C O$ respectively (Figure (4.1)), the elements (and nodes) being numbered in an anti-clockwise direction starting from 0 , i.e. 0 is the first point of element 1 and the second point of element $M$, where $M=M_{1}+M_{2}+M_{3}+M_{4}$. From boundary conditions (4.7), (4.8) and (4.21) we find that

$$
\begin{array}{lll}
\text { on } O A & \phi_{e}^{\prime}=\frac{\omega^{2}}{g} \phi_{e} & e=1,2, \cdots, M_{1} \\
\text { on } A B & \phi_{e}^{\prime}=i k_{0}^{\prime} \phi_{e} & e=M_{1}+1, \cdots, M_{1}+M_{2} \\
\text { on } B C & \phi_{e}^{\prime}=0 & e=M_{1}+M_{2}+1, \cdots, M_{1}+M_{2}+M_{3} \tag{4.36}
\end{array}
$$

On CO we deduce the relationship between $\phi$ and $\frac{\partial \phi}{\partial n}$ as follows. We must first choose $M_{4}=N+1$ which is the number of coefficients $a_{n}$ appearing in equations (4.17) and (4.18). The $y$ coordinates of the $M_{4}$ nodes on $C O$ will be dencted by $y_{1}, y_{2}, \cdots y_{M_{4}}$ corresponding to nodes $M_{1}+M_{2}+M_{3}+1, \cdots, M$. We write equations
(4.17) and (4.18) at each of these nodes :

$$
\begin{aligned}
\phi\left(y_{k}\right)-f_{0}\left(y_{k}\right) & =\sum_{n=1}^{M_{4}} a_{n} f_{n}\left(y_{k}\right) \\
\phi^{\prime}\left(y_{k}\right)+g_{0}\left(y_{k}\right) & =\sum_{n=1}^{M_{4}} a_{n} g_{n}\left(y_{k}\right) \\
k & =1,2, \cdots, M_{4}
\end{aligned}
$$

These may be written in matrix form :

$$
\begin{align*}
\{\phi\}_{C O}-\{f\} & =[F]\{a\}  \tag{4.37}\\
\left\{\phi^{\prime}\right\}_{C O}+\{g\} & =[G]\{a\} \tag{4.38}
\end{align*}
$$

where

$$
\begin{array}{rlrl}
\{\phi\}_{c o} & =\left\{\phi\left(y_{k}\right)\right\}, & k=1,2, \cdots, M_{4} \\
& =\left\{\phi_{e}\right\}, & e=M_{1}+M_{2}+M_{3}+1, \cdots, M \\
\left\{\phi^{\prime}\right\}_{c o} & =\left\{\phi^{\prime}\left(y_{k}\right)\right\}, & & k=1,2, \cdots, M_{4} \\
& =\left\{\phi_{c}^{\prime}\right\}, & e=M_{1}+M_{2}+M_{3}+1, \cdots, M \\
\{f\} & =\left\{f_{0}\left(y_{k}\right)\right\} & k=1,2, \cdots, M_{4} \\
\{g\} & =\left\{g_{0}\left(y_{k}\right)\right\} & k=1,2, \cdots, M_{4} \\
\{a\} & =\left\{a_{n}\right\} & n=1,2, \cdots, M_{4} \\
f_{M_{4}}=f_{0}, & g_{M_{4}}=g_{0}, & a_{M_{4}}=a_{0}
\end{array}
$$

Also, $F_{k n}=f_{n}\left(y_{k}\right), G_{k n}=g_{n}\left(y_{k}\right)$ are the $k-n$ elements of the $M_{4} \times M_{4}$ matrices $[F]$ and $[G]$ respectively.

From (4.37)

$$
\{a\}=[F]^{-1}\left(\{\phi\}_{C O}-\{f\}\right)
$$

Substituting in (4.38) gives

$$
\begin{equation*}
\left\{\phi^{\prime}\right\}_{C O}=[H]\left(\{\phi\}_{C O}-\{f\}\right)-\{g\} \tag{4.39}
\end{equation*}
$$

where

$$
[H]=[G][F]^{-1}
$$

We now note that equation (4.32) may be written in vector form as

$$
\begin{equation*}
\left\{A_{e}\right\}_{\Gamma}^{T}\left\{\phi_{e}\right\}_{\Gamma}-\left\{B_{e}\right\}_{\Gamma}^{T}\left\{\phi_{e}^{\prime}\right\}_{\Gamma}=0 \tag{4.40}
\end{equation*}
$$

where $\left\{A_{e}\right\}_{\Gamma},\left\{B_{e}\right\}_{\Gamma},\left\{\phi_{e}\right\}_{\Gamma},\left\{\phi_{e}^{\prime}\right\}_{\Gamma}$ denote the $M$-dimensional vectors of the quantities $A_{e}, B_{e}, \phi_{e}, \phi_{e}^{\prime}$ on the closed curve $\Gamma$; the superscript " T " denotes the transpose of a vector. In order to apply the boundary conditions we write equation ( 4.40 ) explicitly in terms of quantities defined on each section of the boundary $\Gamma$, i.e.

$$
\begin{align*}
& \left\{A_{e}\right\}_{O A}^{T}\left\{\phi_{e}\right\}_{O A}-\left\{B_{e}\right\}_{O A}^{T}\left\{\phi_{e}^{\prime}\right\}_{O A} \\
+ & \left\{A_{e}\right\}_{A B}^{T}\left\{\phi_{e}\right\}_{A B}-\left\{B_{e}\right\}_{A B}^{T}\left\{\phi_{e}^{\prime}\right\}_{A B} \\
+ & \left\{A_{e}\right\}_{B C}^{T}\left\{\phi_{e}\right\}_{B C}-\left\{B_{e}\right\}_{B C}^{T}\left\{\phi_{e}^{\prime}\right\}_{B C} \\
+ & \left\{A_{e}\right\}_{C O}^{T}\left\{\phi_{e}\right\}_{C O}-\left\{B_{e}\right\}_{C O}^{T}\left\{\phi_{e}^{\prime}\right\}_{C O}=0 \tag{4.41}
\end{align*}
$$

From (4.34) - (4.36) we have

$$
\begin{align*}
\left\{\phi_{e}^{\prime}\right\}_{O A} & =\frac{\omega^{2}}{g}\left\{\phi_{e}\right\}_{O A} \\
\left\{\phi_{e}^{\prime}\right\}_{A B} & =i k_{0}^{\prime}\left\{\phi_{e}\right\}_{A B}  \tag{4.42}\\
\left\{\phi_{e}^{\prime}\right\}_{B C} & =0
\end{align*}
$$

Substituting equations (4.42) together with (4.39) into equation (4.41) gives

$$
\left\{A_{e}-\frac{\omega^{2}}{g} B_{e}\right\}_{O A}^{T}\left\{\phi_{e}\right\}_{O A}+\left\{A_{e}-i k_{0}^{\prime} B_{e}\right\}_{A B}^{T}\left\{\phi_{e}\right\}_{A B}
$$

$$
\begin{array}{r}
+\left\{A_{e}\right\}_{B C}^{T}\left\{\phi_{e}\right\}_{B C}+\left(\left\{A_{e}\right\}_{C O}^{T}-\left\{B_{e}\right\}_{C O}^{T}[H]\right)\left\{\phi_{e}\right\}_{C O} \\
=-\left\{B_{e}\right\}_{C O}^{T}([H]\{f\}+\{g\})
\end{array}
$$

This equation may be written :

$$
\left.\begin{array}{l}
\left\{\left\{A_{e}-\frac{\omega^{2}}{g} B_{e}\right\}_{O A}^{T} \quad\left\{A_{e}-i k_{0}^{\prime} B_{e}\right\}_{A B} \quad\left\{A_{c}\right\}_{B C}^{T}\right.
\end{array} \quad\left\{A_{e}\right\}_{C O}^{T}-\left\{B_{e}\right\}_{C O}^{T}[H]\right]\left(\begin{array}{l}
\left\{\phi_{e}\right\}_{O A} \\
\left\{\phi_{e}\right\}_{A B}  \tag{4.43}\\
\left\{\phi_{e}\right\}_{B C} \\
\left\{\phi_{e}\right\}_{C O}
\end{array}\right)
$$

Equation (4.43) is the $j$ th row of the complete matrix equation

$$
\begin{equation*}
[C]\{\phi\}=\{b\} \tag{4.44}
\end{equation*}
$$

The $(j, k)$-th element of matrix $[C]$ is

$$
C_{j k}= \begin{cases}A_{k}-\frac{\omega^{2}}{g} B_{k} & k=1, \cdots, M_{1}  \tag{4.45}\\ A_{k}-i k_{0}^{\prime} B_{k} & k=M_{1}+1, \cdots, M_{1}+M_{2} \\ A_{k} & k=M_{1}+M_{2}+1, \cdots, M_{1}+M_{2} \\ A_{k}-\sum_{r=1}^{M_{4}} B_{M_{1}+M_{2}+M_{3}+r} H_{r, k-M_{1}-M_{2}-M_{3}} & \\ & k=M_{1}+M_{2}+M_{3}+1, \cdots, M\end{cases}
$$

The $j$ th element of vector $\{b\}$ is

$$
b_{j}=-\sum_{r=1}^{M_{4}} B_{M_{1}+M_{2}+M_{3}+r} h_{r}
$$

Table 4.1: Comparison of BEM with Analytic Formula for $\phi$ on a Flat Seabed

| $x$ <br> coord. of point <br> on Flat Seabed | $\bar{\phi}(B E M) \times 10^{2} \mathrm{~m}^{2} \mathrm{~s}^{-1}$ |  | $\phi($ Analytic $) \times 10^{2} \mathrm{~m}^{2} s^{-1}$ |  |
| ---: | :---: | :---: | :---: | :---: |
|  | Real | Imag | Real | Imag |
| -63.5 | 0.94559 | -0.23239 | 0.94492 | -0.23209 |
| -50.65 | 0.84967 | -0.47520 | 0.84919 | -0.47499 |
| -20.95 | 0.41352 | -0.88093 | 0.41336 | -0.88084 |
| -5.55 | 0.11285 | -0.96644 | 0.11285 | -0.96644 |

where $h_{r}$ is the $r$ th component of the vector $[H]\{f\}+\{g\}$ i.e.

$$
h_{r}=\sum_{i=1}^{M_{4}} H_{r i} f_{0}\left(y_{i}\right)+g_{0}\left(y_{r}\right)
$$

The solution vector is

$$
\{\phi\}=\left\{\phi_{e}\right\}, e=1,2, \cdots, M
$$

The wave pressure $P$ is determined from Bernoulli's equation as

$$
\begin{equation*}
P=-\rho_{f} \frac{\partial \phi}{\partial t}=i \omega \rho_{f} \phi e^{-i \omega t} \tag{4.46}
\end{equation*}
$$

The accuracy of the computer program was checked by running it for the case of a flat seabed under a wave of length 300 m . and height 24 m . in 80 m . of water. The boundary values of $\phi$ were compared with those computed from the analytic formula ( 4.14). Excellent agreement was obtained as illustrated in Table (4.1) .

## Chapter 5

# Boundary Integral Formulation <br> for Two Dimensional 

## Poroelasticity - Sinusoidal

## Loading Conditions

As noted previously, the boundary element method has been applied to Biot's equations by a number of authors. The boundary element formulation for Laplace transformed space was accomplished by Cheng and Liggett(1984). Although the use of Laplace transforms is convenient for treating the time derivatives, such a formulation is not practical for the present problem owing to the difficulty of numerically inverting the transforms for problems involving sinusoidal loading. Cheng
and Liu (1986) have presented the boundary integral equations and fundamental solutions for sinusoidal loading conditions but without derivation. We present here these derivations and in addition, the functions required for determining interior effective stresses and pore pressure.

### 5.1 The Boundary Integral Equations

The governing equations (3.19) and (3.20) are recorded again for convenience :

$$
\begin{align*}
\sigma_{i j, j}+\sigma_{, i} & =-\rho X_{i}  \tag{5.1}\\
\sigma_{, i}+f \rho_{f} X_{i} & =f k\left(\dot{U}_{i}-\dot{u}_{i}\right) \tag{5.2}
\end{align*}
$$

The problem is to solve (5.1) and (5.2) in a domain $\Omega$ bounded by a closed curve $\Gamma$, in two dimensions. Either the displacements $u_{i}, U_{i}$ or the tractions $T_{i}$ is known on $\Gamma$. The tractions are given by (Sokc'nikoff, 1956)

$$
\begin{equation*}
T_{i}=\tau_{i j} n_{j} \tag{5.3}
\end{equation*}
$$

and using (3.11) this becomes

$$
\begin{equation*}
T_{i}=\sigma_{i j} n_{j}+\sigma n_{i} \tag{5.4}
\end{equation*}
$$

where $n_{i}$ is the $i$ th component of the unit outward normal on $\Gamma$. A reciprocal theorem may be deduced as follows. We note that the constitutive laws (3.22) and (3.23) are of the form

$$
\begin{aligned}
\sigma_{i j} & =A e_{i j}+(B e+Q \varepsilon) \delta_{i j} \\
\sigma & =Q e+R \varepsilon
\end{aligned}
$$

If we consider two independent states of stress, denoted by primed and unprimed symbols, we have

$$
\begin{aligned}
\sigma_{i j} e_{i j}^{\prime}+\sigma \varepsilon^{\prime} & =A e_{i j} e_{i j}^{\prime}+(B e+Q \varepsilon) e_{i i}^{\prime}+(Q e+R \varepsilon) \varepsilon^{\prime} \\
& =A e_{i j} e_{i j}^{\prime}+B e e^{\prime}+\left(Q \varepsilon e^{\prime}+Q e \varepsilon^{\prime}\right)+R \varepsilon \varepsilon^{\prime}
\end{aligned}
$$

since $e_{i i}^{\prime}=e^{\prime}$
Since the right hand side is symmetric with respect to primed and un-primed symbols, the left hand side must also be symmetric i.e.

$$
\begin{equation*}
\sigma_{i j} e_{i j}^{\prime}+\sigma \varepsilon^{\prime}=\sigma_{i j}^{\prime} e_{i j}+\sigma^{\prime} \varepsilon \tag{5.5}
\end{equation*}
$$

which is the reciprocal theorem for poroelasticity (Cleary,1977). We shall seek the steady state solution of the governing equations under sinusoidal loading of frequency $\omega$, i.e. we shall assume that a quantity $A$, be it stress or displacement, takes the form $A=\bar{A} e^{-i \omega t}$ where $t$ is time. Equations (5.1),(5.2),(5.4) and (5.5) then take the form

$$
\begin{align*}
\bar{\sigma}_{i j, j}+\bar{\sigma}_{i i} & =-\rho \bar{X}_{i} \text { in } \Omega  \tag{5.6}\\
\bar{\sigma}_{, i}+f \rho_{j} \bar{X}_{i} & =-i \omega f k\left(\bar{U}_{i}-\bar{u}_{i}\right) \text { in } \Omega  \tag{5.7}\\
\bar{T}_{i} & =\bar{\sigma}_{i j} n_{j}+\bar{\sigma} n_{i} \text { on } \Gamma  \tag{5.8}\\
\bar{\sigma}_{i j} \bar{e}_{i j}^{\prime}+\bar{\sigma} \bar{\varepsilon}^{\prime} & =\bar{\sigma}_{i j}^{\prime} \bar{e}_{i j}+\bar{\sigma}^{\prime} \bar{\varepsilon} \tag{5.9}
\end{align*}
$$

Assuming continuity of the first partial derivatives, we apply the divergence theorem to the product $\bar{\sigma}_{i j} \bar{u}_{i}^{\prime}$ :

$$
\int_{\Omega}\left(\bar{\sigma}_{i j} \bar{u}_{i}^{\prime}\right)_{j} d \Omega=\int_{\Gamma} \bar{\sigma}_{i j} \bar{u}_{i}^{\prime} n_{j} d \Gamma
$$

i.e.

$$
\int_{\Omega}\left(\bar{\sigma}_{i j} \bar{u}_{i, j}^{\prime}+\bar{\sigma}_{i j, j} \bar{u}_{i}^{\prime}\right) d \Omega=\int_{\Gamma} \bar{\sigma}_{i j} \bar{u}_{i}^{\prime} n_{j} d \Gamma
$$

Since $i$ and $j$ are dummy indices we can write

$$
\begin{aligned}
\bar{\sigma}_{i j} \bar{u}_{i, j}^{\prime} & =\frac{1}{2}\left(\bar{\sigma}_{i j} \bar{u}_{i, j}^{\prime}+\bar{\sigma}_{j i} \bar{u}_{j, i}^{\prime}\right) \\
& =\frac{1}{2} \bar{\sigma}_{i j}\left(\bar{u}_{i, j}^{\prime}+\bar{u}_{j, i}^{\prime}\right)=\bar{\sigma}_{i j} \bar{e}_{i j}^{\prime}
\end{aligned}
$$

since $\bar{\sigma}_{i j}=\bar{\sigma}_{j i}$
Hence ( 5.10 ) becomes

$$
\begin{equation*}
\int_{\Omega}\left(\tilde{\sigma}_{i j} \bar{e}_{i j}^{\prime}+\bar{\sigma}_{i j, j} \bar{u}_{i}^{\prime}\right) d \Omega=\int_{\Gamma} \bar{\sigma}_{i j} \bar{u}_{i}^{\prime} n_{j} d \Gamma \tag{5.11}
\end{equation*}
$$

Similarly, we may apply the divergence theorem to $\bar{\sigma}_{i j}^{\prime} \bar{u}_{i}$ to get :

$$
\begin{equation*}
\int_{\Omega}\left(\bar{\sigma}_{i j}^{\prime} \bar{e}_{i j}+\bar{\sigma}_{i j, j}^{\prime} \bar{u}_{i}\right) d \Omega=\int_{\Gamma} \bar{\sigma}_{i j}^{\prime} \bar{u}_{i} n_{j} d \Gamma \tag{5.12}
\end{equation*}
$$

Applying the divergence theorem to the product $\bar{\sigma} \bar{U}_{i}^{\prime}$ gives

$$
\int_{\Omega}\left(\bar{\sigma} \bar{U}_{i}^{\prime}\right)_{, i} d \Omega=\int_{\Gamma} \bar{\sigma} \bar{U}_{i}^{\prime} n_{i} d \Gamma
$$

i.e.

$$
\begin{equation*}
\int_{\Omega}\left(\bar{\sigma} \bar{\varepsilon}^{\prime}+\bar{\sigma}_{, i} \bar{U}_{i}^{\prime}\right) d \Omega=\int_{\Gamma} \bar{\sigma} \bar{U}_{i}^{\prime} n_{i} d \Gamma \tag{5.13}
\end{equation*}
$$

Similarly, by applying the divergence theorem to $\bar{\sigma}^{\prime} \bar{U}_{i}$ we have

$$
\begin{equation*}
\int_{\Omega}\left(\bar{\sigma}^{\prime} \bar{\varepsilon}+\bar{\sigma}_{, i}^{\prime} \bar{U}_{i}\right) d \Omega=\int_{\Gamma} \bar{\sigma}^{\prime} \bar{U}_{i} n_{i} d \Gamma \tag{5.14}
\end{equation*}
$$

We now add equations (5.11) and (5.13) to get

$$
\begin{array}{r}
\int_{\Omega}\left(\bar{\sigma}_{i j} \bar{e}_{i j}^{\prime}+\bar{\sigma} \bar{\varepsilon}^{\prime}+\bar{\sigma}_{i j, j} \bar{u}_{i}^{\prime}+\bar{\sigma}_{i} \bar{U}_{i}^{\prime}\right) d \Omega \\
=\int_{\Gamma}\left[\bar{T}_{i} \bar{u}_{i}^{\prime}+\bar{\sigma}\left(\bar{U}_{j}^{\prime}-\bar{u}_{j}^{\prime}\right) n_{j}\right] d \Gamma \tag{5.15}
\end{array}
$$

Adding (5.12) and (5.14) gives

$$
\begin{array}{r}
\int_{\Omega}\left(\bar{\sigma}_{i j}^{\prime} \bar{e}_{i j}+\bar{\sigma}^{\prime} \bar{\varepsilon}+\bar{\sigma}_{i j, i}^{\prime} \bar{u}_{i}+\bar{\sigma}_{i}^{\prime} \bar{U}_{i}\right) d \Omega \\
=\int_{\Gamma}\left[\bar{T}_{i}^{\prime} \bar{u}_{i}+\bar{\sigma}^{\prime}\left(\bar{U}_{j}-\bar{u}_{j}\right) n_{j}\right] d \Gamma \tag{5.16}
\end{array}
$$

We subtract (5.16) from (5.15) and use the reciprocal relation (5.5) to get

$$
\begin{array}{r}
\int_{\Omega}\left[\left(\bar{\sigma}_{i j, j} \bar{u}_{i}^{\prime}-\bar{\sigma}_{i j, j}^{\prime} \bar{u}_{i}\right)+\left(\bar{\sigma}_{i} \bar{U}_{i}^{\prime}-\bar{\sigma}_{i,}^{\prime} \bar{U}_{i}\right)\right] d \Omega \\
=\int_{\Gamma}\left[\bar{T}_{i} \bar{u}_{i}^{\prime}-\bar{T}_{i}^{\prime} \bar{u}_{i}+\bar{\sigma}\left(\bar{U}_{j}^{\prime}-\bar{u}_{j}^{\prime}\right) n_{j}\right. \\
\left.-\bar{\sigma}^{\prime}\left(\bar{U}_{j}-\bar{u}_{j}\right) n_{j}\right] d \Gamma \tag{5.17}
\end{array}
$$

We now let the primed stresses and displacements correspond to certain special loading conditions prescribed via the body force per unit mass $X_{i}$. Referring to Figure (3.1) we note that the masses of the solid and fluid portions of the area $\delta A$ are respectively $(1-f) \rho_{s} \delta A$ and $f \rho_{f} \delta A$, where $\rho_{s}$ is the density of the soil grains and $\rho_{f}$ is the density of the pore water. The bulk density $\rho$ of the soil is given by :

$$
\begin{equation*}
\rho=\rho_{1}+\rho_{2} \tag{5.18}
\end{equation*}
$$

where $\rho_{1}=(1-f) \rho_{\mathrm{g}}$ and $\rho_{2}=f \rho_{f}$.
The body force on the mass $\rho \delta A$ shown in Figure (3.1) is

$$
\rho X_{i} \delta A \equiv\left(\rho_{1} X_{i}+\rho_{2} X_{i}\right) \delta A
$$

Here we interpret the terms $\rho_{1} X_{i}$ and $\rho_{2} X_{i}$ as the body forces (per unit volume) acting on the solid and fluid portions of soil respectively. We let the primed stresses and displacements correspond to the loading

$$
\begin{align*}
\rho_{1} X_{i}^{k} & =2 \pi \delta_{i k} \Delta(\underline{x}, p) e^{-i \omega t}  \tag{5.19}\\
\rho_{2} X_{i}^{k} & =0  \tag{5.20}\\
i, k & =1,2
\end{align*}
$$

Here $\Delta(\underline{x}, p)$ denotes the Dirac delta function, points $\underline{x}, p$ being on the solid skeleton. The superscript $k$ on the left hand side indicates the direction of the point load and equations (5.19), (5.20) thus describe two loading situations. This represents sinusoidal point loading on the solid skeleton and is analogous to the loading commonly used for deriving the fundamental solutions for elasticity problems (Brebbia et al.,1984). A similar approach was taken by Cheng and Liggett (1984) for Laplace transformed space. Since $X_{i}=\bar{X}_{i} e^{-i \omega t}$, we have from (5.19) and (5.20) :

$$
\begin{align*}
\rho_{1} \bar{X}_{i}^{k} & =2 \pi \delta_{i k} \Delta(x, p)  \tag{5.21}\\
\rho_{2} \bar{X}_{i}^{k} & =0  \tag{5.22}\\
i, k & =1,2
\end{align*}
$$

The stresses and displacements due to the loading ( 5.21 ) and (5.22) will be denoted by asterisks and referred to as fundamental quantities. Substituting (5.21) and (5.22) into the governing equations (5.6) and (5.7), and noting (5.18) we have

$$
\begin{align*}
\bar{\sigma}_{i j, j}^{* k} & =-i \omega f k\left(\bar{U}_{i}^{* k}-\bar{u}_{i}^{* k}\right)-2 \pi \delta_{i k} \Delta(x, p)  \tag{5.23}\\
\bar{\sigma}_{, i}^{* k} & =-i \omega f k\left(\bar{U}_{i}^{* k}-\bar{u}_{i}^{* k}\right) \tag{5.24}
\end{align*}
$$

Also, for an arbitrary body force $X_{i}=\bar{X}_{i} e^{-i \omega t}$ we have from (5.6) and (5.7)

$$
\begin{align*}
\bar{\sigma}_{i j, j} & =-\rho_{1} \bar{X}_{i}+i \omega f k\left(\bar{U}_{i}-\bar{u}_{i}\right)  \tag{5.25}\\
\bar{\sigma}_{, i} & =-\rho_{2} \bar{X}_{i}-i \omega f k\left(\bar{U}_{i}-\bar{u}_{i}\right) \tag{5.26}
\end{align*}
$$

We replace the primes in equation (5.17) by asterisks and use equations (5.23) to ( 5.26 ) to get

$$
\int_{\Omega}\left\{\left[-\rho_{1} \bar{X}_{i}+i \omega f k\left(\bar{U}_{i}-\bar{u}_{i}\right)\right] \bar{u}_{i}^{* k}-\left[i \omega f k\left(\bar{U}_{i}^{* k}-\bar{u}_{i}^{* k}\right)-2 \pi \delta_{i k} \Delta(x, p)\right] \bar{u}_{i}\right.
$$

$$
\begin{align*}
& \left.+\left[-\rho_{2} \bar{X}_{i}-i \omega f k\left(\bar{U}_{i}-\bar{u}_{i}\right)\right] \bar{U}_{i}^{* k}+i \omega f k\left(\bar{U}_{i}^{* k}-\bar{u}_{i}^{* k}\right) \bar{U}_{i}\right\} d \Omega \\
= & \int_{\Gamma}\left\{\left(\bar{T}_{i} \bar{u}_{i}^{* k}-\bar{T}_{i}^{* k} \bar{u}_{i}\right)+\bar{\sigma}\left(\bar{U}_{i}^{* k}-\bar{u}_{i}^{* k}\right) n_{i}-\bar{\sigma}^{* k}\left(\bar{U}_{i}-\bar{u}_{i}\right) n_{i}\right\} d \Gamma \tag{5.27}
\end{align*}
$$

By the sifting property of the delta function,

$$
\begin{equation*}
\int_{\Omega} 2 \pi \delta_{i k} \Delta(\underline{x}, p) \bar{u}_{i} d \Omega=2 \pi \delta_{i k} \bar{u}_{i}(p) \tag{5.28}
\end{equation*}
$$

where the integration has been performed with respect to $\underline{x}$, and $\underline{p}$ is the position vector of the point of application of the point load. Both sides of equation (5.27) are thus functions of $\boldsymbol{p} \in \Omega$ and this equation takes the form

$$
\begin{align*}
2 \pi \delta_{i k} \bar{u}_{i}(p)= & \int_{\Gamma}\left(\bar{T}_{i} \bar{u}_{i}^{* k}-\bar{T}_{i}^{* k} \bar{u}_{i}\right)+\bar{\sigma}\left(\bar{U}_{i}^{* k}-\bar{u}_{i}^{* k}\right) n_{i} \\
& -\bar{\sigma}^{* k}\left(\bar{U}_{i}-\bar{u}_{i}\right) n_{i} d \Gamma \\
& +\int_{\Omega} \bar{X}_{i}\left(\rho_{1} \bar{u}_{i}^{* k}+\rho_{2} \bar{U}_{i}^{* k}\right) d \Omega \tag{5.29}
\end{align*}
$$

In order to write ( 5.29 ) for a boundary point $\boldsymbol{p}$ it is necessary to evaluate the Cauchy principal value of integrals involving the fundamental quantities (denoted by asterisks), since the quantities are singular at $p$. This, operation results in modifications to the term in $\bar{u}_{k}(p)$ and the general form of this term is $2 \pi \alpha_{i k} \bar{u}_{i}(p)$, which resembles (5.28) in form.

Thus for any point $p \in \Gamma,(5.29)$ becomes

$$
\begin{align*}
2 \pi \alpha_{i k} \bar{u}_{i}(D)= & P V \int_{\Gamma}\left[\left(\bar{T}_{i} \bar{u}_{i}^{* k}-\bar{T}_{i}^{* k} \bar{u}_{i}\right)+\bar{\sigma}\left(\bar{U}_{i}^{* k}-\bar{u}_{i}^{* k}\right) n_{i}\right. \\
& \left.-\bar{\sigma}^{* k}\left(\bar{U}_{i}-\bar{u}_{i}\right) n_{i}\right] d \Gamma \\
& +P V \int_{\Omega} \bar{X}_{i}\left(\rho_{1} \bar{u}_{i}^{* k}+\rho_{2} \bar{U}_{i}^{* k}\right) d \Omega \tag{5.30}
\end{align*}
$$

The symbol $P V$ denotes the Cauchy principal value of an integral. The coef.cient $\alpha_{i k}$ must be evaluated either from the Cauchy principal value of the integrals or by using a "rigid body motion" analogy (Brebbia et al.,1984).

Equation (5.30) is the integral equation for the soil displacements. To derive a similar equation for pore pressure we consider the fundamental solutions corresponding to the loading

$$
\begin{aligned}
\rho X_{i} & =0 \\
\left(\rho_{2} X_{i}\right)_{, i} & =-2 \pi f k \Delta(\underline{x}, \underline{v}) e^{-i \omega t}
\end{aligned}
$$

i.e.

$$
\begin{align*}
\rho \bar{X}_{i} & =0  \tag{5.31}\\
\left(\rho_{2} \bar{X}_{i}\right)_{, i} & =-2 \pi f k \Delta(x, p) \tag{5.32}
\end{align*}
$$

Since $\rho=\rho_{1}+\rho_{2}$, equation ( 5.31 ) implies that $\rho_{2} \bar{X}_{i}=-\rho_{1} \bar{X}_{i}$, i.e. the body forces on the solid and fluid portions of soil cancel each other. Equation (5.32) describes the loading on the fluid portion of soil. We assume that there exists a scalar function $\chi$ such that

$$
\begin{equation*}
\rho_{2} \bar{X}_{i}=\chi, i \tag{5.33}
\end{equation*}
$$

and hence from (5.32) we have

$$
\begin{equation*}
\nabla^{2} \chi=-2 \pi f k \Delta(\underline{x}, \bar{y}) \tag{5.34}
\end{equation*}
$$

We shall denote the stresses and displacements corresponding to the loading (5.31), (5.32) by superscript 0. These are also called fundamental quantities. From the governing equations (5.6) and (5.7) and the loading described by (5.31) and (5.32) we have

$$
\begin{align*}
& \bar{\sigma}_{i j, j}^{0}=i \omega f k\left(\bar{U}_{i}^{0}-\bar{u}_{i}^{0}\right)+\chi_{, i}  \tag{5.35}\\
& \bar{\sigma}_{, i}^{0}=-i \omega f k\left(\bar{U}_{i}^{0}-\bar{u}_{i}^{0}\right)-\chi, i \tag{5.36}
\end{align*}
$$

Using equations (5.25), (5.26), (5.35) and (5.36) we have from (5.17)

$$
\begin{array}{r}
\int_{\Omega}\left\{\left[-\rho_{1} \bar{X}_{i}+i \omega f k\left(\bar{U}_{i}-\bar{u}_{i}\right)\right] \bar{u}_{i}^{0}-\left[i \omega f k\left(\bar{U}_{i}^{0}-\bar{u}_{i}^{0}\right)+\chi_{i}\right] \bar{u}_{i}\right. \\
\left.+\left[-\rho_{2} \bar{X}_{i}-i \omega f k\left(\bar{U}_{i}-\bar{u}_{i}\right)\right] \bar{U}_{i}^{0}+\left[i \omega f k\left(\bar{U}_{i}^{0}-\bar{u}_{i}^{0}\right)+\chi_{i}\right] \bar{U}_{i}\right\} d \Omega \\
=\int_{\Gamma}\left\{\left(\bar{T}_{i} \bar{u}_{i}^{0}-\bar{T}_{i}^{0} \bar{u}_{i}\right)+\bar{\sigma}\left(\bar{U}_{j}^{0}-\bar{u}_{j}^{0}\right) n_{j}\right. \\
\left.-\bar{\sigma}^{0}\left(\bar{U}_{j}-\bar{u}_{j}\right) n_{j}\right\} d \Gamma \tag{5.37}
\end{array}
$$

By collecting terms we find that the left side (L.H.S.) $\omega_{1}$ (5.37) reduces to

$$
\begin{equation*}
\text { L.H.S. of }(5.37)=\int_{\Omega}\left\{\bar{X}_{i}\left[-\rho \bar{u}_{i}^{0}-\rho_{2}\left(\bar{U}_{i}^{0}-\bar{u}_{i}^{0}\right)\right]+\chi_{, i}\left(\bar{U}_{i}-\bar{u}_{i}\right)\right\} d \Omega \tag{5.38}
\end{equation*}
$$

From (5.36) :

$$
\bar{U}_{i}^{0}-\bar{u}_{i}^{0}=\frac{i}{\omega f k}\left(\bar{\sigma}_{, i}^{0}+\chi_{, i}\right)
$$

From (5.7) :

$$
\bar{U}_{i}-\bar{u}_{i}=\frac{i}{\omega f k}\left(\bar{\sigma}_{i}+f \rho_{f} \bar{X}_{i}\right)
$$

Substituting these expressions into (5.38) gives

$$
\begin{equation*}
\text { L.H.S. of }(5.37)=\int_{\Omega}\left[-\bar{X}_{i}\left(\rho \bar{u}_{i}^{0}+\frac{i \rho_{f}}{\omega k} \bar{\sigma}_{, i}^{0}\right)+\frac{i \chi_{, i} \bar{\sigma}_{, i}}{\omega f k}\right] d \Omega \tag{5.39}
\end{equation*}
$$

By applying the divergence theorem to the product $\psi \eta_{i}$ it is easy to establish the identity

$$
\begin{equation*}
\int_{\Omega} \eta_{i} \psi, i d \Omega=-\int_{\Omega} \psi \eta_{i, i} d \Omega+\int_{\Gamma} \psi \eta_{i} n_{i} d \Gamma \tag{5.40}
\end{equation*}
$$

By putting $\eta_{i}=\chi, i$ and $\psi=\bar{\sigma}$ we have

$$
\begin{align*}
\int_{\Omega} \chi_{, i} \bar{\sigma}_{, i} d \Omega & =-\int_{\Omega} \bar{\sigma} \nabla^{2} \chi d \Omega+\int_{\Gamma} \bar{\sigma} \chi_{, i} n_{i} d \Gamma \\
& =2 \pi f k \int_{\Omega} \bar{\sigma} \Delta(\underline{x}, p) d \Omega+\int_{\Gamma} \bar{\sigma} \chi_{, i} n_{i} d \Gamma \\
& =2 \pi f k \bar{\sigma}(p)+\int_{\Gamma} \bar{\sigma} \chi_{, i} n_{i} d \Gamma \tag{5.41}
\end{align*}
$$

using (5.34) and the sifting property of the delta function.
We now substitute ( 5.41 ) into (5.39) :

$$
\begin{array}{r}
\text { L.H.S. of }(5.37)=\int_{\Omega}-\bar{X}_{i}\left(\rho \bar{u}_{i}^{0}+\frac{i \rho_{f}}{\omega k} \bar{\sigma}_{, i}^{0}\right) d \Omega \\
+\frac{2 \pi i}{\omega} \bar{\sigma}(p)+\frac{i}{\omega f k} \int_{\Gamma} \bar{\sigma} \chi_{, i} n_{i} d \Gamma \tag{5.42}
\end{array}
$$

Equation (5.37) thus becomes

$$
\begin{align*}
\frac{2 \pi i}{\omega} \bar{\sigma}(p)= & \int_{\Gamma}\left[\left(\bar{T}_{j} \bar{u}_{j}^{0}-\bar{T}_{j}^{0} \bar{u}_{j}\right)+\bar{\sigma}\left(\bar{U}_{j}^{0}-\bar{u}_{j}^{0}\right) n_{j}\right. \\
& \left.-\bar{\sigma}^{0}\left(\bar{U}_{j}-\bar{u}_{j}\right) n_{j}-\frac{i}{\omega f k} \bar{\sigma} \chi_{, j} n_{j}\right] d \Gamma \\
& +\int_{\Omega} \bar{X}_{j}\left(\rho \bar{u}_{j}^{0}+\frac{i \rho_{f}}{\omega k} \bar{\sigma}_{j,}^{0}\right) d \Omega  \tag{5.43}\\
j= & 1,2
\end{align*}
$$

If equation ( 5.43 ) is written for a boundary point $p$ the singular integrals involving fundamental quantities must be interpreted as Cauchy principal values and extra terms in $\bar{\sigma}(\underline{p})$ appear. Thus for a boundary point $p$ :

$$
\begin{align*}
\frac{2 \pi i \theta}{\omega} \bar{\sigma}(p)= & P V \int_{\Gamma}\left[\left(\bar{T}_{j} \bar{u}_{j}^{0}-\bar{T}_{j}^{0} \bar{u}_{j}\right)+\bar{\sigma}\left(\bar{U}_{j}^{0}-\bar{u}_{j}^{0}\right) n_{j}\right. \\
& \left.-\bar{\sigma}^{0}\left(\bar{U}_{j}-\bar{u}_{j}\right) n_{j}-\frac{i}{\omega f k} \bar{\sigma} \chi_{, j} n_{j}\right] d \Gamma \\
& +P V \int_{\Omega} \bar{X}_{j}\left(\rho \bar{u}_{j}^{0}+\frac{i \rho_{j}}{\omega k} \bar{\sigma}_{, j}^{0}\right) d \Omega  \tag{5.44}\\
j= & 1,2
\end{align*}
$$

The coefficient $\theta$ must be determined from the evaluation of the Cauchy principal value of the singular integrals or from a "rigid body motion" analogy. For the purpose of determining wave-induced stresses and pore pressures we assume that there are no body forces present. Thus from ( 5.30 ) and ( 5.44 ) we write the
boundary integral equations as for $\boldsymbol{p} \in \Gamma$ as

$$
\begin{align*}
2 \pi \alpha_{i k} \bar{u}_{i}(p)= & P V \int_{\Gamma}\left[\left(\bar{T}_{i} \bar{u}_{i}^{* k}-\bar{T}_{i}^{* k} \bar{u}_{i}\right)+\bar{\sigma}\left(\bar{U}_{i}^{* k}-\bar{u}_{i}^{* k}\right) n_{i}\right. \\
& \left.-\bar{\sigma}^{* k}\left(\bar{U}_{i}-\bar{u}_{i}\right) n_{i}\right] d \Gamma  \tag{5.45}\\
\frac{2 \pi i \theta}{\omega} \bar{\sigma}(p)= & P V \int_{\Gamma}\left[\left(\bar{T}_{i} \bar{u}_{i}^{0}-\bar{T}_{i}^{0} \bar{u}_{i}\right)+\bar{\sigma}\left(\bar{U}_{i}^{0}-\bar{u}_{i}^{0}\right) n_{i}\right. \\
& \left.-\bar{\sigma}^{0}\left(\bar{U}_{i}-\bar{u}_{i}\right) n_{i}-\frac{i}{\omega f k} \bar{\sigma} \chi_{i,} n_{i}\right] d \Gamma \tag{5.46}
\end{align*}
$$

where subscripts $i, k=1,2$, and as previously noted, a repeated subscript indicates summation.

### 5.2 The Fundamental Solutions

As indicated earlier, the fundamental solutions are the stresses and displacements corresponding to the loadings described by equations (5.21), (5.22) and (5.31), (5.32) . In order to determine these solutions the governing equations (5.6) and (5.7) must be written in terms of displacements via the constitutive relations (3.22) and (3.23). This renders (5.6) and (5.7) as

$$
\begin{align*}
\mu \bar{u}_{i, j j}+\left(\lambda+\mu+\frac{Q^{2}}{R}+Q\right) \bar{u}_{j, j i}+(Q+R) \bar{U}_{j, j i} & =-\rho \bar{X}_{i}  \tag{5.47}\\
Q \bar{u}_{j, j i}+R \bar{U}_{j, j i}+f \rho_{f} \bar{X}_{i} & =-i \omega f k\left(\bar{U}_{i}-\bar{u}_{i}\right) \tag{5.48}
\end{align*}
$$

where, we recall, $Q$ and $R$ are given by (3.24). The first set of fundamental solutions, denoted by asterisks, is obtained by substituting (5.21) and (5.22) into (5.47) and (5.48) to give

$$
\begin{align*}
\mu \bar{u}_{i, j j}^{* k}+\left(\lambda+\mu+\frac{Q^{2}}{R}+Q\right) \bar{u}_{j, j i}^{* k}+(Q+R) \bar{U}_{j, j i}^{* k} & =-2 \pi \delta_{i k} \Delta(x, p)  \tag{5.49}\\
Q \bar{u}_{j, j i}^{* k}+R \bar{U}_{j, j i}^{* k}+i \omega f k\left(\bar{U}_{i}^{* k}-\bar{u}_{i}^{\nabla k}\right) & =0 \tag{5.50}
\end{align*}
$$

Following Biot (1956) we define

$$
\begin{equation*}
\bar{v}_{i}^{* k}=\left(\frac{H-Q-R}{H}\right) \bar{u}_{i}^{* k}+\left(\frac{Q+R}{H}\right) \bar{U}_{i}^{* k} \tag{5.51}
\end{equation*}
$$

where

$$
\begin{equation*}
H=\lambda+2 \mu+\frac{(Q+R)^{2}}{R} \tag{5.52}
\end{equation*}
$$

and

$$
\begin{equation*}
\bar{w}_{i}^{\bullet k}=\bar{U}_{i}^{* k}-\bar{u}_{i}^{* k} \tag{5.53}
\end{equation*}
$$

from which we have

$$
\begin{align*}
\bar{u}_{i}^{* k} & =\bar{v}_{i}^{* k}-\left(\frac{Q+R}{H}\right) \bar{w}_{i}^{* k}  \tag{5.54}\\
\bar{U}_{i}^{* k} & =\bar{v}_{i}^{* k}+\left(1-\frac{Q+R}{H}\right) \bar{w}_{i}^{* k} \tag{5.55}
\end{align*}
$$

In terms of $\bar{v}_{i}^{* k}$ and $\bar{w}_{i}^{* k}$ equations (5.49) and (5.50) become

$$
\begin{align*}
\mu \bar{v}_{i, j j}^{* k}+(H-\mu) \bar{v}_{j, j i}^{* k}-\frac{\mu(Q+R)}{H}\left(\bar{w}_{i, j j}^{* k}-\bar{w}_{j, j i}^{* k}\right) & =-2 \pi \delta_{i k} \Delta(x, \underline{x})  \tag{5.56}\\
(Q+R) \bar{v}_{j, j i}^{* k}+\frac{R(\lambda+2 \mu)}{H} \bar{w}_{j, j i}^{* k}+i \omega f k \bar{w}_{i}^{* k} & =0 \tag{5.57}
\end{align*}
$$

But we note from (5.24) that

$$
\bar{w}_{i}^{* k}=\frac{\bar{\sigma}_{i .}^{* k}}{-i \omega f k}
$$

so that

$$
\bar{w}_{j, j i}^{* k}=\bar{w}_{i, j j}^{* k}
$$

Substituting this into (5.56) and (5.57) gives

$$
\begin{align*}
\mu \bar{v}_{i, j j}^{* k}+(H-\mu) \bar{v}_{j, j i}^{\otimes k} & =-2 \pi \delta_{i k} \Delta(\boldsymbol{x}, p)  \tag{5.58}\\
(Q+R) \tilde{v}_{j, j i}^{* k}+\frac{R(\lambda+2 \mu)}{H} \bar{w}_{i, j j}^{* k}+i \omega f k \bar{w}_{i}^{* k} & =0 \tag{5.59}
\end{align*}
$$

Equation (5.58) may be compared to the elasticity equation

$$
\begin{equation*}
\mu u_{i, j j}^{k}+(\lambda+\mu) u_{j, j i}^{k}=-2 \pi \delta_{i k} \Delta(\underline{x}, p) \tag{5.60}
\end{equation*}
$$

Employing the Papkovitch-Neuber technique we write

$$
\begin{equation*}
u_{i}^{k}=\frac{\partial}{\partial z_{i}}\left(\psi_{0}^{k}+z_{j} \psi_{j}^{k}\right)-4(1-\nu) \psi_{i}^{k} \tag{5.61}
\end{equation*}
$$

where

$$
\begin{equation*}
\nu=\frac{\lambda}{2(\lambda+\mu)} \tag{5.62}
\end{equation*}
$$

The quantities $\psi_{0}^{k}, \psi_{i}^{k}$ are functions of the spatial coordinates $z_{i}$. We may take any one of these functions to be zero without loss of completeness (Eubanks and Sternberg, 1956). Accordingly, we set $\psi_{0}^{k} \equiv 0$. We define the coordinate system by

$$
z_{i}=x_{i}-p_{i}
$$

where $x_{i}, p_{i}$ are the coordinates of the points $\underline{x}$ and $\underline{p}$ respectively. Using (5.61) and writing

$$
\begin{equation*}
\beta=\frac{2 \mu(1-\nu)}{1-2 \nu} \tag{5.63}
\end{equation*}
$$

we find that ( 5.60 ) becomes

$$
\beta \frac{\partial}{\partial x_{i}}\left[\left(x_{j}-p_{j}\right) \nabla^{2} \psi_{j}^{k}\right]-2(1-2 \nu) \beta \nabla^{2} \psi_{i}^{k}=-2 \pi \delta_{i k} \Delta(x, p)
$$

We now integrate over a circular region $\Omega$,centre $p$ and radius $\epsilon$ :

$$
\begin{equation*}
\beta \int_{\Omega} \frac{\partial}{\partial x_{i}}\left[\left(x_{j}-p_{j}\right) \nabla^{2} \psi_{j}^{k}\right] d \Omega-2(1-2 \nu) \beta \int_{\Omega} \nabla^{2} \psi_{i}^{k} d \Omega=-2 \pi \delta_{i k} \tag{5.64}
\end{equation*}
$$

since

$$
\int_{\Omega} \Delta(x, p) d \Omega=1
$$

We may define local polar coordinates with $p$ as origin by letting

$$
\begin{aligned}
& x_{1}-p_{1}=r \cos \theta \\
& x_{2}-p_{2}=r \sin \theta
\end{aligned}
$$

Then, assuming $\psi_{i}^{k}$ is independent of $\theta$, we find that the second integral on the left hand side of ( 5.64 ) is easily evaluated as

$$
\begin{align*}
\int_{\Omega} \nabla^{2} \psi_{i}^{k} d \Omega & =\int_{\Omega}\left(\psi_{i, j}^{k}\right)_{, j} d \Omega \\
& =\int_{\Gamma} \psi_{i, j}^{k} n_{j} d \Gamma \\
& =\int_{\Gamma} \frac{\partial \psi_{i}^{k}}{\partial n} d \Gamma \\
& =\int_{0}^{2 \pi}\left(\frac{\partial \psi_{i}^{k}}{\partial r}\right)_{r=e} \epsilon d \theta \\
& =2 \pi \epsilon\left(\frac{\partial \psi_{i}^{k}}{\partial r}\right)_{r=e} \tag{5.65}
\end{align*}
$$

where $\Gamma$ is the circular boundary of $\Omega$.
The first integral on the left hand side of (5.64) is evaluated by noting that for any continuously differentiable function $\phi$

$$
\int_{\Omega} \frac{\partial \phi}{\partial x_{i}} d \Omega=\int_{\Gamma} \phi n_{i} d \Gamma
$$

Thus

$$
\begin{equation*}
\int_{\Omega} \frac{\partial}{\partial x_{i}}\left[\left(x_{j}-p_{j}\right) \nabla^{2} \psi_{j}^{k}\right] d \Omega=\int_{\Gamma}\left(x_{j}-p_{j}\right) \nabla^{2} \psi_{j}^{k} n_{i} d \Gamma \tag{5.66}
\end{equation*}
$$

Changing to polar coordinates and recalling that $\psi_{i}^{k}$ is independent of $\theta,(5.66)$ becomes

$$
\int_{\Omega} \frac{\partial}{\partial x_{i}}\left[\left(x_{j}-p_{j}\right) \nabla^{2} \psi_{j}^{k}\right] d \Omega
$$

$$
\begin{align*}
& =\int_{\Gamma}\left[r \cos \theta \nabla^{2} \psi_{1}^{k}+r \sin \theta \nabla^{2} \psi_{2}^{k}\right]\binom{\cos \theta}{\sin \theta} d \Gamma \\
& =\left[\begin{array}{l}
\int_{0}^{2 \pi}\left[\epsilon \cos \theta\left(\nabla^{2} \psi_{1}^{k}\right)_{r=e}+\epsilon \sin \theta\left(\nabla^{2} \psi_{2}^{k}\right)_{r=e}\right] \cos \theta \cdot \epsilon d \theta \\
\int_{0}^{2 \pi}\left[\epsilon \cos \theta\left(\nabla^{2} \psi_{1}^{k}\right)_{r=e}+\epsilon \sin \theta\left(\nabla^{2} \psi_{2}^{k}\right)_{r=e}\right] \sin \theta . \epsilon d \theta
\end{array}\right] \\
& =\pi \epsilon^{2}\left[\begin{array}{l}
\left(\nabla^{2} \psi_{1}^{k}\right)_{r=e} \\
\left(\nabla^{2} \psi_{2}^{k}\right)_{r=e}
\end{array}\right]=\pi \epsilon^{2}\left(\nabla^{2} \psi_{i}^{k}\right)_{r=e} \tag{5.67}
\end{align*}
$$

Then using (5.65) and (5.67), equation (5.64) simplifies to

$$
\begin{equation*}
\beta \pi \epsilon^{2}\left(\nabla^{2} \psi_{i}^{k}\right)_{r=e}-4(1-2 \nu) \beta \pi \epsilon\left(\frac{\partial \psi_{i}^{k}}{\partial r}\right)_{r=e}=-2 \pi \delta_{i k} \tag{5.68}
\end{equation*}
$$

We choose $\psi_{i}^{k}$ such that

$$
\begin{equation*}
\nabla^{2} \psi_{i}^{k}=0 \text { at least when } x \neq p \text { i.e. } r \neq 0 \tag{5.69}
\end{equation*}
$$

Then (5.68) becomes

$$
\begin{equation*}
2(1-2 \nu) \beta\left(\frac{\partial \psi_{i}^{k}}{\partial r}\right)_{r=e}=\frac{1}{\epsilon} \delta_{i k} \tag{5.70}
\end{equation*}
$$

This equation is satisfied by

$$
\begin{align*}
\psi_{i}^{k} & =\frac{1}{2(1-2 \nu) \beta} \ln r \delta_{i k} \\
& =\frac{1}{4 \mu(1-\nu)} \ln r \delta_{i k} \tag{5.71}
\end{align*}
$$

which also satifies condition (5.69). We thus have from (5.61) a solution of (5.60) as

$$
\begin{align*}
u_{i}^{k} & =\frac{1}{4 \mu(1-\nu)}\left(z_{j} \ln r \delta_{j k}\right)_{, i}-\frac{1}{\mu} \ln r \delta_{i k} \\
& =\frac{1}{4 \mu(1-\nu)}\left[r_{, i} r_{, k}+(-3+4 \nu) \ln r \delta_{i k}\right] \tag{5.72}
\end{align*}
$$

where we have used $r^{2}=z_{i} z_{i}, z_{i, j}=\delta_{i j}$ and $z_{i}=r r_{i,}$.
By comparing (5.60) with (5.58) we have the analogy :

$$
\begin{aligned}
u_{i}^{k} & \longrightarrow \bar{v}_{i}^{* k} \\
\mu & \longrightarrow \mu \\
\lambda+\mu & \longrightarrow H-\mu
\end{aligned}
$$

Hence

$$
\begin{aligned}
\lambda & \rightarrow H-2 \mu \\
\nu=\frac{\lambda}{2(\lambda+\mu)} & \rightarrow \frac{H-2 \mu}{2(H-\mu)}
\end{aligned}
$$

We thus construct from (5.72) the function $\bar{v}_{i}^{* k}$ as

$$
\begin{equation*}
\bar{v}_{i}^{* k}=\left(\frac{H-\mu}{2 \mu H}\right) r_{, i} r_{, k}-\left(\frac{H+\mu}{2 \mu H}\right) \ln r \delta_{i k} \tag{5.73}
\end{equation*}
$$

To determine $\bar{w}_{i}^{* k}$ (defined by 5.53 ) we again note from (5.24) that $\bar{w}_{i}^{* k}$ is the gradient of a scalar function. Writing

$$
\begin{equation*}
\bar{w}_{i}^{* k}=\phi_{, i}^{k} \tag{5.74}
\end{equation*}
$$

we have from (5.59)

$$
(Q+R) \frac{\partial}{\partial x_{i}}\left(\bar{v}_{j, j}^{* k}\right)+\frac{R(\lambda+2 \mu)}{H} \frac{\partial}{\partial x_{i}}\left(\nabla^{2} \phi^{k}\right)+i \omega f k \frac{\partial}{\partial x_{i}}\left(\phi^{k}\right)=0
$$

Integrating with respect to $x_{i}$ and setting the integration constant equal to zero (since we are only interested in a particular solution) gives

$$
\begin{equation*}
(Q+R) \bar{v}_{j, j}^{*}+\frac{R(\lambda+2 \mu)}{H} \nabla^{2} \phi^{k}+i \omega f k \phi^{k}=0 \tag{5.75}
\end{equation*}
$$

To evaluate $\bar{v}_{j, j}^{* k}$ we make use of the following results:

$$
\begin{aligned}
r_{, i} & =\frac{x_{i}-p_{i}}{r} \\
r_{, i} r_{, i} & =1 \\
r_{, i j} & =\frac{\delta_{i j}}{r}-\frac{r_{, i} r_{, j}}{r} \\
\left(r_{, j} r_{, k}\right)_{, j} & =\frac{r_{, k}}{r}
\end{aligned}
$$

which are easily deduced from the definition

$$
r^{2}=\left(x_{i}-p_{i}\right)\left(x_{i}-p_{i}\right)
$$

Then differentiating $\bar{v}_{j}^{-k}$ (from (5.73)) with respect to $x_{j}$ we get

$$
\bar{v}_{i, i}^{* k}=-\frac{1}{H}(\ln r)_{, k}
$$

Substituting in (5.75) gives

$$
\begin{equation*}
\nabla^{2} \phi^{k}+\frac{i \omega f k H}{R(\lambda+2 \mu)} \phi^{k}=\left[\frac{Q+R}{R(\lambda+2 \mu)}\right](\ln r)_{, k} \tag{5.76}
\end{equation*}
$$

We let

$$
\begin{equation*}
\phi^{k}=\phi_{1}^{k}+\phi_{2}^{k} \tag{5.77}
\end{equation*}
$$

where

$$
\begin{align*}
\phi_{1}^{k} & =\left(\frac{Q+R}{i \omega f k H}\right)(\ln r)_{, k} \\
& =\frac{-i(Q+R)}{w f k H} \frac{r_{, k}}{r} \tag{5.78}
\end{align*}
$$

Since (see Appendix D)

$$
\nabla^{2}(\ln r)=2 \pi \Delta(x, p)
$$

we have

$$
\begin{align*}
\nabla^{2} \phi_{1}^{k} & =\left(\frac{Q+R}{i \omega f k H}\right)\left(\nabla^{2} \ln r\right)_{, k} \\
& =\frac{2 \pi(Q+R)}{i \omega f k H} \frac{\partial}{\partial x_{k}}[\Delta(x, p)] \tag{5.79}
\end{align*}
$$

where the right hand side must be interpreted as the derivative of a generalized function. Substituting ( 5.77 ) into ( 5.76 ) and then using ( 5.78 ) and ( 5.79 ) gives

$$
\begin{equation*}
\nabla^{2} \phi_{2}^{k}+\frac{i \omega f k H}{R(\lambda+2 \mu)} \phi_{2}^{k}=-\frac{2 \pi(Q+R)}{i \omega f k H} \frac{\partial}{\partial x_{k}}[\Delta(x, p)] \tag{5.80}
\end{equation*}
$$

We define the constant $a$ by

$$
\begin{equation*}
a^{2}=-\frac{i \omega f k H}{R(\lambda+2 \mu)} \tag{5.81}
\end{equation*}
$$

This will be expressed in a more convenient form later. Equation ( 5.80 ) becomes

$$
\begin{equation*}
\nabla^{2} \phi_{2}^{k}-a^{2} \phi_{2}^{k}=\frac{2 \pi i(Q+R)}{\omega f k H} \frac{\partial}{\partial x_{k}}[\Delta(\underline{x}, \underline{p})] \tag{5.82}
\end{equation*}
$$

Integration with respect to $x_{k}$ gives

$$
\begin{equation*}
\nabla^{2} \eta^{k}-a^{2} \eta^{k}=2 \pi \Delta(\underline{x}, p) \tag{5.83}
\end{equation*}
$$

where

$$
\begin{equation*}
\eta^{k}=\frac{\omega f k H}{i(Q+R)} \int \phi_{2}^{k} d x_{k}(\text { no sum on } k) \tag{5.84}
\end{equation*}
$$

A solution to (5.83) is (Bleistein, 1984)

$$
\begin{equation*}
\eta^{k}=-K_{0}(a r) \tag{5.85}
\end{equation*}
$$

where $K_{0}$ is the modified Bessel function of the second kind of order zero, and $r=|x-p|$. From (5.83) and (5.84) it is clear that

$$
\begin{align*}
\phi_{2}^{k} & =\frac{-i(Q+R)}{\omega f k H} \frac{\partial}{\partial x_{k}}\left[K_{0}(a r)\right] \\
& =\frac{i(Q+R)}{\omega f k H} a r_{, k} K_{1}(a r) \tag{5.86}
\end{align*}
$$

where $K_{1}$ is the modified Bessel function of the second kind of oruer 1. Now substituting (5.77) and (5.85) into (5.76) we have

$$
\begin{equation*}
\phi^{k}=\frac{i(Q+R)}{\omega f k H} a r_{, k}\left[K_{1}(a r)-\frac{1}{a r}\right] \tag{5.87}
\end{equation*}
$$

The function $\bar{w}_{i}^{p k}$ may now be obtained from (5.74) as

$$
\begin{align*}
\bar{w}_{i}^{* k}= & \frac{i(Q+R) a^{2}}{\omega f k H}\left\{\left[\frac{2}{a^{2} r^{2}}-K_{2}(a r)\right] r_{, i} r_{, k}\right. \\
& \left.+\delta_{i k}\left[\frac{K_{1}(a r)}{a r}-\frac{1}{a^{2} r^{2}}\right]\right\} \tag{5.88}
\end{align*}
$$

where we have used the results

$$
\begin{aligned}
K_{2}(a r) & =K_{0}(a r)+\frac{2}{a r} K_{1}(a r) \\
\frac{\partial}{\partial x_{i}}\left[K_{1}(a r)\right] & =-\frac{1}{2} a r_{, i}\left[K_{0}(a r)+K_{2}(a r)\right]
\end{aligned}
$$

Using (5.73) and (5.88) we recover $\bar{u}_{i}^{* k}$ and $\bar{U}_{i}^{* k}$ from (5.54) and (5.55) :

$$
\begin{align*}
\bar{u}_{i}^{* k}= & \left(\frac{H-\mu}{2 \mu H}\right) r_{, i} r_{, k}-\left(\frac{H+\mu}{2 \mu H}\right) \ln r \delta_{i k} \\
& -\frac{i(Q+R)^{2} a^{2}}{\omega f k H^{2}}\left\{\left[\frac{2}{a^{2} r^{2}}-K_{2}(a r)\right] r_{, i} r_{, k}\right. \\
& \left.+\delta_{i k}\left[\frac{K_{1}(a r)}{a r}-\frac{1}{a^{2} r^{2}}\right]\right\}  \tag{5.89}\\
\bar{U}_{i}^{* k}= & \left(\frac{H-\mu}{2 \mu H}\right) r_{, i} r_{, k}-\left(\frac{H+\mu}{2 \mu H}\right) \ln r \delta_{i k} \\
+ & \frac{i(H-Q-R)(Q+R) a^{2}}{\omega f k H^{2}}\left\{\left[\frac{2}{a^{2} r^{2}}-K_{2}(a r)\right] r_{, i} r_{, k}\right. \\
+ & \left.\delta_{i k[ }\left[\frac{K_{1}(a r)}{a r}-\frac{1}{a^{2} r^{2}}\right]\right\} \tag{5.90}
\end{align*}
$$

The tractions $\bar{T}_{i}^{* k}$ and $\bar{\sigma}^{* k}$ are now obtained from (5.8) and the constitutive laws (3.22) and (3.23) i.e.

$$
\begin{align*}
\bar{T}_{i}^{* k} & =2 \mu \bar{e}_{i j}^{* k} n_{j}+\left(\lambda+\frac{Q^{2}}{R}+Q\right) \bar{e}^{* k} n_{i}+(Q+R) \bar{\varepsilon}^{* k} n_{i}  \tag{5.91}\\
\bar{\sigma}^{* k} & =Q \bar{e}^{* k}+R \bar{e}^{* k} \tag{5.92}
\end{align*}
$$

where

$$
\begin{aligned}
& \bar{e}_{i j}^{* k}=\frac{1}{2}\left(\bar{u}_{i, j}^{* k}+\bar{u}_{j, i}^{* k}\right) \\
& \bar{e}^{* k}=\bar{u}_{i, i}^{* k} \\
& \bar{\varepsilon}^{* k}=\bar{U}_{i, i}^{* k}
\end{aligned}
$$

Thus substituting (5.89) and (5.90) into (5.91) and (5.92) we obtain after some manipulations

$$
\begin{align*}
\bar{T}_{i}^{* k}= & -2\left(\frac{H-\mu}{H}\right) \frac{r_{i,} r_{, k}}{r} \frac{\partial r}{\partial n}+\frac{\mu}{H} \frac{1}{r}\left(r_{, k} n_{i}-r_{, i} n_{k}-\frac{\partial r}{\partial n} \delta_{i k}\right) \\
& -\frac{i 2 \mu(Q+R)^{2} a^{3}}{\omega f k H^{2}}\left\{\frac{\partial r}{\partial n} \delta_{i k}\left[\frac{2}{a^{3} r^{3}}-\frac{K_{2}(a r)}{a r}\right]\right. \\
& +\frac{\partial r}{\partial n} r_{, i} r_{, k}\left[K_{3}(a r)-\frac{8}{a^{3} r^{3}}\right] \\
& +r_{, k} n_{i}\left[\frac{2}{a^{3} r^{3}}+\frac{3 K_{2}(a r)}{a r}-K_{3}(a r)\right] \\
& \left.+r_{, i} n_{k}\left[\frac{2}{a^{3} r^{3}}-\frac{K_{2}(a r)}{a r}\right]\right\}  \tag{5.93}\\
\bar{\sigma}^{* k}= & a\left(\frac{Q+R}{H}\right) r_{, k}\left[K_{1}(a r)-\frac{1}{a r}\right] \tag{5.94}
\end{align*}
$$

In these derivations the following relations were used

$$
\begin{aligned}
& K_{n+1}(z)=K_{n-1}(z)+\frac{2 n}{z} K_{n}(z) \\
& K_{n}^{\prime}(z)=-\frac{1}{2}\left[K_{n-1}(z)+K_{n+1}(z)\right]
\end{aligned}
$$

The second set of fundamental solutions, denoted by superscript 0 , is obtained by substituting ( 5.31 ) and (5.33) into (5.47) and (5.48). This gives

$$
\begin{array}{r}
\mu \nabla^{2} \bar{u}_{i}^{0}+\left(\lambda+\mu+\frac{Q^{2}}{R}+Q\right) \bar{u}_{j, j i}^{0}+(Q+R) \bar{U}_{j, j i}^{0}=0 \\
Q \bar{u}_{j, j i}^{0}+R \bar{U}_{j, j i}^{0}+\chi, i+i \omega f k\left(\bar{U}_{i}^{0}-\bar{u}_{i}^{0}\right)=0 \tag{5.96}
\end{array}
$$

where $\chi$ is found from (5.34). The solution of (5.34) is (see Appendix D)

$$
\begin{equation*}
x=-f k \ln r \tag{5.97}
\end{equation*}
$$

We define

$$
\begin{align*}
\bar{v}_{i}^{0} & =\bar{u}_{i}^{0}+\left(\frac{Q+R}{H}\right)\left(\bar{U}_{i}^{0}-\bar{u}_{i}^{0}\right)  \tag{5.98}\\
\bar{w}_{i}^{0} & =\bar{U}_{i}^{0}-\bar{u}_{i}^{0} \tag{5.99}
\end{align*}
$$

from which we have

$$
\begin{align*}
\bar{u}_{i}^{0} & =\bar{v}_{i}^{0}-\left(\frac{Q+R}{H}\right) \bar{w}_{i}^{0}  \tag{5.100}\\
\bar{U}_{i}^{0} & =\bar{v}_{i}^{0}+\left(1-\frac{Q+R}{H}\right) \bar{w}_{i}^{0} \tag{5.101}
\end{align*}
$$

From (5.36) we note that

$$
\bar{w}_{i}^{0}=\left(\frac{\bar{\sigma}^{0}+\chi_{, i}}{-i \omega f k}\right)_{, i}
$$

which implies that

$$
\begin{equation*}
\nabla^{2} \bar{w}_{i}^{0}=\bar{w}_{j, j i}^{0} \tag{5.102}
\end{equation*}
$$

Using (5.100), (5.101) and (5.102) we express (5.95) and (5.96) in terms of $\bar{v}_{i}^{0}, \bar{w}_{i}^{0}$ :

$$
\begin{align*}
\mu \nabla^{2} \bar{v}_{i}^{0}+(H-\mu) \bar{v}_{j, j i}^{0} & =0  \tag{5.103}\\
(Q+R) \bar{v}_{j, j i}^{0}+\frac{R(\lambda+2 \mu)}{H} \nabla^{2} \bar{w}_{i}^{0}+i \omega f k \bar{w}_{i}^{o}-\frac{f k r_{, i}}{r} & =0 \tag{5.104}
\end{align*}
$$

We note that $\bar{v}_{i}^{0}=0$ is a solution of (5.103), and (5.104) then becomes

$$
\begin{equation*}
\nabla^{2} \bar{w}_{i}^{0}-a^{2} \bar{w}_{i}^{0}+\frac{a^{2}}{i \omega} \frac{r_{i}}{r}=0 \tag{5.105}
\end{equation*}
$$

where we have used (5.81). We let

$$
\begin{equation*}
\bar{w}_{i}^{0}=\frac{r_{i, i}}{i \omega r}+\eta_{i} \tag{5.106}
\end{equation*}
$$

By using the results

$$
\begin{gathered}
r_{, i j}=\frac{\delta_{i j}}{r}-\frac{r_{, i}^{m}, j}{r} ; r_{, j j}=\frac{1}{r} ; r_{, j} r_{, j}=1 \\
\left(r_{, i} r_{, j}\right)_{, j}=\frac{r_{, i}}{r} ; r_{, i j j}=-\frac{r_{, i}}{r^{2}}
\end{gathered}
$$

it is easily shown that

$$
\nabla^{2}\left(\frac{r, i}{r}\right)=0
$$

Hence from (5.106) we have

$$
\begin{equation*}
\nabla^{2} \bar{w}_{i}^{n}=\nabla^{2} \eta_{i} \tag{5.107}
\end{equation*}
$$

We substitute (5.106) and (5.107) inio (5.105) to find

$$
\begin{equation*}
\nabla^{2} \eta_{i}-a^{2} \eta_{i}=0 \tag{5.108}
\end{equation*}
$$

From (5.106) we see that it would be convenient to seek a solution of (5.108) of the form

$$
\eta_{i}=r_{, i} f(r)
$$

Substituting this into (5.108) gives

$$
r^{2} \frac{\mathrm{~d}^{2} f}{\mathrm{~d} r^{2}}+r \frac{\mathrm{~d} f}{\mathrm{~d} r}-\left(a^{2} r^{2}+1\right) f=0
$$

Putting $x=a r$ and $g(x)=f\left(\frac{x}{a}\right)$ this equation becomes

$$
\begin{equation*}
x^{2} \frac{\mathrm{~d}^{2} g}{\mathrm{~d} x^{2}}+x \frac{\mathrm{~d} g}{\mathrm{~d} x}-\left(x^{2}+1\right) g=0 \tag{5.109}
\end{equation*}
$$

We recognise (5.109) as the modified Bessel equation of order 1 and hence we can write a solution as

$$
g(x)=K_{1}(x)
$$

Thus $f(r)=K_{1}(a r)$ and we write a solution of (5.108) as

$$
\eta_{i}=r_{, i} K_{1}(a r)
$$

In Appendix E we verify by direct substitution that this is a solution of (5.108).
It is clear that

$$
\begin{equation*}
\eta_{i}=\frac{-a r_{, i}}{i \omega} K_{1}(a r) \tag{5.110}
\end{equation*}
$$

also satisfies (5.108) so that from (5.106) we have

$$
\begin{equation*}
\bar{w}_{i}^{0}=\frac{-a r_{, i}}{i \omega}\left[K_{1}(a r)-\frac{1}{a r}\right] \tag{5.111}
\end{equation*}
$$

The fundamental solutions $\bar{u}_{i}^{0}$ and $\tilde{U}_{i}^{0}$ can now be written from (5.100) and (5.101):

$$
\begin{align*}
\bar{u}_{i}^{0} & =\left(\frac{Q+R}{i \omega H}\right) a r_{, i}\left[K_{1}(a r)-\frac{1}{a r}\right]  \tag{5.112}\\
\bar{U}_{i}^{0} & =\left(\frac{Q r R}{H}-1\right) \frac{a r_{, i}}{i \omega}\left[K_{1}(a r)-\frac{1}{a r}\right] \tag{5.113}
\end{align*}
$$

since $\bar{v}_{i}^{0}=0$.
Equation (5.8) and the constitutive laws (3.22), (3.23) imply

$$
\begin{align*}
& \tilde{T}_{i}^{0}=2 \mu \bar{e}_{i j}^{0} n_{j}+\left(\lambda+\frac{Q^{2}}{R}+Q\right) \bar{e}^{0} n_{i}+(Q+R) \bar{\varepsilon}^{0} n_{i}  \tag{5.114}\\
& \bar{\sigma}^{0}=Q \bar{e}^{0}+R \bar{\varepsilon}^{0} \tag{5.115}
\end{align*}
$$

where

$$
\begin{aligned}
\bar{e}_{i j}^{0} & =\frac{1}{2}\left(\bar{u}_{i, j}^{0}+\bar{u}_{j, i}^{0}\right) \\
\bar{e}^{0} & =\bar{u}_{i, i}^{0} \\
\bar{\varepsilon}^{0} & =\bar{U}_{i, i}^{0}
\end{aligned}
$$

Using (5.112) and (5.113) in (5.114) and (5.115) gives, after the necessary algebra,

$$
\begin{align*}
\bar{T}_{i}^{0}= & -i \frac{2 \mu a^{2}}{\omega}\left(\frac{Q+R}{H}\right)\left\{r_{, i} \frac{\partial r}{\partial n}\left[\frac{2}{a^{2} r^{2}}-K_{2}(a r)\right]\right. \\
& \left.+n_{i}\left[\frac{1}{2} K_{2}(a r)+\frac{1}{2} K_{0}(a r)-\frac{1}{a^{2} r^{2}}\right]\right\}  \tag{5.116}\\
\bar{\sigma}^{0}= & -f k K_{0}(a r) \tag{5.117}
\end{align*}
$$

It is desirable to express the fundamental solutions in terms of familiar material constants. We first recall that

$$
\begin{equation*}
k=\frac{\rho_{\rho} g f}{k_{0}} \tag{5.118}
\end{equation*}
$$

where $k_{0}$ is the soil permeability is $m s^{-1}, f$ is the soil porosity, $\rho_{f}$ is the density of the pore water and $g$ is the gravitational acceleration. Also, from (3.24) we write the constant $H$, defined in (5.52), as

$$
\begin{equation*}
H=\lambda+2 \mu+\frac{K_{f}}{f} \tag{5.119}
\end{equation*}
$$

where $\lambda$ and $\mu$ are the Lamé constant and shear modulus respectively of the soil skeleton, and $K_{f}$ is the bulk modulus of the pore water. The parameter $\lambda$ may be written in terms of the mo.. 'umiliar Poisson's ratio $\nu$ :

$$
\begin{equation*}
\lambda=\frac{2 \mu \nu}{1-2 \nu} \tag{5.120}
\end{equation*}
$$

and equation (5.119) may be re-written:

$$
\begin{equation*}
H=\frac{2 \mu(1-\nu)}{1-2 \nu}+\frac{K_{\rho}}{f} \tag{5.121}
\end{equation*}
$$

Using (3.24), (5.118), (5.120) and (5.121) we express the parameter $a^{2}$, defined in (5.81), as

$$
\begin{equation*}
a^{2}=\frac{-i \omega \rho_{f} g}{k_{0}}\left[\frac{1-2 \nu}{2 \mu(1-\nu)}+\frac{f}{K_{f}}\right] \tag{5.122}
\end{equation*}
$$

We can now summarize all the fundamental solutions as follows :

$$
\begin{align*}
& \bar{u}_{i}^{* k}=\left(\frac{H-\mu}{2 \mu H}\right) r_{, i} r_{, k}-\left(\frac{H+\mu}{2 \mu H}\right) \ln r \delta_{i k} \\
& -\frac{i k_{0} K_{j}^{2} a^{2}}{\omega f^{2} \rho_{f} g H^{2}}\left\{\left[\frac{2}{a^{2} r^{2}}-K_{2}(a r)\right] r_{, i} r_{, k}\right. \\
& \left.+\delta_{i k}\left[\frac{K_{1}(a r)}{a r}-\frac{1}{a^{2} r^{2}}\right]\right\}  \tag{5.123}\\
& \bar{U}_{i}^{* k}=\left(\frac{H-\mu}{2 \mu H}\right) r_{, i} r_{, k}-\left(\frac{H+\mu}{2 \mu H}\right) \ln r \delta_{i k} \\
& +\frac{i\left(H-K_{f}\right) K_{f} k_{0} a^{2}}{\omega f^{2} \rho_{f} g H^{2}}\left\{\left[\frac{2}{a^{2} r^{2}}-K_{2}(a r)\right] r_{, i} r_{, k}\right. \\
& \left.+\delta_{i k}\left[\frac{K_{1}(a r)}{a r}-\frac{1}{a^{2} r^{2}}\right]\right\}  \tag{5.124}\\
& \bar{T}_{i}^{* k}=-2\left(\frac{H-\mu}{H}\right) \frac{r_{i} r_{, k}}{r} \frac{\partial r}{\partial n}+\frac{\mu}{H} \frac{1}{r}\left(r_{, k} n_{i}-r_{, i} n_{k}-\frac{\partial r}{\partial n} \delta_{i k}\right) \\
& -\frac{i 2 \mu K_{f}^{2} k_{0} a^{3}}{\omega f^{2} \rho_{j} g H^{2}}\left\{\frac{\partial r}{\partial n} \delta_{i k}\left[\frac{2}{a^{3} r^{3}}-\frac{K_{2}(a r)}{a r}\right]\right. \\
& +\frac{\partial r}{\partial n} r_{i,} r_{, k}\left[K_{3}(a r)-\frac{8}{a^{3} r^{3}}\right] \\
& +r_{, k} n_{i}\left[\frac{2}{a^{3} r^{3}}+\frac{3 K_{2}(a r)}{a r}-K_{3}(a r)\right] \\
& \left.+r_{, i} n_{k}\left[\frac{2}{a^{3} r^{3}}-\frac{K_{2}(a r)}{a r}\right]\right\}  \tag{5.125}\\
& \bar{\sigma}^{* k}=\left(\frac{a K_{f}}{H}\right) r_{, k}\left[K_{1}(a r)-\frac{1}{a r}\right]  \tag{5.126}\\
& \bar{u}_{i}^{0}=-\left(\frac{i K_{f} a}{\omega H}\right) r_{, i}\left[K_{1}(a r)-\frac{1}{a r}\right]  \tag{5.127}\\
& \bar{U}_{i}^{0}=\frac{i}{\omega}\left(1-\frac{K_{f}}{H}\right) a r_{, i}\left[K_{1}(a r)-\frac{1}{a r},\right.  \tag{5.128}\\
& \bar{T}_{i}^{0}=-i \frac{2 \mu K_{f} a^{2}}{\omega H}\left\{r_{, i} \frac{\partial r}{\partial n}\left[\frac{2}{a^{2} r^{2}}-K_{2}(a r)\right]\right. \\
& \left.+n_{i}\left[\frac{1}{2} K_{2}(a r)+\frac{1}{2} K_{0}(a r)-\frac{1}{a^{2} r^{2}}\right]\right\}  \tag{5.129}\\
& \bar{\sigma}^{0}=\frac{-\rho_{\rho} g f^{2}}{k_{0}} K_{0}(a r) \tag{5.130}
\end{align*}
$$

where the constants $H$ and $a$ are given by (5.121) and (5.122) respectively, and
suffixes $i$ and $k$ take the values 1 and 2.

### 5.3 Evaluation of $\alpha_{i k}$ and $\theta$

As previously mentioned, the coefficients $\alpha_{i k}$ and $\theta$ appearing in the boundary integral equations (5.45) and (5.46) may be determined from the evaluation of the Cauchy principal values of the integrals. However,as will be seen below, it is not necessary to know these quantities explicitly. They will be determined by a method known as the " rigid body motion" analogy, which is frequently used in the boundary element literature. First we rewrite the boundary integral equations (5.45) and ( 5.46 ) so that pore pressure $p$ appears explicitly. To do this we substitute $\bar{\sigma}=-f \bar{p}$ and define the fundamental solutions $\overrightarrow{\boldsymbol{p}}^{\boldsymbol{k}}$ and $\vec{p}^{0}$ by

$$
\begin{gathered}
\bar{\sigma}^{* k}=-f \vec{p}^{* k} \\
\bar{\sigma}^{o}=-f \bar{p}^{0}
\end{gathered}
$$

Substituting these in (5.45) and (5.46), multiplying (5.46) by $\frac{-\omega}{i f}$ and using (5.97) gives the boundary integral equations in the form

$$
\begin{array}{r}
2 \pi \alpha_{i k} \bar{u}_{i}(P)=P V \int_{\Gamma}\left(\bar{T}_{i} \bar{u}_{i}^{* k}-\bar{T}_{i}^{* k} \bar{u}_{i}\right)-\left(\bar{p} \bar{v}_{n}^{* k}-\bar{p}^{* k} \bar{v}_{n}\right) d \Gamma \\
i, k=1,2 \\
2 \pi \theta \bar{p}(P)=P V \int_{\Gamma}\left(\bar{T}_{i} \bar{u}_{i}^{01}-\bar{T}_{i}^{01} \bar{u}_{i}\right)-\left(\bar{p} \bar{v}_{n}^{01}-\bar{p}^{01} \bar{v}_{n}\right) d \Gamma  \tag{5.132}\\
i=1,2
\end{array}
$$

where we have written

$$
\begin{align*}
\bar{v}_{n} & =f\left(\bar{U}_{j}-\bar{u}_{j}\right) n_{j} \\
\bar{v}_{n}^{* k} & =f\left(\bar{U}_{j}^{* k}-\bar{u}_{j}^{* k}\right) n_{j} \\
\bar{u}_{i}^{01} & =\frac{-\omega}{i f} \bar{u}_{i}^{0} \\
\bar{U}_{i}^{01} & =\frac{-\omega}{i f} \bar{U}_{i}^{0}  \tag{5.133}\\
\bar{T}_{i}^{01} & =\frac{-\omega}{i f} \bar{T}_{i}^{0} \\
\bar{p}^{01} & =\frac{-\omega}{i f} \bar{p}^{0}=\frac{\omega \bar{\sigma}^{0}}{i f^{2}} \\
\bar{v}_{n}^{01} & =f\left(\bar{U}_{i}^{01}-\bar{u}_{i}^{01}\right) n_{i}-\frac{1}{r} \frac{\partial r}{\partial n}
\end{align*}
$$

The position vector of the point $P$ on the boundary (at which $r$ is zero) is denoted by $P$. To determine the coefficient $\alpha_{i k}$ we notice that the governing equations ( 5.47 ) and ( 5.48 ) admit the following trivial constant solutions for the case of zero body forces:
(a) $\bar{u}_{1}=\bar{U}_{1}=1$ and $\bar{u}_{2}=\bar{U}_{2}=0 \Rightarrow \bar{v}_{n}=0$
(b) $\bar{u}_{1}=\bar{U}_{1}=0$ and $\bar{u}_{2}=\bar{U}_{2}=1 \Rightarrow \bar{v}_{n}=0$
where it is implicit that $\boldsymbol{\omega}=0$. No stresses or pore pressures are induced by these cases, as is easily verified from the constitutive relations, i.e. $\bar{T}_{i}=0$ and $\bar{p}=0$. Substituting the first case (a) into ( 5.131 ) gives :

$$
2 \pi \alpha_{1 k}=-\int_{\Gamma} \tilde{T}_{1}^{* k} d \Gamma
$$

Similarly, the second case (b) produces from ( 5.131)

$$
2 \pi \alpha_{2 k}=-\int_{\Gamma} \bar{T}_{2}^{* k} d \Gamma
$$

Thus we have

$$
\begin{equation*}
2 \pi \sigma_{1, k}=-\int_{\Gamma} \vec{T}_{i}^{* k} d \Gamma \quad i, k=1,2 \tag{5.134}
\end{equation*}
$$

To determine the parameter $\theta$ we consider (again for no body forces) the case of unit pore pressure and zero displacement of the soil skeleton, i.e., $\bar{p}=1, \bar{u}_{i}=0$ and $\omega=0$. From the constitutive laws (3.22) and (3.23) we have (using $\bar{\sigma}=-f \bar{p}$ )

$$
\bar{\varepsilon}=-\frac{1}{K_{j}} \text { and } \bar{\sigma}_{i j}=-(1-f) \delta_{i j}
$$

Thus for a compressible fluid ( $K_{f} \neq \infty$ ), the fluid displacement $\bar{U}_{i}$ is non-zero. We note also ,from (5.8), that the tractions $\bar{T}_{i}$ are non-zero. Since $\omega=0$, the parameter $a$ defined in (5.122) is zero. Hence, in the boundary integral equations we must use the limit of the fundamental solutions as $a \rightarrow 0$. We now examine the fundamu . $\operatorname{tal}$ solutions appearing in (5.132). It will be shown in section (6.2) that the funci:nns $\bar{u}_{i}^{01}, \bar{T}_{i}^{01}, \bar{p}^{01}$ possess logarithmic singularity at ar $=0$,(See 6.29). From the definitions (5.133) and (5.127), (5.128), (5.130),it is clear that $\bar{u}_{i}^{01}, \bar{T}_{i}^{01}, \bar{p}^{01}$ all approach zero as $a \rightarrow 0$. The function $\bar{v}_{n}^{01}$,on the other hand, behaves differently. From (5.133), (5.127) and (5.128) we have

$$
\bar{v}_{n}^{01}=-a r_{, i}\left[K_{1}(a r)-1 / a r\right] n_{i}-\frac{1}{r} \frac{\partial r}{\partial n}
$$

The first term tends to zero as $a \rightarrow 0$, as can be seen from ( 6.22 ). Thus putting $\bar{p}=1, \bar{u}_{i}=0$ in (5.132) and taking limits of the fundamental solutions as $a \rightarrow 0$ we have

$$
\begin{equation*}
2 \pi \theta=\int_{\Gamma} \frac{1}{r} \frac{\partial r}{\partial n} d \Gamma \tag{5.135}
\end{equation*}
$$

## Chapter 6

## Boundary Element Procedure

In this chapter we outline the numerical details of the boundary element method used for solving a boundary value problem. The accuracy of the method is illustrated by comparison with the analytic solution for the wave loading of a horizontal isotropic poroelastic seabed which is briefly described in Appendix I

### 6.1 Discretisation of Boundary Integral Equations

Using (5.134) and (5.135), the boundary integral equations (5.45) and (5.46) become

$$
\begin{equation*}
\int_{\Gamma}\left\{\tilde{\bar{T}}_{i} \tilde{\bar{u}}_{i}^{* k}+\tilde{\bar{T}}_{i}^{\Delta k}\left[\tilde{\bar{u}}_{i}(P)-\tilde{\bar{u}}_{i}\right]-\tilde{\bar{p}}_{\tilde{\bar{v}}_{n}^{* k}}+\tilde{\bar{p}}^{* k} \tilde{\bar{v}}_{n}\right\} d \Gamma=0 \tag{6.1}
\end{equation*}
$$

$$
\begin{equation*}
\int_{\Gamma}\left\{\tilde{\bar{T}}_{i} \tilde{\bar{u}}_{i}^{01}-\tilde{\bar{T}}_{i}^{02} \tilde{\bar{u}}_{i}-\tilde{\bar{p}} f\left(\tilde{\bar{U}}_{i}^{01}-\tilde{\bar{u}}_{i}^{01}\right) n_{i}+\tilde{\bar{p}}^{01} \tilde{\bar{v}}_{n}+\frac{r_{0}}{r} \frac{\partial r}{\partial n}[\tilde{\bar{p}}-\tilde{\bar{p}}(P)]\right\} d \Gamma=0 \tag{6.2}
\end{equation*}
$$

where ( 5.46 ) was multiplied by $\frac{r_{0}}{\mu}$ ( $r_{0}$ is an arbitrary length) and we have defined the following dimensionless quantities :

$$
\begin{array}{cc}
\tilde{\bar{u}}_{i}^{* k}=\mu \bar{u}_{i}^{* k} & \tilde{\bar{v}}_{n}^{* k}=\mu \bar{v}_{n}^{* k} \\
\tilde{\bar{T}}_{i}^{* k}=r_{0} \bar{T}_{i}^{k} & \tilde{\bar{p}}^{* k}=r_{0} \bar{p}^{* k} \\
\tilde{\bar{u}}_{i}^{01}=r_{0} \bar{u}_{i}^{01} & \tilde{\bar{U}}_{i}^{01}=r_{0} \bar{u}_{i}^{01}  \tag{6.3}\\
\tilde{T}_{i}^{01}=\frac{r_{0}^{2}}{\mu} \bar{T}_{i}^{01} & \tilde{\bar{p}}^{01}=\frac{r_{0}^{2}}{\mu} \bar{p}^{01}
\end{array}
$$

It is understood that the integrals are evaluated in the Cauchy principal value sense and we recall that $P$ denotes the position vector of the point at which $r=0$. We now discretise the boundary $\Gamma$ into $M$ straight line segments $\Gamma_{e}, e=1$ to $M$. Each segment is termed a "boundary element" and a node is located at its midpoint. On each element we make the assumption

$$
\tilde{\bar{T}}_{i}=\tilde{\bar{T}}_{i}, \quad \tilde{\bar{u}}_{i}=\tilde{\bar{u}}_{i}^{e}
$$

where $\tilde{\bar{T}}_{i}^{e}$ and $\tilde{\bar{u}}_{i}^{e}$ denote the values of $\tilde{\bar{T}}_{i}$ and $\tilde{\bar{u}}_{i}$ at node $e$. For this reason the elements are termed "constant" elements.

If we let point $\mathbf{P}$ be the $j$ th node (i.e. $r=0$ at node $j$ ) then equation (6.1) becomes

$$
\begin{align*}
& \sum_{e=1}^{M} \tilde{\bar{T}}_{i}^{e} \int_{\Gamma_{e}} \tilde{\tilde{u}}_{i}^{* k} d \Gamma_{e}+\sum_{e=1}^{M}\left(\tilde{\bar{u}}_{i}^{j}-\tilde{\bar{u}}_{i}^{e}\right) \int_{\Gamma_{e}} \tilde{\bar{T}}_{i}^{* k} d \Gamma_{e}  \tag{6.4}\\
& \quad+\sum_{e=1}^{M}-\tilde{\tilde{p}}^{e} \int_{\Gamma_{e}} \tilde{\tilde{v}}_{n}^{* k} d \Gamma_{e}+\sum_{e=1}^{M} \tilde{\bar{v}}_{n}^{e} \int_{\Gamma_{e}} \tilde{\bar{p}}^{* k} d \Gamma_{e}=0
\end{align*}
$$

The second term on the left hand side is

$$
\begin{aligned}
\sum_{e=1}^{M}\left(\tilde{\tilde{u}}_{i}^{j}-\tilde{\bar{u}}_{i}^{e}\right) \int_{\Gamma_{e}} \tilde{\bar{T}}_{i}^{* k} d \Gamma_{e} & =\sum_{\substack{e=1 \\
e \neq j}}^{M}\left(\tilde{\bar{u}}_{i}^{j}-\tilde{\bar{u}}_{i}^{e}\right) \int_{\Gamma_{e}} \tilde{\bar{T}}_{i}^{* k} d \Gamma_{e} \\
& =\tilde{\bar{u}}_{i}^{j} \sum_{\substack{i=1 \\
\overline{\neq j}}}^{M} \int_{\Gamma_{e}} \tilde{\bar{T}}_{i}^{* k} d \Gamma_{e}-\sum_{\substack{e=1 \\
e \neq j}}^{M} \tilde{\bar{u}}_{i}^{e} \int_{\Gamma_{e}} \tilde{\bar{T}}_{i}^{* k} d \Gamma_{e} \\
& =\sum_{e=1}^{M} S_{k i}^{e} \tilde{\bar{u}}_{i}^{e}
\end{aligned}
$$

where we have defined $S_{k i}^{e}$ as

$$
\begin{gathered}
S_{k i}^{e}= \begin{cases}-\int_{\Gamma_{i}} \tilde{\bar{T}}_{i}^{* k} d \Gamma_{e} & e \neq j \\
\sum_{\substack{r=1 \\
r \neq j}}^{M} \int_{\Gamma_{r}} \tilde{\bar{T}}_{i}^{* k} d \Gamma_{r} & e=j\end{cases} \\
i, k=1,2
\end{gathered}
$$

We now note that (6.1) may be written in the form

$$
\begin{align*}
& \sum_{e=1}^{M}\left(R_{k i}^{e} \tilde{\tilde{T}}_{i}^{e}\right.  \tag{6.5}\\
&\left.i=S_{k i}^{e} \tilde{u}_{i}^{e}\right)=0 \\
& i=1 \text { to } 3, k=1,2
\end{align*}
$$

where we have defined $\tilde{\bar{u}}_{3}=\tilde{\bar{v}}_{n}$ and

$$
\begin{gathered}
S_{k 3}^{e}=\int_{\Gamma_{\mathrm{e}}} \tilde{\bar{p}}^{* k} d \Gamma_{e} \\
R_{k i}^{e}= \begin{cases}\int_{\Gamma_{\mathrm{e}}} \tilde{\bar{u}}_{i}^{* k} d \Gamma_{e} & i=1,2 \\
\int_{\Gamma_{e}} \tilde{\bar{v}}_{n}^{* k} d \Gamma_{\mathrm{e}} & i=3\end{cases} \\
k=1,2
\end{gathered}
$$

Writing $\tilde{\bar{T}}_{3}=-\tilde{\bar{p}}$ and discretising ( 6.2 ), we obtain

$$
\begin{align*}
& \sum_{e=1}^{M} \tilde{T}_{i}^{e} \int_{\Gamma_{e}} \tilde{\bar{u}}_{i}^{01} d \Gamma_{e}+\sum_{e=1}^{M} \tilde{\bar{u}}_{i}^{e} \int_{\Gamma_{e}}-\bar{T}_{i}^{01} d \Gamma_{e} \\
& \top \sum_{e=1}^{M} \tilde{\bar{T}}_{3}^{e} \int_{\Gamma_{e}} f\left(\tilde{U}_{i}^{01}-\tilde{u}_{i}^{01}\right) n_{i} d \Gamma_{e}  \tag{6.6}\\
& +\sum_{e=1}^{M} \tilde{\bar{u}}_{3}^{e} \int_{\Gamma_{e}} \tilde{\bar{p}}^{01} d \Gamma_{e}+\sum_{e=1}^{M}\left(\tilde{T}_{3}^{e}-\tilde{\bar{T}}_{3}^{j}\right) \int_{\Gamma_{e}}-\frac{r_{0}}{r} \frac{\partial r}{\partial n} d \Gamma_{e}=0
\end{align*}
$$

Proceeding as done earlier we write the last term on the left hend side as

$$
\begin{aligned}
\sum_{e=1}^{M}\left(\tilde{\bar{T}}_{3}^{e}-\tilde{\bar{T}}_{3}^{j}\right) \int_{\Gamma_{e}}-\frac{r_{0}}{r} \frac{\partial r}{\partial n} d \Gamma_{e} & =\sum_{\substack{a=1 \\
e \neq j}}^{M}\left(\tilde{\bar{T}}_{3}^{e}-\tilde{\bar{T}}_{3}^{j}\right) \int_{\Gamma_{e}}-\frac{r_{0}}{r} \frac{\partial r}{\partial n} d \Gamma_{e} \\
& =\sum_{\substack{i=1 \\
e \neq j}}^{M} \tilde{T}_{3} \int_{\Gamma_{e}}-\frac{r_{0}}{r} \frac{\partial r}{\partial n} d \Gamma_{e}+\tilde{\bar{T}}_{3}^{j} \sum_{\substack{e=1 \\
e \neq j}}^{M} \int_{\Gamma_{e}} \frac{r_{0}}{r} \frac{\partial r}{\partial n} d \Gamma_{e} \\
& =\sum_{\substack{e=1}}^{M} \tilde{T}_{3}^{e} \tilde{v}^{2}
\end{aligned}
$$

where $V^{e}$ is defined as

$$
V^{e}= \begin{cases}\int_{\Gamma_{e}}-\frac{r_{0}}{r} \frac{\partial r}{\partial n} d \Gamma_{e} & e \neq j  \tag{6.7}\\ \sum_{\substack{j=1 \\ j \neq j}}^{M} \int_{\Gamma_{0}} \frac{r_{0}}{r} \frac{\partial r}{\partial n} d \Gamma_{e} & e=j\end{cases}
$$

Equation (6.6) may thus be written in the form

$$
\begin{gather*}
\sum_{e=1}^{M} R_{3 i}^{e} \tilde{\bar{T}}_{i}^{e}+\sum_{e=1}^{M} S_{3 i}^{e} \tilde{u}_{i}^{e}=0  \tag{6.8}\\
i=1 \text { io } 3
\end{gather*}
$$

where

$$
\begin{gathered}
R_{3 i}^{e}= \begin{cases}\int_{\Gamma_{e}}^{\tilde{\bar{u}}_{i}^{01} d \Gamma_{e}} & i=1,2 \\
\int_{\Gamma_{e}} f\left(\tilde{\bar{U}}_{i}^{01}-\tilde{\bar{u}}_{i}^{01}\right) n_{i} d \Gamma_{e}+V^{e} & i=3\end{cases} \\
S_{3 i}^{e}= \begin{cases}\int_{\Gamma_{e}}-\tilde{\bar{T}}_{i}^{01} d \Gamma_{e} & i=1,2 \\
\int_{\Gamma_{e}} \tilde{\bar{p}}^{01} d \Gamma_{e} & i=3\end{cases}
\end{gathered}
$$

Combining (6.5) and (6.8) we have

$$
\begin{gather*}
\sum_{e=1}^{M}\left(R_{k i}^{e} \frac{\tilde{T}_{i}^{e}}{T_{i}}+S_{k i}^{e} \tilde{u_{i}^{e}}\right)=0  \tag{6.9}\\
i, k=1 \text { to } 3
\end{gather*}
$$

where we recall that $\tilde{\bar{u}}_{3}=\tilde{\bar{v}}_{n}$ and $\tilde{\bar{T}}_{3}=-\tilde{\bar{p}}$. Equation (6.9) is in fact the discretised boundary integral equation. The formulae for $R_{k i}^{e}$ and $S_{k i}^{e}$ may be written more compactly if we make the following definitions:

$$
\begin{align*}
& \tilde{\bar{u}}_{i}^{* 3}=\tilde{\tilde{u}}_{i}^{01}, \quad \tilde{\bar{T}}_{i}^{* 3}=\tilde{\bar{T}}_{i}^{01},  \tag{6.10}\\
& \tilde{\bar{u}}_{3}^{* k}= \begin{cases}\tilde{\bar{v}}_{n}^{* k} & k=1,2 \\
f\left(\tilde{\bar{U}}_{i}^{01}-\tilde{\bar{u}}_{i}^{01}\right) n_{i} & k=3\end{cases}  \tag{6.11}\\
& \tilde{\bar{T}}_{3}^{* k}= \begin{cases}-\tilde{\bar{p}}^{* k} & k=1,2 \\
-\overline{\bar{p}}^{01} & k=3\end{cases} \tag{6.12}
\end{align*}
$$

Then

$$
\begin{gather*}
R_{k i}^{e}= \begin{cases}\int_{\Gamma_{e}} \tilde{\tilde{u}}_{3}^{* 3} d \Gamma_{e}+V^{e} & i=k=3 \\
\int_{\Gamma_{e}} \tilde{\bar{u}}_{i}^{* k} d \Gamma_{e} & \text { otherwise }\end{cases}  \tag{6.13}\\
S_{k i}^{e}= \begin{cases}\sum_{\substack{=1 \\
a \neq j}}^{M} \int_{G a m m a} \tilde{\bar{T}}_{i}^{* k} d \Gamma_{s} & i, k=1,2 \\
-\int_{\Gamma_{e}} \tilde{\bar{T}}_{i}^{* k} d \Gamma_{e} & e=j\end{cases}  \tag{6.14}\\
\text { itherwise }
\end{gather*}
$$

### 6.2 Function Approximations

To evaluate integrals $R_{k i}^{e}, S_{k i}^{e}$ we need to take a closer look at the integrands. It will be noted that the fundamental solutions are functions of $a r, r_{, i}$ and $n_{i}$, these quantities being defined in Chapter 5. From (5.81) we note that $a^{2}$ may be written in the form

$$
\begin{equation*}
a^{2}=-i c^{2} \tag{6.15}
\end{equation*}
$$

where

$$
\begin{equation*}
c=\frac{\omega f k H}{R(\lambda+2 \mu)} \tag{6.16}
\end{equation*}
$$

The constant $c$ is real and positive and has dimensions of (length) ${ }^{-1}$. From (6.15) we have

$$
\begin{equation*}
a r=c r e^{-i \frac{\pi}{4}} \tag{6.17}
\end{equation*}
$$

The Bessel functions of argument ar may now be separated into real and imaginary parts via the relations (Abramowitz and Stegun, 1970).

$$
\begin{equation*}
K_{0}\left(x e^{-i \frac{\pi}{4}}\right)=k e r x-i k e i x \tag{6.18}
\end{equation*}
$$

$$
\begin{align*}
K_{1}\left(x e^{-i \frac{\pi}{4}}\right) & =-\left(k e i_{1} x+i k e r_{1} x\right) \\
& =-\frac{1}{\sqrt{2}}\left(\text { ker }^{\prime} x+k e i^{\prime} x\right)-\frac{i}{\sqrt{2}}\left(\text { ker }^{\prime} x-k e i^{\prime} x\right) \tag{6.19}
\end{align*}
$$

and

$$
\begin{array}{r}
K_{n+1}(z)=K_{n-1}(z)+\frac{2 n}{z} K_{n}(z)  \tag{6.20}\\
\quad n \text { integer, } z \text { real or complex }
\end{array}
$$

To evaluate the Kelvin functions $k e r, k e i, k e r^{\prime}, k e i^{\prime}$ we use the polynomial approximations given by Abramowitz and Stegun, Sec. 9.11. These have been tested against IMSL routines for the Kelvin functions and have been found to be extremely accurate. The use of the IMSL routines directly in the BEM program is not computationally efficient owing to the large number of function evaluations required. Using the polynomial representations, therefore, we write expressions for the following functional groups involving the Bessel functions, which appear in the fundamental solutions. The algebra involved is quite extensive and is not recorded. The final results are

$$
\begin{array}{r}
K_{0}(a r)=\ln \left(\frac{c r}{2}\right) \tilde{u}_{0}(c r)+\tilde{v}_{0}(c r) \\
\frac{K_{1}(a r)}{a r}-\frac{1}{a^{2} r^{2}}=\ln \left(\frac{c r}{2}\right) \tilde{u}_{1}(c r)+\tilde{v}_{1}(c r) \\
K_{2}(a r)-\frac{2}{a^{2} r^{2}}=\ln \left(\frac{c r}{2}\right) \tilde{u}_{2}(c r)+\tilde{v}_{2}(c r) \\
\frac{K_{2}(a r)}{a r}-\frac{2}{a^{3} r^{3}}=\ln \left(\frac{c r}{2}\right) \tilde{u}_{3}(c r)+\tilde{w}_{3}(c r)+\frac{1}{c r} \tilde{z}_{3}(c r) \\
K_{3}(a r)-\frac{3 K_{2}(a r)}{a r}-\frac{2}{a^{3} r^{3}}=\ln \left(\frac{c r}{2}\right) \tilde{u}_{4}(c r)+\tilde{w}_{4}(c r)+\frac{1}{c r} \tilde{z}_{4}(c r) \\
\frac{1}{2} K_{2}(a r)+\frac{1}{2} K_{0}(a r)-\frac{1}{a^{2} r^{2}}=\ln \left(\frac{c r}{2}\right) \tilde{u}_{5}(c r)+\tilde{v}_{5}(c r) \\
K_{3}(a r)-\frac{8}{a^{3} r^{3}}=\ln \left(\frac{c r}{2}\right) \tilde{u}_{6}(c r)+\tilde{w}_{6}(c r)+\frac{1}{c r} \tilde{z}_{8}(c r) \tag{6.27}
\end{array}
$$

where for any $n$

$$
\begin{array}{r}
\tilde{u}_{n}(x)=u_{n}(x)+i u_{n}^{\prime}(x) \\
\tilde{v}_{n}(x)=v_{n}(x)+i v_{n}^{\prime}(x) \\
\tilde{w}_{n}(x)=w_{n}(x)+i w_{n}^{\prime}(x)  \tag{6.28}\\
\tilde{z}_{n}(x)=z_{n}(x)+i z_{n}^{\prime}(x)
\end{array}
$$

The functions $u_{n}, u_{n}^{\prime}, v_{n}, v_{n}^{\prime}, w_{n}, w_{n}^{\prime}, z_{n}, z_{n}^{\prime}$ are recorded in Appendix $\mathbf{F}$. These expressions enable us to express the fundamental solutions $\tilde{\bar{u}}_{i}^{* k}$ and $\tilde{\bar{T}}_{i}^{* k}$ as

$$
\begin{align*}
\tilde{\bar{u}}_{i}^{* k} & =\ln \left(\frac{c r}{2}\right) \tilde{H}_{i k}+\tilde{G}_{i k} \\
\tilde{\bar{T}}_{i}^{* k} & =\ln \left(\frac{c r}{2}\right) \tilde{P}_{i k}+\tilde{Q}_{i k}+\frac{1}{c r} \tilde{Z}_{i k}(c r) \tag{6.29}
\end{align*}
$$

where

$$
\begin{align*}
\tilde{H}_{i k} & =H_{i k}+i H_{i k}^{\prime} \quad \tilde{G}_{i k}=G_{i k}+i G_{i k}^{\prime} \\
\tilde{P}_{i k} & =P_{i k}+i P_{i k}^{\prime} \quad \tilde{Q}_{i k}=Q_{i k}+i Q_{i k}^{\prime}  \tag{6.30}\\
\tilde{Z}_{i k} & =Z_{i k}+i Z_{i k}^{\prime}
\end{align*}
$$

These functions (6.30) are defined in Appendix G.

### 6.3 Evaluation of Integrals

As can be seen from Appendix $F$ and Appendix $G$ the functions $\tilde{H}_{i}^{* k}, \tilde{G}_{i}^{+k}, \tilde{P}_{i}^{* k}$, $\tilde{Q}_{i}{ }^{\kappa k}, \tilde{Z}_{i}^{* k}$, defined in (6.30), possess no singularities. We see from (6.7) and (6.29)
that the integrands of $\boldsymbol{R}_{k i}^{e}$ and $S_{k i}^{e}$ (equations 6.13 and 6.14) possess terms of order $\frac{1}{r}$ for certain values of $i$ and $k$. It will be noted,however, that it is not neccessary to integrate these expressions over the element on which $r=0$ (element $j$ ). Hence all integrals may be considered of the form

$$
\begin{equation*}
I=\int_{\Gamma_{e}}\left[\ln \left(\frac{c r}{2}\right) f\left(\frac{r}{r_{0}} ; r_{, i} ; n_{i}\right)+g\left(\frac{r}{r_{0}} ; r_{, i} ; n_{i}\right)\right] d \Gamma_{e} \tag{6.31}
\end{equation*}
$$

where the functions $f$ and $g$ are non-singular at $r=0$ and $\Gamma_{e}$ is an arbitrary element $e$. Special procedures are required for the case $e=j$, owing to the presence of the logarithmic singularity on this elemer.t. In addition, we make note of the following facts concerning the nature of functions $f$ and $g$ in (6.31). The functional approximations used (Appendix F) change form at the value $\mathrm{cr}=8$. When writing an integral in the form (6.31) we have employed the following programming device. For the range $c r>8$ the function $f$ has been set to zero and the entire functional form of the integrand has been assigned to the function $g$. The result is that both $f$ and $g$ are discontinuous at $c r=8$, while the entire form of the integrand in ( 6.31 ) is not. This must be taken into consideration only on element $j$, since it is only on this element that it is necessary to separate the integrands into logarithmic and non-logarithmic parts.

We now describe the details of the integration procedure. We consider that the boundary $\Gamma$ lies in the $x-z$ plane with the $x$ and $z$ coordinate axes directed to the right and vertically downward respectively. We recall that the problem is two dimensional, i.e. conditions are uniform in the $y$ direction. We denote the coordinates of the end points of an arbitrary element $\Gamma_{e}$ by $\left(x_{e}, z_{e}\right)$ and $\left(x_{e+1}, z_{e+1}\right)$ and define the integration direction from the former point to the latter. Since there are $M$ elements, $\left(x_{M+1}, z_{M+1}\right) \equiv\left(x_{1}, z_{1}\right)$. The $x$ and $z$ coordinates of a point on $\Gamma_{e}$
are represented parametrically as

$$
\begin{array}{r}
x=\frac{1}{2}(1-\xi) x_{e}+\frac{1}{2}(1+\xi) x_{e+1} \\
z=\frac{1}{2}(1-\xi) z_{e}+\frac{1}{2}(1+\xi) z_{e+1} \\
\text { where }-1 \leq \xi \leq 1
\end{array}
$$

The point $p$ from which $r$ is measured is located at node $j$ (mid-point of element $\left.\Gamma_{j}\right)$. Denoting the coordinates of $p$ by $\left(x_{0}, z_{0}\right)$ we have

$$
\begin{equation*}
x_{0}=\frac{1}{2}\left(x_{j}+x_{j+1}\right) \quad z_{0}=\frac{1}{2}\left(z_{j}+z_{j+1}\right) \tag{6.32}
\end{equation*}
$$

Since $r=|\underline{x}-p|$ we have

$$
\begin{array}{r}
r=\left[\left(x-x_{0}\right)^{2}+\left(z-z_{0}\right)^{2}\right]^{\frac{1}{2}} \\
r_{, 1} \equiv \frac{\partial r}{\partial x}=\frac{x-x_{0}}{r}  \tag{6.33}\\
r_{, 2} \equiv \frac{\partial r}{\partial z}=\frac{z-z_{0}}{r}
\end{array}
$$

It is easily verified that the unit outward normal $n_{i}$ for anticlockwise integration around $\Gamma$ is given by

$$
\begin{equation*}
n_{1}=\frac{1}{2 J_{e}}\left(z_{e}-z_{e+1}\right) \quad ; \quad n_{3}=\frac{1}{2 J_{e}}\left(x_{c+1}-x_{e}\right) \tag{6.34}
\end{equation*}
$$

where

$$
\begin{equation*}
J_{e}=\frac{1}{2}\left[\left(x_{e+1}-x_{e}\right)^{2}+\left(z_{e+1}-z_{e}\right)^{2}\right]^{\frac{1}{2}} \tag{6.35}
\end{equation*}
$$

The line element $d \Gamma_{c}$ is given by

$$
\begin{equation*}
d \Gamma_{e}=J_{e} d \xi \tag{6.36}
\end{equation*}
$$

For the case $e \neq j$ it is thus quite easy to express (6.31) in the form

$$
\begin{equation*}
I=\int_{-1}^{1} F(\xi) d \xi \tag{6.37}
\end{equation*}
$$

by using ( 6.32 ) through ( 6.36 ). Standard Gaussian quadrature can now be applied to ( 6.37 ). Integration along element $\Gamma_{j}$ must be considered in two cases, viz. integration before and after the mid-point $\left(x_{0}, z_{0}\right)$, at which the logarithmic singularity occurs. We shall denote these two integrals by $I_{a}$ and $I_{b}$ respectively.
Case(a): Integration from $\left(x_{j}, z_{j}\right)$ to $\left(x_{0}, z_{0}\right)$
We define $r_{\text {max }}$ to be the distance from $\left(x_{j}, z_{j}\right)$ to $\left(x_{0}, z_{0}\right)$, i.e.

$$
\begin{equation*}
r_{\text {max }}=\frac{1}{2}\left[\left(x_{j}-x_{0}\right)^{2}+\left(z_{j}-z_{0}\right)^{2}\right]^{\frac{1}{2}} \tag{6.38}
\end{equation*}
$$

We find by putting $\left(x_{e}, z_{e}\right) \equiv\left(x_{j}, z_{j}\right)$ and $\left(x_{e+1}, z_{e+1}\right) \equiv\left(x_{0}, z_{0}\right)$ in (6.32) to (6.36) that

$$
\begin{array}{ll}
r=\frac{1}{2}(1-\xi) r_{\max } & \\
r_{, 1}=\frac{x_{j}-x_{0}}{r_{\max }} & r_{, 2}=\frac{z_{j}-z_{0}}{r_{\max }}  \tag{0.39}\\
n_{1}=r_{, 2} & n_{2}=-r_{, 1} \\
\frac{\partial r}{\partial n} \equiv r_{, i} n_{i}=0 &
\end{array}
$$

Hence we have

$$
\begin{align*}
I_{a}= & \frac{1}{2} r_{\max } \int_{-1}^{1}\left\{\ln \left[\frac{c(1-\xi) r_{\max }}{\dot{q}}\right] f\left[\frac{(1-\xi) r_{\max }}{2 r_{0}} ; r_{, i} ; n_{i}\right]\right. \\
& \left.+g\left[\frac{(1-\xi) r_{\max }}{2 r_{0}} ; r_{, i} ; n_{i}\right]\right\} d \xi \tag{0.40}
\end{align*}
$$

We note that $r_{, i}$ and $n_{i}$ are constant on $\Gamma_{j}$. Because of the discontinuities in the functions $f$ and $g$ at $c r=8$, the integral $I_{a}$ must be evaluated in two cases also:
(1) $c r_{\text {max }}<8$

In this case the discontinuity does not lie in the range of the integration. By isolating the logarithmic singularity (at $\xi=1$ ) and making appropriate changes of
variable we can express (6.40) as

$$
\begin{align*}
I_{a}= & \frac{1}{2} r_{\max } \int_{-1}^{1} \ln \left[\frac{c r_{\max }}{4}\right] f\left[\frac{(1-\xi) r_{\max }}{2 r_{0}} ; r_{, i} ; n_{i}\right]+g\left[\frac{(1-\xi) r_{\max }}{2 r_{0}} ; r_{, i} ; n_{i}\right] d \xi \\
& +\frac{1}{4} r_{\max } \int_{-1}^{1} \ln \left(\frac{3-\xi}{2}\right) f\left[\frac{(3-\xi) r_{\max }}{4 r_{0}} ; r_{, i} ; n_{i}\right] d \xi \\
& -\frac{1}{2} r_{\max } \int_{0}^{1} \ln \left(\frac{1}{\xi}\right) f\left[\frac{\xi r_{\max }}{2 r_{0}} ; r_{, i} ; n_{i}\right] d \xi \tag{6.41}
\end{align*}
$$

(2) $c r_{\text {max }}>8$

The discontinuity at $c r=8$ now lies within the range of integration. We define the value $\xi_{0}$ such that $\mathrm{cr}=8$ at $\xi=\xi_{0}$. From ( 6.39 ) we thus have

$$
\begin{equation*}
\xi_{0}=1-\frac{16}{c r_{\max }} \tag{6.42}
\end{equation*}
$$

We express (6.40) as the sum of two integrals: one from -1 to $\xi_{0}$ and the other from $\xi_{0}$ to 1 , making use of the fact that for $\xi<\xi_{0}$ (i.e. $\operatorname{cr}>8$ ), $f$ is identically zero. Again by isolating the logarithmic singularity and making appropriate changes of variable we wrice ( 6.40 ) as

$$
\begin{align*}
I_{a}= & \frac{1}{4}\left(1+\xi_{0}\right) r_{\max } \int_{-1}^{1} g\left[\frac{(1-s) r_{\max }}{2 r_{0}} ; r_{, i} ; n_{i}\right] d \xi \\
& +\frac{1}{4}\left(1-\xi_{0}\right) r_{\max } \int_{-1}^{1} \ln \left(\frac{c r_{\max }}{4}\right) f\left[\frac{(1-t) r_{\max }}{2 r_{0}} ; r_{, i} ; n_{i}\right]+g\left[\frac{(1-t) r_{\max }}{2 r_{0}} ; r_{, i} ; n_{i}\right] d \xi \\
& +\frac{1}{4} r_{\max }\left(1-\xi_{0}\right) \ln \left(1-\xi_{0}\right) \int_{-1}^{1} f\left[\frac{\left(u r_{\max }\right.}{2 r_{0}} ; r_{, i} ; n_{i}\right] d \xi \\
& -\frac{1}{2} r_{\max }\left(1-\xi_{0}\right) \int_{0}^{1} \ln \left(\frac{1}{\xi}\right) f\left[\frac{v r_{\max }}{2 r_{0}} ; r_{, i} ; n_{i}\right] d \xi \tag{6.43}
\end{align*}
$$

where

$$
\begin{aligned}
& s=\frac{1}{2}\left(1+\xi_{0}\right)(1+\xi)-1 \\
& t=\frac{1}{2}\left[1+\xi_{0}+\xi\left(1-\xi_{0}\right)\right] \\
& u=\frac{1}{2}\left(1-\xi_{0}\right)(1+\xi) \\
& v=\xi\left(1-\xi_{0}\right)
\end{aligned}
$$

Case (b): Integration from $\left(x_{0}, z_{0}\right)$ to $\left(x_{j+1}, z_{j+1}\right)$
Here we define $r_{\max }$ as the distance from $\left(x_{0}, z_{0}\right)$ to $\left(x_{j+1}, z_{j+1}\right)$ so that

$$
\begin{equation*}
r_{\max }=\left[\left(x_{j+1}-x_{0}\right)^{2}+\left(z_{j+1}-z_{0}\right)^{2}\right]^{\frac{1}{2}} \tag{6.44}
\end{equation*}
$$

We put $\left(x_{e}, z_{e}\right)=\left(x_{0}, z_{0}\right)$ and $\left(x_{e+1}, z_{e+1}\right)=\left(x_{j+1}, z_{j+1}\right)$ in (6.32) to (6.36) to find

$$
\begin{array}{cc}
r=\frac{1}{2}(1+\xi) r_{\max } & \\
r_{, 1}=\frac{x_{j+1}-x_{0}}{r_{\max }} & r_{, 2}=\frac{z_{j+1}-z_{0}}{r_{m a x}} \\
n_{1}=-r_{, 2} & n_{2}=r_{, 1} \\
\frac{\partial r}{\partial n} \equiv r_{, i} n_{i}=0 & \tag{6.45}
\end{array}
$$

As before, it is necessary to consider two sub-cases :
(1) $c r_{\max }<8$

The result is

$$
\begin{align*}
I_{b}= & \frac{1}{2} r_{\max } \int_{-1}^{1} \ln \left(\frac{c r_{\max }}{4}\right) f\left[\frac{(1+\xi) r_{\max }}{2 r_{0}} ; r_{, i} ; n_{i}\right]+g\left[\frac{(1+\xi) r_{\max }}{2 r_{0}} ; r_{, i} ; n_{i}\right] d \xi \\
& +\frac{1}{4} r_{\max } \int_{-1}^{1} \ln \left(\frac{3+\xi}{2}\right) f\left[\frac{(3+\xi) r_{\max }}{4 r_{0}} ; r_{, i} ; n_{i}\right] d \xi \\
& -\frac{1}{2} r_{\max } \int_{0}^{1} \ln \left(\frac{1}{\xi}\right) f\left[\frac{\left(\xi r_{\max }\right.}{2 r_{0}} ; r_{, i} ; n_{i}\right] d \xi \tag{6.46}
\end{align*}
$$

(2) $c r_{\text {max }}>8$

Again, we define $\xi_{0}$ by the relation $c r=8$ at $\xi=\xi_{0}$. From ( 6.45 ) we find that

$$
\begin{equation*}
\xi_{0}=\frac{16}{c r_{\max }}-1 \tag{6.47}
\end{equation*}
$$

Noting again that $f \equiv 0$ when $\boldsymbol{\xi}>\xi_{0}$ we have the final result as

$$
I_{b}=\frac{1}{4}\left(1-\zeta_{0}\right) r_{\max } \int_{-1}^{1} g\left[\frac{(1+s) r_{\max }}{2 r_{0}} ; r_{, i} ; n_{i}\right] d \xi
$$

$$
\begin{align*}
& +\frac{1}{4}\left(1+\xi_{0}\right) r_{\max } \int_{-1}^{1} \ln \left(\frac{c r_{\max }}{4}\right) f\left[\frac{(1+t) r_{\max }}{2 r_{0}} ; r_{i} ; n_{i}\right]+g\left[\frac{(1+t) r_{\max }}{2 r_{0}} ; r_{, i} ; n_{i}\right] d \xi \\
& +\frac{1}{4}\left(1+\xi_{0}\right) \ln \left(1+\xi_{0}\right) r_{\max } \int_{-1}^{1} f\left[\frac{\left(u r_{\max }\right.}{2 r_{0}} ; r_{, i} ; n_{i}\right] d \xi \\
& -\frac{1}{2}\left(1+\xi_{0}\right) r_{\max } \int_{0}^{1} \ln \left(\frac{1}{\xi}\right) f\left[\frac{\left(v r_{\max }\right.}{2 r_{0}} ; r_{, i} ; n_{i}\right] d \xi \tag{6.48}
\end{align*}
$$

where

$$
\begin{aligned}
& s=\frac{1}{2}\left[\xi\left(1-\xi_{0}\right)+1+\xi_{0}\right] \\
& t=\frac{1}{2}\left[\xi\left(1+\xi_{0}\right)-1+\xi_{0}\right] \\
& u=\frac{1}{2}\left(1+\xi_{0}\right)(1+\xi) \\
& v=\xi\left(1+\xi_{0}\right)
\end{aligned}
$$

We now write ( 6.31 ) for the case $e=j$, i.e.

$$
I \text { on } \Gamma_{j}=I_{a}+I_{b}
$$

where $I_{a}$ is given by (6.41) or (6.43) and $I_{b}$ is given by (6.46) or (6.48) depending on whether $c r_{\text {max }}$ is less than or greater than 8 . We note that in (6.41) and (6.43) the quantities $r, r_{i}, n_{i}$ and $\xi_{0}$ are given by ( 6.38 ), ( 6.39 ) and ( 6.42 ); whereas in (6.46) and (6.48) these quantities are given by (6.44), (6.45), and (6.47). The integrals $I_{a}$ and $I_{b}$ are evaluated by standard Gaussian quadrature. We draw attention to the fact that the last lerm in the expressions for $I_{a}$ and $I_{b}$ involves a logarithmic singularity. These terms are written in a form suitable for the application of the quadrature rules designed for such cases (Brebbia ct al., i乡̄̄4).

### 6.4 Determination of Unknown Boundary Data

The discretised form (6.9) of the boundary integral equation may be written

$$
\begin{align*}
& \sum_{e=1}^{M}\left(R_{k i}^{e}(j) \tilde{\tilde{T}}_{i}^{e}+S_{k i}^{e}(j) \tilde{\bar{u}}_{i}^{e}\right)=0  \tag{6.49}\\
& i, k=1 \text { to } 3 ; j=1 \text { to } M
\end{align*}
$$

where the argument $j$ has been included to emphasize the fact that ( 6.49 ) must be written for each node on the boundary $\Gamma$, i.e. each node must serve in turn as node $j$, the node at which $r=0$. The coeffic: $\urcorner$ nts $R_{k i}^{e}$ and $S_{k i}^{e}$ are determined using the methods described in the previous section. In a well posed problem, exactly half of the boundary data is known. At each node the boundary data consists of six quantities: $\left\{\tilde{\tilde{T}}_{i}^{e}, \tilde{u}_{i}^{e}, i=1\right.$ to 3$\}$.

We let $\left\{x_{i}^{\prime e}, i=1\right.$ to 3$\}$ denote the three known values at node $e$, and $\left\{x_{i}^{e}, i=\right.$ 1 to 3 \} denote the three unknown values at node $e$. We denote the coefficients of $x_{i}^{\prime e}$ and $x_{i}^{e}$ in (6.49) by $F_{k i}^{\prime e}$ and $F_{k i}^{e}$ respectively. Then equation (6.49) may be written

$$
\begin{array}{r}
\sum_{e=1}^{M} F_{k i}^{e}(j) x_{i}^{e}=b_{k}(j)  \tag{6.50}\\
i, k=1 \text { to } 3 ; j=1 \text { to } M
\end{array}
$$

where

$$
\begin{equation*}
b_{k}(j)=-\sum_{e=1}^{M} F_{k i}^{\prime e}(j) x_{i}^{\prime e} \tag{6.51}
\end{equation*}
$$

For a given value of $j$, equation ( 6.50 ) generates three rows of the final matrix equation as $k$ takes the values 1 to 3. Specifically, (6.50) is the $n$th row of the matrix equation, where

$$
n=3 j-(3-k) \quad k=1 \text { to } 3 ; j=1 \text { to } M
$$

The final $3 M \times 3 M$ matrix equation may be written in the form

$$
\begin{equation*}
[G]\{y\}=\{c\} \tag{6.52}
\end{equation*}
$$

It can be verified that the elements of the matrix $[G]$ and vector $\{c\}$ are produced by the following algorithm :

$$
\begin{aligned}
& \text { For } j=1 \text { to } M \\
& \text { For } k=1 \text { to } 3 \\
& n=3 j-(3-k) \\
& c_{n}=-\sum_{e=1}^{M} F_{k i}^{\prime e} x_{i}^{\prime e} \\
& \text { For } t=1 \text { to } M \\
& \text { For } p=3 t-2 \text { to } 3 t \\
& G_{n p}=F_{k, p-3(t-1)}^{t}
\end{aligned}
$$

A given input value $x_{i}^{\prime e}$ must be identified either as $\tilde{T}_{i}^{e}$ or $\tilde{\tilde{u}}_{i}^{e}$ so that proper assignments can be made to the quantities $F_{k i}^{\prime e}$ and $F_{k i}^{e}$. The solution vector $\{y\}$ consists of the $3 M$ unknown nodal values $\left\{x_{i}^{e}, i=1\right.$ to $3, e=1$ to $\left.M\right\}$.

### 6.5 Determination of Interior Stresses and Pore

## Pressure

We recall that for an interior point $\underline{x}$ the coefficients $\alpha_{i k}$ and $\theta$ in equations (5.131) and (5.132) respectively become

$$
\begin{aligned}
\alpha_{i k} & =\delta_{i k} \\
\theta & =1
\end{aligned}
$$

Using the definitions (6.3) and (6.10),(6.11), (6.12) we write (5.131) and (5.132) for an interior point $\underline{x}$ :

$$
\begin{array}{r}
2 \pi \bar{u}_{k}(\underline{x})=\int_{\Gamma} \tilde{\tilde{T}}_{\alpha}\left(\underline{x}^{\prime}\right) \tilde{\bar{u}}_{\alpha}^{* k}\left(r, \frac{\partial r}{\partial \underline{x}^{\prime}}, \underline{n}\right)-\tilde{\bar{T}}_{\alpha}^{* k}\left(r, \frac{\partial r}{\partial \underline{x}^{\prime}}, \underline{n}\right) \tilde{\bar{u}}_{\alpha}\left(\underline{x}^{\prime}\right) d \Gamma \\
k=1,2 \\
\frac{2 \pi r_{0}}{\mu} \bar{p}(\underline{x})=\int_{\Gamma} \tilde{\bar{T}}_{\alpha}\left(\underline{x}^{\prime}\right) \tilde{\tilde{u}}_{\alpha}^{* 3}\left(r, \frac{\partial r}{\partial \underline{x}^{\prime}}, \underline{n}\right)-\tilde{\bar{T}}_{\alpha}^{* 3}\left(r, \frac{\partial r}{\partial \underline{x}^{\prime}}, \underline{n}\right) \tilde{\tilde{u}}_{\alpha}\left(\underline{x}^{\prime}\right) d \Gamma \tag{6.54}
\end{array}
$$

where $r=\left|x^{\prime}-x\right|$. In the above equations the integration is performed with respect to $\underline{x}^{\prime} \in \Gamma$, and the subscript $\alpha$ is summed from 1 to 3 . In order to determine the effective stresses at point $\boldsymbol{x}$, we must evaluate derivatives of $\bar{u}_{k}$ with respect to $\boldsymbol{x}$ and then use the constitutive laws. We note that a comma followed by a subscript denotes differentiation with respect to one of the coordinates of $\underline{x}^{\prime}=\left(x_{1}^{\prime}, x_{2}^{\prime}\right)$. Differentiation with respect to $\boldsymbol{x}$ will be written explicitly. For example, since $r^{2}=\left(x_{j}^{\prime}-x_{j}\right)\left(x_{j}^{\prime}-x_{j}\right)$, we have

$$
\begin{array}{r}
r_{, i}=\frac{x_{i}^{\prime}-x_{i}}{r}, \frac{\partial r}{\partial x_{i}}=-r_{, i} \\
\text { and } \frac{\partial}{\partial x_{j}}\left(r_{, i}\right)=\frac{1}{r}\left(r_{, i} r_{, j}-\delta_{i j}\right)
\end{array}
$$

Again, repeated suffixes denotes summation. Using the constitutive relation (3.9) with

$$
\lambda=\frac{2 \mu \nu}{1-2 \nu}
$$

we can write from (6.53)

$$
\begin{gather*}
\frac{2 \pi}{\mu} \tilde{\tau}_{k j}^{\prime}=\int_{\Gamma}\left(\tilde{\bar{T}}_{\alpha} D_{\alpha k j}-\tilde{\bar{u}}_{\alpha} S_{\alpha k j}\right) d \Gamma  \tag{6.55}\\
k, j=1,2, \alpha=1,2,3
\end{gather*}
$$

where

$$
\begin{align*}
& D_{\alpha k j}=\frac{\partial \tilde{u}_{\alpha}^{* k}}{\partial x_{j}}+\frac{\partial \tilde{u}_{\alpha}^{* j}}{\partial x_{k}}+\left(\frac{2 \nu}{1-2 \nu}\right) \frac{\partial \tilde{\tilde{u}}_{\alpha}^{* i}}{\partial x_{i}} \delta_{k j}  \tag{6.56}\\
& S_{\alpha k j}=\frac{\partial \tilde{\bar{T}}_{\alpha}^{* k}}{\partial x_{j}}+\frac{\partial \tilde{T}_{\alpha}^{* j}}{\partial x_{k}}+\left(\frac{2 \nu}{1-2 \nu}\right) \frac{\partial \tilde{T}_{\alpha}^{* i}}{\partial x_{i}} \delta_{k j} \tag{6.57}
\end{align*}
$$

Using the definitions (6.29), (6.30) and Appendix $G$ we can perform the differentiations indicated in (6.56) and (6.57). The algebra involved is again quite extensive. The results are listed in Appendix H.
Discretisation of (6.55) leads to

$$
\begin{equation*}
\frac{2 \pi}{\mu} \tilde{\tau}_{k j}^{\prime}(\underline{x})=\sum_{e=1}^{M}\left(\tilde{\bar{T}}_{\alpha}^{e} P_{\alpha k j}^{e}-\tilde{\tilde{u}}_{\alpha}^{e} Q_{\alpha k j}^{e}\right) \tag{6.58}
\end{equation*}
$$

where

$$
\begin{array}{r}
P_{\alpha k j}^{e}=\int_{\Gamma_{e}} D_{\alpha k j} d \Gamma_{e} \\
Q_{\alpha k j}^{e}=\int_{\Gamma_{e}} S_{\alpha k j} d \Gamma_{e}  \tag{6.59}\\
k, j=1,2 ; \quad \alpha=1,2,3 ; \quad e=1 \text { to } M
\end{array}
$$

Similarly, the pore pressure at $x$ is found by discretising ( 6.54):

$$
\begin{equation*}
\frac{2 \pi r_{0}}{\mu} \bar{p}(\underline{x})=\sum_{e=1}^{M} \tilde{\tilde{T}}_{\alpha}^{e} A_{\alpha}^{e}-\tilde{\bar{u}}_{\alpha}^{e} B_{\alpha}^{e} \tag{6.60}
\end{equation*}
$$

where

$$
\begin{align*}
A_{\alpha}^{e} & =\int_{\Gamma_{e}} \tilde{u}_{\alpha}^{* 3} d \Gamma_{e} \\
B_{\alpha}^{e} & =\int_{\Gamma_{e}} \tilde{T}_{\alpha}^{* 3} d \Gamma_{e} \\
\alpha=1 \text { to } 3, e=1 \text { to } M & \tag{6.61}
\end{align*}
$$

Since $\underline{x}$ is not on the boundary $\Gamma$, the integrals ( 6.60 ) and ( 6.61 ) are all nonsingular and are evaluated by standard Gaussian quadrature.

### 6.6 Wave loading of Flat Homogeneous Isotropic

## Seabed. Comparison with Analytic Solu-

## tion.

We consider a horizontal distance of one quarter of the wavelength $L$, as illustiated in Figure (6.1). The $x$ axis is at the mudline as shown and the $z$ axis is vertically downward. The wave pressure on $O C$ is of the form

$$
P=P_{0} e^{i(k x-\omega t)}
$$

where $k=\frac{2 \pi}{L}$ and $\omega$ is the circular frequency. Hence

$$
\begin{equation*}
\bar{P}=P_{0} e^{i k x} \tag{6.62}
\end{equation*}
$$

From linear wave theory, the formula for $P_{0}$ is given by

$$
\begin{equation*}
P_{0}=\frac{\rho_{f}\left(\frac{H}{2}\right) g}{\cosh (k h)} \tag{6.63}
\end{equation*}
$$



Fig.(6.1) BEM Test for Flat Seabed.
where $H$ is the wave height and $h$ is the water depth. From (3.7) and (5.3) we have

$$
\begin{equation*}
T_{i}=\tau_{i j}^{\prime} n_{j}-p n_{i} \quad(i=1,2) \tag{6.64}
\end{equation*}
$$

and we recall that

$$
\begin{equation*}
T_{3}=-p \tag{6.65}
\end{equation*}
$$

The boundary conditions are as follows:
On $O C, \bar{\tau}_{12}^{\prime}=0, \bar{\tau}_{22}^{\prime}=0, \bar{p}=P_{0} e^{i k x}$.
In terms of $\tilde{\bar{T}}_{i}$ these become

$$
\begin{equation*}
\tilde{\tilde{T}}_{1}=0, \tilde{\tilde{T}}_{2}=\frac{P_{0}}{\mu} e^{i k x}, \tilde{\tilde{T}}_{3}=-\frac{P_{0}}{\mu} e^{i k x} \tag{6.66}
\end{equation*}
$$

On $A B$ :

$$
\begin{equation*}
\tilde{\bar{u}}_{i}=0 \text { ior } i=1,2,3 \tag{6.67}
\end{equation*}
$$

Periodicity conditions must be used on the side boundaries $O A$ and $B C$. A given function $f(x, z)$ may be expressed as

$$
\begin{equation*}
f(x, z)=F(z) e^{i k x} \tag{6.68}
\end{equation*}
$$

where $f(x, z)$ represents $\bar{u}_{i}, \bar{U}_{i}, \bar{\tau}_{i j}^{\prime}$ or $\bar{p}$. From (6.68) we deduce that since $k=\frac{2 \pi}{L}$

$$
\begin{equation*}
f\left(\frac{L}{4}, z\right)=i f(0, z) \tag{6.69}
\end{equation*}
$$

Equation (6.69) can be used to express the boundary conditions on $B C$ in terms of those on $O A$. It is easy to show that

$$
\begin{align*}
& \tilde{\bar{u}}_{i}\left(\frac{L}{4}, z\right)=\alpha_{u}^{i} \tilde{\bar{u}}_{i}(0, z) \text { no sum on } i  \tag{6.70}\\
& \tilde{\bar{T}}_{i}\left(\frac{L}{4}, z\right)=\alpha_{T}^{i} \tilde{\bar{T}}_{i}(0, z) \text { no sum on } i
\end{align*}
$$

where $\alpha_{u}^{n}$ and $\alpha_{T}^{n}$ are complex constants defined by

$$
\begin{aligned}
& \alpha_{u}^{n}=\left\{\begin{array}{cc}
i & n=1,2 \\
-i & n=3
\end{array}\right. \\
& \alpha_{T}^{n}=\left\{\begin{array}{cc}
-i & n=1,2 \\
i & n=3
\end{array}\right.
\end{aligned}
$$

Conditions (6.66), (6.67) and (6.70) must now be used in the discretised boundary integral equation (6.9). Since the boundary conditions on $B C$ are expressed in terms of those on $O A$ we choose the same number of elements $M_{1}$ on each of these boundary segments. On $A B$ and $C O$ we choose $M_{2}$ and $M_{3}$ elements respectively. The total number of elements is thus $M=2 M_{1}+M_{2}+M_{3}$. The first element is taken on $O A$ and the last one on $C O$. The direction of integration is anticlockwise around the boundary.

Using (6.70) we have

$$
\begin{aligned}
\tilde{\bar{T}}_{i}^{M_{1}+M_{2}+e} & =\alpha_{T}^{i} \tilde{\bar{T}}_{i}^{M_{1}+1-e} \\
\tilde{\bar{u}}_{i}^{M_{2}+M_{2}+e} & =\alpha_{u}^{i} \tilde{u}_{i}^{M_{1}+1-e} \\
e & =1 \text { to } M_{1}, \text { no sum on } i
\end{aligned}
$$

We let

$$
\begin{aligned}
R_{k i}^{M_{1}+M_{2}+e} & =R_{k i}^{\prime} M_{1}+1-e \\
S_{k i}^{M_{1}+M_{2}+e} & =S_{k i}^{\prime M_{1}+1-e} \\
e & =1 \text { to } M_{1}
\end{aligned}
$$

Then it is easily shown that

$$
\begin{aligned}
& \sum_{e=1}^{M_{1}} R_{k i}^{M_{1}+M_{2}+e} \tilde{\bar{T}}_{i}^{M_{1}+M_{2}+e}=\sum_{e=1}^{M_{1}} \alpha_{T}^{i} R_{k i}^{\prime e} \tilde{T}_{i}^{e} \\
& \sum_{e=1}^{M_{1}} S_{k i}^{M_{1}+M_{2}+e} \tilde{\bar{u}}_{i}^{M_{1}+M_{2}+e}=\sum_{e=1}^{M_{1}} \alpha_{u}^{i} S_{k i}^{\prime e} \tilde{\bar{u}}_{i}^{e}
\end{aligned}
$$

Substituting into (6.9) gives

$$
\begin{align*}
& \sum_{e=1}^{M_{1}}\left(R_{k i}^{e}+\alpha_{T}^{i} R_{k i}^{\prime e}\right) \tilde{T}_{i}^{e}+\left(S_{k i}^{e}+\alpha_{u}^{i} S_{k i}^{\prime e}\right) \tilde{\bar{u}}_{i}^{e} \\
& +\sum_{e=1}^{M_{2}}\left(R_{k i}^{M_{1}+e} \tilde{\bar{T}}_{i}^{M_{1}+e}+S_{k i}^{M_{1}+e} \tilde{\bar{u}}_{i}^{M_{1}+e}\right)  \tag{6.71}\\
& +\sum_{e=1}^{M_{3}}\left(R_{k i}^{2 M_{1}+M_{2}+e} \tilde{\bar{T}}_{i}^{2 M_{1}+M_{2}+e}+S_{k i}^{2 M_{1}+M_{2}+e}{\tilde{u_{i}}}_{i=1}^{2 M_{1} M_{2}+e}\right)=0
\end{align*}
$$

We now deline the following
For $e=1$ to $M_{1}$

$$
\begin{aligned}
& R_{k i}^{* e}=R_{k i}^{e}+\alpha_{T}^{i} R_{k i}^{\prime e} \\
& S_{k i}^{* e}=S_{k i}^{e}+\alpha_{u}^{i} S_{k i}^{\prime e} \\
& T_{i}^{* e}=\tilde{\tilde{T}}_{i}^{e} \\
& u_{i}^{* e}=\tilde{\tilde{u}}_{i}^{e}
\end{aligned}
$$

For $e=1$ to $M_{2}$ :

$$
\begin{aligned}
& R_{k i}^{* M_{1}+e}=R_{k i}^{M_{1}+e} \\
& S_{k i}^{* M_{1}+e}=S_{k i}^{M_{1}+e} \\
& T_{i}^{* M_{1}+e}=\tilde{T}_{i}^{M_{1}+e} \\
& u_{i}^{* M_{1}+e}=\tilde{\bar{u}}_{i}^{M M_{1}+e}
\end{aligned}
$$

For $e=1$ to $M_{3}$

$$
\begin{aligned}
& R_{k i}^{* M_{1}+M_{2}+e}=R_{k i}^{2 M_{1}+M_{2}+e} \\
& S_{k i}^{w M_{1}+M_{2}+e}=S_{k i}^{2 M_{1}+M_{2}+e} \\
& T_{i}^{* M_{1}+M_{2}+e}=\tilde{T}_{i}^{2 M_{1}+M_{2}+e} \\
& u_{i}^{* M_{1}+M_{2}+e}=\tilde{\bar{u}}_{i}^{2 M_{2}+M_{2}+e}
\end{aligned}
$$

With these definitions, equation (6.71) becomes

$$
\begin{equation*}
\sum_{e=1}^{M^{\bullet}} R_{k i}^{* e}(j) T_{i}^{* e}+S_{k i}^{* e}(j) u_{i}^{* e}=0 \quad i, k=1 \text { to } 3 \tag{6.72}
\end{equation*}
$$

where $M^{*}=M_{1}+M_{2}+M_{3}$
The unknowns are

$$
\begin{align*}
& \left\{T_{i}^{* e}, u_{i}^{* e}, e=1 \text { to } M_{1}\right\} \\
& \left\{T_{i}^{* M_{1}+e}, e=1 \text { to } M_{2}\right\}  \tag{6.73}\\
& \left\{u_{i}^{* M_{1}+M_{2}+e}, e=1 \text { to } M_{3}\right\}
\end{align*}
$$

where $i=1$ to 3 , i.e. a total of $\left(6 M_{1}+3 M_{2}+3 M_{3}\right)$ or $3 M$ unknowns. Equation (6.72) must be written with each boundary node serving as $r=0$, i.e. as $j$ takes the values 1 to $M$. This produces $3 M$ equations. For a given $j$, row number $3 j-(3-k)$ of the matrix equation is generated.

For $e=1$ to $\left(M_{2}+M_{3}\right)$ we define

$$
\left.\begin{array}{l}
x_{i}^{\prime e}=T_{i}^{* M_{1}+e} \\
F_{k i}^{\prime e}=R_{k i}^{* M_{1}+e} \\
x_{i}^{e}=u_{i}^{* M_{1}+e} \\
F_{k i}^{e}=S_{k i}^{* M_{1}+e}
\end{array}\right\} \text { if } T_{i}^{* M_{1}+e} \text { is known }
$$

and

$$
\left.\begin{array}{l}
x_{i}^{\prime e}=u_{i}^{* M_{1}+e} \\
F_{k i}^{\prime e}=\cdot{ }_{k i}^{* M_{1}+e} \\
x_{i}^{e}=T_{i}^{* M_{1}+e} \\
F_{k i}^{e}=R_{k i}^{* M_{1}+e}
\end{array}\right\} \text { if } u_{i}^{* M_{1}+e} \text { is known }
$$

Then equation ( 6.72 ) becomes

$$
\begin{equation*}
\sum_{e=1}^{M_{1}}\left(R_{k i}^{* e} T_{i}^{* e}+S_{k i}^{* e} u_{i}^{* e}\right)+\sum_{e=1}^{M_{2}+M_{3}} F_{k i} x_{i}^{e}=-\sum_{e=1}^{M_{2}+M_{3}} F_{k i}^{\prime e} x_{i}^{\prime e} \tag{6.74}
\end{equation*}
$$

This may be written in matrix form

$$
\begin{equation*}
[G]\{y\}=\{b\} \tag{6.75}
\end{equation*}
$$

It can be verified that the following algorithm generates the matrix $[G]$ and vector \{b\}:

For $j=1$ to $M$
For $k=1$ to 3

$$
\begin{aligned}
& n=3 j-(3-k) \\
& b_{n}=-\sum_{e=1}^{M_{2}+M_{3}} F_{k i}^{\prime e} x_{i}^{\prime e} \\
& G_{n p}=\left\{\begin{array}{l}
R_{k, p-3 t+3}^{* t} p=3 t-2 \text { to } 3 t, t=1 \text { to } M_{1} \\
S_{k, p-3\left(M_{1}+t-1\right)}^{* t} p=3 M_{1}+3 t-2 \text { to }\left(3 M_{1}+3 t\right), t=1 \text { to } M_{1} \\
F_{k, p-3\left(2 M_{1}+t-1\right)}^{t} p=\left(6 M_{1}+3 t-2\right) \text { to }\left(6 M_{1}+3 t\right), t=1 \text { to }\left(M_{2}+M_{3}\right)
\end{array}\right.
\end{aligned}
$$

The solution vector is given by (6.73). The interior stresses are computed as described in Section (6.5). In order to test the accuracy of the BEM we compare

Table 6.1: Comparison of BEM with Analytic Solution at (1,2)

| Stresses <br> $\times 10^{4}$ <br> $\mathrm{Nm}^{-2}$ |  | $\overline{\mathrm{BEM}}$ (20r |  | 'ements) |
| :---: | :---: | :---: | :---: | :---: |
|  | Real | Imag | Analytic |  |
| $\bar{\tau}_{11}^{\prime}$ | 0.9047 | 0.1253 | 0.9068 | 0.1310 |
| $\bar{\tau}_{22}^{\prime}$ | -0.9104 | 0.2007 | -0.9064 | 0.1952 |
| $\bar{\tau}_{12}^{\prime}$ | -0.0075 | 0.1535 | -0.0075 | 0.1527 |
| $\bar{p}$ | 4.5129 | 0.3004 | 4.5116 | 0.3044 |

Table 6.2: Comparison of BEM with Analytic Solution at $(5,5)$

| Stresses <br> $\times 10^{4}$ <br> $\times 1 m^{-2}$ |  | BEM (200 elements) |  | Analytic |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | Real | Imag | Real | Imag |  |
| $\bar{\tau}_{11}^{\prime}$ | 0.8639 | 0.0845 | 0.8639 | 0.0852 |  |
| $\bar{\tau}_{\tau 2}^{0}$ | -0.9307 | -0.1081 | -0.9310 | -0.1088 |  |
| $\bar{\tau}_{12}^{\prime}$ | -0.0431 | 0.3670 | -0.0430 | 0.3666 |  |
| $\bar{p}$ | 4.4466 | 0.4330 | 4.4449 | 0.4342 |  |

the results with the analytic solution which is briefly described in Appendix I. The data for the test case was $O A=25 \mathrm{~m}$, water depth $=70 \mathrm{~m}$, wave period $=$ 15 sec . (which corresponds to a wavelength $L$ of 311.812 m ), wave height $=24 \mathrm{~m}$. Details of the input data for the soil (fine sand) can be found in Chapter 8. Tables (6.1), (6.2) and (6.3) show the stresses in $N / m^{2}$ at three interior points (coordinates defined by axes in Figure (6.1) ) obtained by the two methods. There is excellent agreement between the BEM and the analytic solution. The BEM results were obtained with eight point Gaussian quadrature.

Table 6.3: Comparison of BEM with Analytic Solution at $(35,12)$

| Stresses <br>  <br>  |  | BEM (200 elements) |  | Analytic |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | Real | Imag | Real | Imag |  |
| $\bar{\tau}_{11}^{\prime}$ | 0.5517 | 0.4725 | 0.5517 | 0.4725 |  |
| $\bar{\tau}_{22}^{\prime}$ | -0.5790 | -0.4955 | -0.5790 | -0.4955 |  |
| $\bar{\tau}_{12}^{\prime}$ | -0.5811 | 0.6753 | -0.5809 | 0.6749 |  |
| $\bar{p}$ | 3.4689 | 2.9490 | 3.4686 | 2.9489 |  |

## Chapter 7

## Stress Analysis of Sloping Seabed Under Wave Loading

### 7.1 Wave Induced Effective Stresses and Pore Pressure in a Bed of Arbitrary Slope

The problem domain is illustrated in Figure (7.1). The $x$ and $z$ axes are chosen with origin at $O^{\prime}$ as shown. The incident wave is travelling in the negative $x$ direction. The boundary conditions on the seabed $B D E C$ are that the normal and shear effective stresses are zero (there is no applied inter-granular load) and the pore pressure $\bar{p}$ is equal to the wave pressure $\bar{P}$ acting normal to the bed (determined from equation 4.46). In terms of the tractions $\bar{T}_{i}$, these become :

On BD : $\bar{T}_{1}=0, \bar{T}_{2}=\left.i \omega \rho_{f} \bar{\phi}\right|_{B D}, \bar{T}_{3}=-\left.i \omega \rho_{f} \bar{\phi}\right|_{B D}$


On DE : $\bar{T}_{1}=-\left.i \omega \rho_{f} \bar{\phi}\right|_{D E} \sin \beta, \bar{T}_{2}=\left.i \omega \rho_{f} \bar{\phi}\right|_{D E} \cos \beta, \bar{T}_{3}=-\left.i \omega \rho_{f} \bar{\phi}\right|_{D E}$
On EC : $\bar{T}_{1}=0, \bar{T}_{2}=\left.i \omega \rho_{f} \bar{\phi}\right|_{E C}, \bar{T}_{3}=-\left.i \omega \rho_{f} \bar{\phi}\right|_{E C}$
where $\bar{\phi}$ is the velocity potential in the sea as determined from the techniques of Chapter 4. At the interface with the bedrock base $R S$ we have:

$$
\bar{u}_{1}=\bar{u}_{2}=\bar{u}_{3}=0
$$

To provide boundary conditions on the side boundaries $B R$ and $S C$ we make use of the author's analytic solution (hereafter referred to as A.S.) described in Appendix I. We first determine from (4.15) the velocity potential on the seabed in the region to the right of $O^{\prime} C$. The coefficients $a_{n}$ in (4.15) are determined by solving (4.37) for the vector $\{a\}$. The numerical results indicate that the infinite sum on the right hand side of (4.15) is negligible compared to the other terms. Neglecting this term, we write the velocity potential on the seabed to the right of $O^{\prime} C$ as :

$$
\begin{equation*}
\left.\Phi\right|_{b e d}=\frac{-i g \eta_{0}}{\omega \cosh \left(k_{0} h_{0}\right)}\left[a_{0} e^{i k_{0} x}+e^{-i k_{0} x}\right] e^{-i \omega t} \tag{7.1}
\end{equation*}
$$

Hence from (4.46) the wave pressure on the seabed in this region is :

$$
\begin{equation*}
\left.P\right|_{b e d}=\frac{\rho_{J} g \eta_{0}}{\cosh \left(k_{0} h_{0}\right)}\left[a_{0} e^{i k_{0} x}+e^{-i k_{0} x}\right] e^{-i \omega t} \tag{7.2}
\end{equation*}
$$

which is in a form suitable for applying the A.S. We can determine the effective stresses and pore pressure (and hence $\bar{T}_{i}$ ) on SC by superposing the results obtained from the A.S. for the following two cases:
(1) Incident Wave conditions:

$$
\begin{aligned}
k & =-k_{0} \\
P_{0} & =\frac{\rho_{f} g \eta_{0}}{\cosh \left(k_{0} h_{0}\right)}
\end{aligned}
$$

(2) Reflected Wave conditions:

$$
\begin{aligned}
k & =k_{0} \\
P_{0} & =\frac{a_{0} \rho_{j} g \eta_{0}}{\cosh \left(k_{0} h_{0}\right)}
\end{aligned}
$$

The $\bar{\tau}_{i j}^{\prime}$ and $\vec{p}$ at desired soil depths are evaluated at $x=0$ in each case. We can now write the boundary conditions on $S C$, noting that $\left(n_{1}, n_{2}\right)=(1,0)$ :

$$
\bar{T}_{1}=\bar{\tau}_{11}^{\prime}-\bar{p}, \bar{T}_{2}=\bar{\tau}_{21}^{\prime}, \bar{T}_{3}=-\bar{p}
$$

For the region to the left of $B R$, the wave velocity potential on the seabed is, from (4.16)

$$
\begin{equation*}
\left.\Phi\right|_{b e d}=a_{0}^{\prime} e^{-i k_{0}^{\prime} x} e^{-i \omega t} \tag{7.3}
\end{equation*}
$$

where we have neglected the infinite sum on the right hand side of (4.16). The coefficient $a_{0}^{\prime}$ may be determined from the numerically determined value of $\bar{\phi}$ at $B$ :

$$
\begin{equation*}
\left.\bar{\phi}\right|_{B}=a_{0}^{\prime} e^{i k_{0}^{\prime} d} \tag{7.4}
\end{equation*}
$$

Again, from (4.46) and (7.3) the wave pressure on the seabed in this region is:

$$
\begin{align*}
\left.P\right|_{\text {bed }} & =i \omega \rho_{f} a_{0}^{\prime} e^{-i k_{0}^{\prime} x} e^{-i \omega t} \\
& =\left.i \omega \rho_{j} \bar{\phi}\right|_{B} e^{-i k_{0}^{\prime} d} e^{-i\left(k_{0}^{\prime} x+\omega t\right)} \tag{7.5}
\end{align*}
$$

where we have used (7.4).
The effective stresses and pore pressure on $B R$ may thus be determined from the A.S. using

$$
k=-k_{0}^{\prime}, \quad P_{0}=\left.i \omega \rho_{j} \bar{\phi}\right|_{B} \mathrm{e}^{-i k_{0}^{\prime} d}, \quad x=-d
$$

Cn $B R,\left(n_{1}, n_{2}\right)=(-1,0)$ and we write the boundary conditions as

$$
\bar{T}_{1}=-\bar{\tau}_{11}^{\prime}+\bar{p}, \quad \bar{T}_{2}=-\bar{\tau}_{21}^{\prime}, \quad \bar{T}_{3}=-\bar{p}
$$

With the boundary conditions thus specified, we determine the unknown boundary data and wave-induced interior stresses as descrioed in Chapter 6.

### 7.2 Initial Stresses

The in-situ stresses in soil depend not only on gravitational forces, but also on the stress history due to the geological processes involved in the formation of the soil deposit. Thus the standard constants of elasticity theory must be modified to adequately model the in-situ stress field. For example, for a flat seabed under hydrostatic conditions the effecti e stresses at depth $z$ are, in conventional notation,

$$
\begin{align*}
\sigma_{z}^{\prime} & =\gamma^{\prime} z \\
\sigma_{x}^{\prime} & =K_{0} \gamma^{\prime} z  \tag{7.6}\\
\tau_{x z}^{\prime} & =0
\end{align*}
$$

where $\gamma^{\prime}$ is the buoyant unit weight of the saturated soil and $\sigma_{x}^{\prime}, \sigma_{x}^{\prime}, \tau_{x z}^{\prime}$ are the effective vertical, horizontal and shear stresses respectively at depth $z$ (soil mechanics sign convention). The parameter $K_{0}$ is the coefficient of lateral earth pressure at rest which attempts to account for the stress history of the soil. It is usually empirically determined. For a sloping bed it is possible to determine the gravitational stress field from the theory of elasticity using a numerical technique (finite element or boundary element method) since no analytic solution exists for the geometry
of the problem considered. In this approach it would be necessary to modify the elastic constants for the reasons discussed above. However, no adequate method of doing this is currently available and we employ the concept of conjugate stresses introduced by Taylor (1948) in which it is possible to account for the stress history of the soil via a parameter analogous to $K_{0}$. The concept is strictly valid only for infinite slopes. Accordingly, the representation of in-situ stresses by this method in the vicinity of the lines DG and EF (Figure 7.1) is only approximate.

We shall first consider an infinite slope of dry soil inclined at an angle $\beta$ as illustrated in Figure (7.2).

The element $A B C D$, located at distance $z$ below the surface of the slope as shown, has sides $A B$ and $C D$ parallel to the slope. Taylor (1948) has shown that the total stresses on the element may be represented by $\sigma_{v}$ in the vertical direction and $\sigma_{\beta}$ parallel to the slope. The stresses $\sigma_{v}$ and $\sigma_{\beta}$ are called conjugate stresses. Considering the weight of material above $A B$ it is easily seen that

$$
\begin{equation*}
\sigma_{v}=\gamma_{\mathrm{d}} z \cos \beta \tag{7.7}
\end{equation*}
$$

where $\gamma_{d}$ is the unit weight of the dry soil. Following Chowdhury (1977), we assume that there exists a constant $K$, called the conjugate stress ratio, such that

$$
\sigma_{\beta}=K \sigma_{v}
$$

i.e.

$$
\begin{equation*}
\sigma_{\beta}=K \gamma_{d} z \cos \beta \tag{7.8}
\end{equation*}
$$

We now determine the Cartesian stress tensor $\tau_{i j}$ at a point. As before, the unit outward normal on a plane is denoted by ( $n_{1}, n_{2}$ ) and the tractions $T_{i}$ are given by


Fig.(7.2) Conjugate Stresses in Infinite Slope
(5.3). On $A B$, Figure (7.2),

$$
T_{1}=0, T_{2}=\sigma_{v}, \quad n_{1}=\sin \beta, n_{2}=-\cos \beta
$$

This gives

$$
\begin{align*}
& \tau_{11} \sin \beta-\tau_{12} \cos \beta=0  \tag{i}\\
& \tau_{21} \sin \beta-\tau_{22} \cos \beta=\gamma_{d} z \cos \beta \tag{ii}
\end{align*}
$$

On $B C$,

$$
T_{1}=-\sigma_{\beta} \cos \beta, T_{2}=-\sigma_{\beta} \sin \beta, \quad n_{1}=1, n_{2}=0
$$

This gives

$$
\begin{align*}
& \tau_{11}=-K \gamma_{d} z \cos ^{2} \beta  \tag{iii}\\
& \tau_{21}=-K \gamma_{d} z \sin \beta \cos \beta \tag{iv}
\end{align*}
$$

From (ii) and (iv) we have

$$
\begin{equation*}
\tau_{2}:=-\gamma_{d} z\left(1+K \sin ^{2} \beta\right) \tag{v}
\end{equation*}
$$

It is easily verified that the expressions giveu by (iii),(iv) and (v) satisfy (i).
The above expressions give the effective stress tensor at a point in a slope of dry soil. If we now consider the effective stress field in a submerged slope we note that $\gamma_{d}$ is now replaced by $\boldsymbol{\gamma}^{\prime}$, the buoyant unit weight of the soil. Hence the effective stress tensor in a submerged slope is given by expressions (iii), (iv) and (v) above with $\gamma_{d}$ replaced by $\gamma^{\prime}$. Using the conventional soil mechanics notation and sign conventions (described in Section 3.2) we have the in-sisu effective stress field in the region $D E F G$ (Figure 7.1) as

$$
\sigma_{z}^{\prime}=\gamma^{\prime} z\left(1+K \sin ^{2} \beta\right)
$$

$$
\begin{align*}
\sigma_{x}^{\prime} & =K \gamma^{\prime} z \cos ^{2} \beta  \tag{7.9}\\
\tau_{x z}^{\prime} & =K \gamma^{\prime} z \sin \beta \cos \beta
\end{align*}
$$

where we recall that $z$ denotes depth below the sloping surface $D E$. We can define $K_{0}$,the coefficient of earth pressure at rest, in the usual way as the ratio of the horizontal to vertical effective stress, i.e.

$$
K_{0}=\frac{K \cos ^{2} \beta}{1+K \sin ^{2} \beta}
$$

The value of $K$ can thus be determined from measurements of $K_{0}$, or from commonly accepted empirical formulae for $K_{0}$ (Bowles, 1984).

The principal stresses $\sigma_{1}^{\prime}, \sigma_{3}^{\prime}$ are given by :

$$
\sigma_{1,3}^{\prime}=\frac{1}{2}\left(\sigma_{x}^{\prime}+\sigma_{x}^{\prime} \pm \sqrt{\frac{1}{4}\left(\sigma_{x}^{\prime}-\sigma_{z}^{\prime}\right)^{2}+\tau_{x z}^{\prime 2}}\right.
$$

which gives

$$
\begin{equation*}
\sigma_{1,3}^{\prime}=\frac{1}{2} \gamma^{\prime} z\left[1+K \pm \sqrt{\left.K^{2}-2 K \cos 2 \beta+1\right]}\right. \tag{7.10}
\end{equation*}
$$

The limiting values of the conjugate stress ratio $K$ are determined by the MohrCoulomb failure criterion which is written for cohesionless soils as

$$
\begin{equation*}
\sigma_{1}^{\prime}-\sigma_{3}^{\prime}=\left(\sigma_{1}^{\prime}+\sigma_{3}^{\prime}\right) \sin \phi^{\prime} \tag{7.11}
\end{equation*}
$$

where $\phi^{\prime}$ is the (effective) angue of internal friction of the soil. From (7.10) and (7.11) we have

$$
\begin{equation*}
K^{2} \cos ^{2} \phi^{\prime}-2 K\left(\cos 2 \beta+\sin ^{2} \phi^{\prime}\right)+\cos ^{2} \phi^{\prime}=0 \tag{7.12}
\end{equation*}
$$

which has roots

$$
\begin{equation*}
K_{1}, K_{2}=\frac{1}{\cos ^{2} \phi^{\prime}}\left[\cos 2 \beta+\sin ^{2} \phi^{\prime} \pm 2 \cos \beta \sqrt{\left.\sin ^{2} \phi^{\prime}-\sin ^{2} \beta\right]}\right. \tag{7.13}
\end{equation*}
$$

Table 7.1: Limiting Values of $K$

| Slope Angle | Limits for $K$ |  |
| :---: | :---: | :---: |
| $\beta$ (deg.) | $\phi=30^{\circ}$ | $\phi=35^{\circ}$ |
| 5 | $0.34-2.95$ | $0.27-3.64$ |
| 10 | $0.35-2.82$ | $0.29-3.50$ |
| 15 | $0.39-2.59$ | $0.31-3.25$ |
| 20 | $0.44-2.27$ | $0.34-2.92$ |
| 25 | $0.54-1.84$ | $0.40-2.50$ |

We note that the roots are real if $\phi^{\prime} \geq \beta$. Equation ( 7.13) gives the limiting values of $K$, i.e. the range of values of $K$ for which the in-situ stresses do not violate the Mohr-Coulomb failure criterion. These values are presented in Table (7.1) for various slope angles and values of the angle $\phi^{\prime}$. The upper and lower limits for $K$ correspond to states of passive and active failure respectively in the undisturbed soil. The initial stresses in the regions $B D G R$ and ECSF (Figure (7.2)) are estimated from equations (7.6),where $z$ is depth below $B D$ and $E C$ respectively.

### 7.3 Failure Analysis

The basis for the failure analysis will be the Mohr-Coulomb failure criterion which is widely accepted in the geotechnical community. We shall identify "zones of overstress" in the soil, i.e. regions of soil in which the failure criterion is violated. It should be noted, however, that when such violation of the failure criterion has taken place, the stress field is, strictly speaking, invalid. The zone of overstress must
therefore be regarded as the approximate region of failure. Further, we note that owing to the effect of stress transfer when the failure condition is violated, the zone of overstress denotes the minimum region of soil failure. An elastoplastic analysis is then required to determine the failure zone more accurately. In this thesis we provide the first stage of such an analysis, via the location of the overstressed zones determined by a single application of the failuire criterion. This type of failure analysis has also been employed by Yamamoto (1978), Mynett and Mei (1982) and Mei and McTigue (1984). The limitations of such analysis was also pointed out by Mei and McTigue (1984) who nevertheless recognised the utility of the approach in providing a good first approximation of the extent of the failure zone.

We employ the conventional soil mechanics notation and sign convention described in Section (3.2). The initial effective stresses determined from Section
 be denoted by $\sigma_{x}^{\prime(1)}, \sigma_{z}^{\prime(1)}, \tau_{x z}^{\prime(1)}$ where we take the real part of the complex stresses computed by the BEM.

The resultant effective stresses $\sigma_{x}^{\prime}, \sigma_{x}^{\prime}, \tau_{x z}^{\prime}$ are thus given by :

$$
\begin{align*}
& \sigma_{x}^{\prime}=\sigma_{x}^{\prime(0)}+\sigma_{x}^{\prime(1)} \\
& \sigma_{z}^{\prime}=\sigma_{z}^{\prime(0)}+\sigma_{z}^{\prime(1)}  \tag{7.14}\\
& \tau_{x z}^{\prime}=\tau_{x z}^{\prime(0)}+\tau_{x z}^{\prime(1)}
\end{align*}
$$

These stresses may be plotted on a Mohr circle as illustrated in Figure (7.3). The ccordinates ( $\sigma, \tau$ ) of any point $D$ on the circle are the normal and shear stress acting on a particular plane, the inclination of which is determined from the angle $D \hat{C} A$ in Figure (7.3).


Fig.(7.3) The Stress Angle $\theta$

The Mohr-Coulomb failure envelope is the line

$$
\begin{equation*}
r=\sigma \tan \phi^{\prime} \tag{7.15}
\end{equation*}
$$

and is illustrated in the figure. It intersects the $\tau$ axis at the origin $O$. We define the stress angle $\theta$ as the angle between the $\sigma$ axis and the tangent to the circle from the failure envelope intercept $O$. The stress angle-is_funetion of the stress state $\left(\sigma_{x}^{\prime}, \sigma_{x}, \tau_{x x}^{\prime \prime}\right)$, and by comparing its value with $\phi$ " we determine the "distance" between the current stress state and the failure state. The condition of failure is thus conveniently expressed as

$$
\begin{equation*}
\theta \geq \phi^{\prime} \tag{7.16}
\end{equation*}
$$

From the geometry of Figure (7.3) we find that

$$
\begin{equation*}
\sin \theta=\frac{\left[\left(\sigma_{z}^{\prime}-\sigma_{x}^{\prime}\right)^{2}+4 \tau_{x z}^{\prime 2}\right]^{\frac{1}{2}}}{\sigma_{x}^{\prime}+\sigma_{z}^{\prime}} \tag{7.17}
\end{equation*}
$$

from which $\theta$ can be calculated. The overstressed zones are then determined from ( 7.16). We would comment at this point that for the purpose of determining the stress angle no significant reduction in accuracy is observed by using four point Gaussian quadrature (instead of eight point quadrature). when computing the waveinduced effective stresses by the BEM.

## Chapter 8

## Results and Discussion

We first present the wave pressures on a sloping seabed as computed by the methods of Chapter 4. Figures (8.1) and (8.2) illustrate the wave pressures on a 12 degree slope due to wave lengths ( $L$ ) of 300 m and 150 m respectively. The coordinates of the points on the slope are as defined in Figure (7.1). The bed slopes upward between $x=-5 m$ and $x=-60 m$ (points E and D in Figure (7.1)). The incident wave height is $24 m$ and the water depth between $x=0 m$ and $x=-5 m$ is $80 m$. At time $t=0$, the wave crest is vertically above the point $x=0 m$ (point C in Figure (7.1)). The graphs illustrate the progression in time of the wave pressures, the maximum pressures being experienced by points under the wave crest. The maxima at the left end of the slope are slightly greater than at the right end owing to the smaller water depth at the left end. We note also that negative pressures are experienced at certain instants of the wave cycle. At $\omega t=\pi$ and $\omega t=\frac{3 \pi}{2}$ the wave pressures are the negative of those illustrated for $\omega t=0$ and $\omega t=\frac{\pi}{2}$ respectively.

Thus we would expect, for example, positive wave-induced pore pressures at $\omega t=\frac{\pi}{2}$ and negative wave-induced pore pressures at $\omega t=\frac{3 \pi}{2}$. We have also investigated the wave pressures on a 12 degree slope in shallower water. A wave of length 150 m in 80 m of water would have a length of 133.5 m in 30 m of water over a flat bed. Figure (8.3) illustrates the wave pressures on a 12 degree slope, the wave length and water depth at the base of the slope being 133.5 m and 30 m respectively. The incident wave height is set at 16 m . Even at this reduced wave height, the maximum wave pressures are more than twice those illustrated in Figure (8.2), owing to the smaller water depths.

We now examine the wave-induced failure zones in the soil. As described in Section (7.3), the approximate extent of the failure zone may be determined by computing the stress angle $\theta$ at several points in the soil and comparing with the angle of friction $\phi^{\prime}$. For this purpose it is convenient to plot contours of the stress angle. If the angle of friction of the soil is, for example, $\phi^{\prime}=30^{\circ}$, we can determine the failure zone approximately as the region in which $\theta \geq 30^{\circ}$. We emphasize again that this provides only an estimate of the minimum failure zone, owing to the fact that violation of the Mohr-Coulomb criterion invalidates the stress field. The meaning of "failure" in the Mohr-Coulomb sense is that frictional resistance in the soil has been exceeded on some plane. The failure zone is thus an unstable region and further research is needed to study the manner in which failure progresses. Stress angle contours for a fine sand and coarse sand with slope angles of $\beta=$ $2^{\circ}, 5^{\circ}, 12^{\circ}, 20^{\circ}$ and wave lengths of $L=150 \mathrm{~m}, 225 \mathrm{~m}, 300 \mathrm{~m}$ are given in this chapter and in Appendix J. The contours were produced by the graphics package SURFACE II at Memorial University. Because of the interpolation and smoothing performed


Figure 8.1: Wave Pressure on 12 deg. Slope : $L=300 \mathrm{~m} ., H=24 \mathrm{~m}$.
Water Depth at Slope Base $(x=0)=80 \mathrm{~m}$.


Figure 8.2: Wave Pressure on 12 deg. Slope : $L=150 \mathrm{~m} ., H=24 \mathrm{~m}$.
Water Depth at Slope Base $(x=0)=80 \mathrm{~m}$.


Figure 8.3: Wave Pressure on 12 deg. Slope : $L=133.5 \mathrm{~m} ., H=16 \mathrm{~m}$.
Water Depth at Slope Base $(x=0)=30 \mathrm{~m}$.
by SURFACE II the contours are to be regarded as an approximate representation of the actual numerical results. In all cases (except figure 8.11) a wave height of $24 m$ was used. The conjugate stress ratio $K$ was taken to be 0.5 except in figures (8.13) to (8.15). The distances $d, d_{0}, d_{0}$ in Figure (7.1) are $65 m, 5 m$ and $5 m$ respectively. The water depth at $C$ (Figure 7.1) was kept at 80 m and the depth at $B$ varied from $78 m$ for the $2^{\circ}$ slope to 60 m for the $20^{\circ}$ slope. Similarly, the soil depth $l_{0}$ at $C$ was kept at 20 m , with the depth $l_{0}^{\prime}$ at $B$ varying from $22 m$ for the $2^{\circ}$ slope to 40 m for the $20^{\circ}$ slope. In only two cases was it necessary to deviate from this pattern, viz. the cases of fine sand, $\beta=20^{\circ}$ with $L=225 m$ and $150 m$. In these cases we chose $l_{0}=17 m, l_{0}^{\prime}=37 m$ and $l_{0}=14 m, l_{0}^{\prime}=34 m$ respectively. The reason for these changes is that numerical overfow problems were encountered when generating the boundary conditions on $B R$ (using the method of Appendix I). The effect of the changes is that the boundary conditions on $R S$ (Figure 7.1) is applied at a smaller depth than would otherwise have been the case. In order to investigate this further we have tested the case of Figure (8.4) (fine sand, $L=300 m, \beta=20^{\circ}$ ) with the location of the boundary $R S$ determined by $l_{0}=14 \mathrm{~m}$ and $l_{0}^{\prime}=34 \mathrm{~m}$. The contours obtained (shown in Figure J.1) are of the same pattern as in Figure (8.4), although in some places the failure zone ( $\theta \geq 30^{\circ}$ ) is not as deep.

The soil data is as follows :

Shear modulus $G\left(\mathrm{Nm}^{-2}\right)$
Poisson's ratio $\nu$
Porosity $f$
Permeability $k_{0}\left(m s^{-1}\right)$
$1.0 \times 10^{7}$
0.33 0.3

$$
1.0 \times 10^{-2}(\text { coarse sand })
$$

Density of soil grains $\rho_{s}\left(\mathrm{kgm}^{-3}\right) \quad 2.7 \times 10^{3}$
The bulk modulus of the pore water, $K_{f}$, was taken as $2.3 \times 10^{9} \mathrm{Nm}^{-2}$.
As expacted, the plots indicate that for a given slope angle the longer wavelengths produce the greater failure zones, and for a given wavelength the greater failures are observed in the steeper slopes. In most cases, there is no significant difference in the extent of the failure zones (hereafter termed the failure profile) for fine sand ( $k_{0}=10^{-4} m s^{-1}$ ) and coarse sand ( $k_{0}=10^{-2} \mathrm{~ms}^{-1}$ ). We observe, however, for $\beta=20^{\circ}, L=300 \mathrm{~m}$ and 225 m the failure zones $\left(\theta \geq 30^{\circ}\right.$ ) in fine sand (Figures 8.4 and 8.5 ) are slightly deeper than in coarse sand (Figures ס̄.i and 8.7). Also, in these cases, failure conditions in fine sand persist at $\omega t=\pi$, whereas this is not so in coarse sand.

The plots also illustrate the variation in the failure profile throughout the course of a wave cycle. The contours are shown for the instants $\omega t=0, \frac{\pi}{2}, \pi, \frac{3 \pi}{2}$. As the wave cycle progresses in time the responses of the soil skeleton and pore water also go through a cycle, out of phase with the wave cycle. The phase lag is quite easily computed from the arguments of the complex effective stresses and pore pressures determined by the boundary element method. We emphasize that the skeleton and pore water do not respond independently : the responses are coupled as described


Figure 8.4: Stress Angle Contours(degrees): Fine Sand, $K=0.5, L=300 \mathrm{~m}, \beta=20^{\circ}$


Figure 8.5: Stress Angle Contours(degrees): Fine Sand, $h^{\prime}=0.5, L=225 \mathrm{~m}, \beta=20^{\circ}$

## Horizontal Distance (m.)



Horizontal Distance (m.)


## Horizontal Distance (m.)



## Horizontal Distance (m.)



Figure 8.6: Stress Angle Contours(degrees):Coarse Sand, $K=0.5, L=300 \mathrm{~m}, \beta=20^{\circ}$


Figure 8.7: Stress Angle Contours(degrees): ('oarse Sand, $K^{\prime}=0.5, L=225 \mathrm{~m}, \beta=20^{\circ}$
by the Biot equations. At certain instants in the wave cycle the effective stresses at several points are such that the Mohr circles touch or cross the failure criterion. At other instants the effective stresses produce Mohr circles which are quite distant from the failure criterion, this being the result of negative wave-induced pore pressures. It is therefore necessary to examine the stability characterisitcs of the slope throughout the wave cycle in order to assess the likelihood of failure. For the time instants plotted, the greatest failure zones are observed at $\omega t=\frac{\pi}{2}$ in most cases.

The stress angle contours in a flat bed of coarse sand (computed from the analytic solution), with $L=300 m, K=K_{0}=0.5$ are shown in Figure (8.8). If we compare this with the plots for slopes of $5^{\circ}$ and $2^{\circ}$ in coarse sand with $L=300 \mathrm{~m}$ (Figures 8.9 and 8.10) we notice that the failure profiles and failure depths are similar. This means that for gentle slopes of the order of $5^{\circ}$ or less we can determine approximately the extent of the failure zones by using the analytic solution for a flat seabed described in Appendix 1.

In Figure(8.11), we illustrate the stress angle contours in coarse sand corresponding to the wave loading described by Figure (8.3), i.e. a wave of length $133.5 m$ and $16 m$ height in 30 m of water incident on a $12^{\circ}$ slope. We have already noted that such a wave would produce more than twice the maximum pressures than it wrould in 80 m of water at a wavelength of 150 m , and a greater height of 24 m . The stress angle contours for this latter case are illustrated in Figure(8.12). Accordingly, the failure zones illustrated in Figure (8.11) are much greater than those illustrated in Figure (8.12).

We have also investigated the effect of the initial stress distribution as determined by the value of the conjugate stress ratio $K$. Stress angle contours in coarse

Table 8.1: In Situ Stress Angle $\boldsymbol{\theta}_{0}$ (deg)

| $\beta$ (deg) | $K=0.5$ | $K=0.7$ | $K=1.0$ |
| :---: | ---: | ---: | ---: |
| 2 | 19.6 | 10.4 | 2.0 |
| 5 | 20.1 | 11.3 | 5.0 |
| 12 | 22.7 | 15.7 | 12.0 |
| 20 | 27.6 | 22.3 | 20.0 |

sand at $\omega t=\frac{\pi}{2}(L=300 m)$ are shown in Figures (8.13) and (8.14) for $K=0.7$ and 1.0. We find that at these values of $K$ the extent of the failure zones is considerably smaller than at $K=0.5$ (compare, for example, Figures (8.13),(8.14), $\beta=20^{\circ}$ with Figure (8.6), $\omega t=\frac{\pi}{2}$ ). The reason for this becomes obvious when we examine the stress angle $\theta_{0}$ under in-situ stress conditions. This is easily computed from equations (7.9) and (7.17) as

$$
\begin{equation*}
\theta_{0}=\sin ^{-1}\left[\frac{\left(K^{2}-2 K \cos 2 \beta+1\right)^{\frac{1}{2}}}{K+1}\right] \tag{8.1}
\end{equation*}
$$

For normally consolidated soils the in-situ horizontal effective stress is less than the vertical effective stress, i.e. $\sigma_{x}^{\prime(0)}<\sigma_{z}^{\prime}(0)$. From (7.9) we find that

$$
\begin{equation*}
\sigma_{x}^{\prime(0)}<\sigma_{z}^{\prime(0)} \Longrightarrow K<\frac{1}{\cos 2 \beta} \tag{8.2}
\end{equation*}
$$

Values of $\theta_{0}$ for different values of $K$ and $\beta$ satisfying ( 8.2) are given in table (8.1). During the parts of the wave cycle in which the horizontal effective stress is decreased and the vertical effective stress is increased, the stress angle $\theta$ increases above its in-situ value. If it increases to the extent that $\theta \geq \phi^{\prime}$, failure occurs in a manner analogous to "active" failure, e.g. at $\omega t=\frac{\pi}{2}$. We note from table (8.1) that the in-situ stress angle $\theta_{0}$ decreases as $K$ goes from 0.5 to 1.0 . This means that
at $K=0.5$ the unloaded slope is closer to wave induced failure than at $K=1.0$. In some cases, e.g. $L=150 \mathrm{~m}, \beta=12^{\circ}, K=0.7$ and 1.0 , there is only minimal failure (Figures J. 21 and J.22).

In some soils the in-situ horizontal effective stress is larger than the vertical effective stress i.e. $\sigma_{x}^{\prime(0)}>\sigma_{x}^{\prime(0)}$. S'ıch a situation may occur if the soil has been subjected in the past to heavy overburden stresses which have since been removed over the course of its history. In this case the stress angle is increased during the parts of the wave cycle in which the horizontal effective stresses are increased and the vertical effective stresses reduced. Again if $\theta \geq \phi^{\prime}$ failure occurs, but this tirne in a manner analogous to "passive" failure. Such a situation is shown in Figure (8.15) for a coarse sand, $L=300 m, \beta=20^{\circ}$ at $\omega t=\pi$, where we have taken $K=2$. As an example we examine the stress state at the point with horizontal and vertical coordinates of $(-54.5,-7.6)$ in Figure (8.15). The in-situ stresses are (for $K=2$ )

$$
\begin{aligned}
\sigma_{x}^{\prime(0)} & =0.1155 \times 10^{6} \mathrm{Nm}^{-2} \\
\sigma_{x}^{\prime(0)} & =0.8064 \times 10^{5} \mathrm{Nm}^{-2} \\
\tau_{x z}^{\prime(0)} & =0.4199 \times 10^{5} \mathrm{Nm}^{-2}
\end{aligned}
$$

We note that $\sigma_{x}^{\prime(0)}>\sigma_{x}^{\prime(0)}$. The in-situ stress angle $\theta_{0}$ is $27.6^{\circ}$. The wave induced stresses are shown in table (8.2) at several instants in the wave cycle. At $\omega t=0$ and $\frac{\pi}{2}$ the wave reduces the horizontal effective stresses and increases the vertical effective stresses thus tending to reduce the stress angle from its in-situ value. The resultant stress angles are in fact $21.7^{\circ}$ and $25.1^{\circ}$ at $\omega t=0$ and $\frac{\pi}{2}$ respectively. At $\omega t=\pi$ and $\frac{3 \pi}{2}$ the wave pressures are reversed and the Mohr circle expands as the horizontal stress increases and the vertical stress decreases. The resultant

Table 8.2: Wave Induced Stresses $\left(\times 10^{4} \mathrm{Nm}^{-2}\right)$ at $(-54.5,-7.6)$ in Figure (8.15)

| Stress Component | $\omega t=0$ | $\omega t=\frac{\pi}{2}$ | $\omega t=\pi$ | $\omega t=\frac{3 \pi}{2}$ |
| :---: | ---: | ---: | ---: | ---: |
| $\sigma_{x}^{\prime}$ | -0.9694 | -2.143 | 0.9694 | 2.143 |
| $\sigma_{z}^{\prime}$ | 0.5819 | 0.3287 | -0.5819 | -0.3287 |
| $\boldsymbol{\tau}_{x z}^{\prime \prime}$ | -0.7634 | -0.4510 | 0.7634 | 0.4510 |

stress angles are $33.8^{\circ}$ and $31.0^{\circ}$ respectively, indicating that the failure criterion ( $\theta \geq 30^{\circ}$ ) has been violated. The failure zone at $\omega t=\pi$ is particularly extensive as shown in Figure (8.15).

Horizontal Distance (m.)
Seabed


Horizontal Distance (m.)
Seabed


## Horizontal Distance (m.).

## Seabed



Horizontal Distance (m.)
Seabed


Figure 8.8: Stress Angle Contours(degrees),Analytic Solution for Flat Bed :Coarse Sand, $K=0.5, L=300 \mathrm{~m}$

Horizontal Distance (m.)


## Horizontal Distance (m.)



Horizontal Distance (m.)


## Horizontal Distance (m.)



Figure 8.9: Stress Angle Contours(degrees): Coarse Sand, $K=0.5, L=300 \mathrm{~m}, \beta=5^{\circ}$

Horizontal Distance (m.)


Horizontal Distance (m.)


Horizontal Distance (m.)


Horizontal Distance (m.)


Figure 8.10: Stress Angle Contours(degrees): Coarse Sand, $K=0.5, L=300 \mathrm{~m}, \beta=2^{\circ}$

Horizontal Distance (m.)



Horizontal Distance (m.)


## Horizontal Distance (m.)



Figure 8.11: Stress Angle Contours(degrees): Coarse Sand, $K=0.5, L=133.5 \mathrm{~m}, \beta=12^{\circ}$

$$
h_{0}=30 \mathrm{~m}, h_{0}^{\prime}=18 \mathrm{~m}, \eta_{0}=8 \mathrm{~m}
$$

Horizontal Distance (m.)


Horizontal Distance (m.)


Horizontal Distance (m.)


Figure 8.12: Stress Angle Contours(degrees): Coarse Sand, $K=0.5, L=150 \mathrm{~m}, \beta=12^{\circ}$

$$
h_{0}=80 \mathrm{~m}, h_{0}^{\prime}=68 \mathrm{~m}, \eta_{0}=12 \mathrm{~m}
$$



Figure 8.13: Stress Angle Contours(degrees): Coarse Sand, $L=300 \mathrm{~m}, \omega t=\frac{\pi}{2}, K=0.7$

Horizontal Distance (m.)


Horizontal Distance (m.)


Horizontal Distance (m.)


Horizontal Distance (m.)


Figure 8.14: Stress Angle Contours(degrees): Coarse Sand, $L=300 \mathrm{~m}, \omega t=\frac{\pi}{2}, K=1.0$

Horizontal Distance (m.)


Horizontal Distance (m.)


Horizontal Distance (m.)


Horizontal Distance (m.)


Figure 8.15: Stress Angle Contours(degrees): Coarse Sand, $L=300 \mathrm{~m}, \beta=20^{\circ}, K=2.0$

## Chapter 9

## Summary and Conclusions

In this thesis we have accomplished the following:

1. Implementation of a boundary element procedure for determining the forces due to uni-directional waves on a plane seabed of arbitrary inclination within the context of linear (Airy) wave theory.
2. Implementation of a boundary element procedure for determining the waveinduced effective stresses and pore pressures in a sloping poroelastic seabed using the results from (1) above.
3. Determination of the approximate extent of the minimum failure zones in various slopes under different soil and wave conditions.

In addition, we have provided a derivation of the boundary integral equations and fundamental solutions for Biot's linear theory of poroelasticity under quasi-static sinusoidal loading conditions. This derivatior has not previously been recorded in the literature.

The following conclusions may be drawn :
(a). The boundary element technique is very accurate, as evidenced by the comparisons made with analytic solutions for flat beds.
(b). As expected, for a given slope angle the longer wave lengths produce the greater failure zones and for a given wavelength, the greater failures are observed in the steeper slopes. Also, a wave that is relatively harmless in deep water can cause significant failure in shallower water.
(c). There is no significant difference in the extent of the failure zones for fine sand ( $k_{0}=10^{-4} \mathrm{~ms}^{-1}$ ) and coarse sand ( $k_{0}=10^{-2} \mathrm{~ms}^{-1}$ ) although in some cases the failure zone in fine sand is slightly deeper.
(d). It is necessary to evaluate the stability characteristics of the slope throughout the wave cycle in order to assess the likelihood of failure.
(e). For gentler slopes, of the order of $5^{\circ}$ or less, the failure zones may be located with sufficient accuracy by using the analytic solution of Appendix I, i.e. we can avoid the complexity of the BEM for such cases.
(f). The failure profile is strongly dependent on the initial state of atress in the slope as determined by the conjugate stress ratio $K$. For normally consolidated soils, the failure zone decreases in sizes as $K$ increases, with $K$ remaining in the range for which $\sigma_{x}^{\prime}(0)<\sigma_{x}^{\prime}(0)$. In such soils the maximum wave-induced failure occurs in a manner analogous to active failure. For some overconsolidated soils we may have $\sigma_{x}^{\prime(0)}>\sigma_{x}^{\prime(0)}$ and in such cases the
maximum wave induced failure occurs in a manner analogous to passive failure. In these cases, larger $K$ values mean that the in-situ soil condition is closer to passive failure.

## Suggestions for Further Research

(1) In order to account for the volume changes of soil under cyclic loading modifications to the original Biot formulation are required, as suggested by Verruijt (1985). Perhaps the phenomenon of pore pressure buildup under cyclic loading may be analysed in this way. Further research is also required to model the non-linear, time dependent and anisotropic behaviour of soil.
(2) The propagation of the soil failure illustrated in this thesis needs to be studied by means of some form of elasto-plastic analysis.
(3) Wave-induced stresses of layered soils can be easily studied by the methods described in the thesis. It will be necessary to divide the domain into subdomains, generate a system of equations for each sub-domain, and match boundary conditions at common boundaries.
(4) Earthquake-induced stresses in a poroelastic medium can be determined by including the acceleration terms in the governing equations. This would involve the derivation of a new boundary integral equation and fundamental solutions.
(5) The study of axisymmetric and three-dimensional problems is a fruitful extension of the methods presented herein.

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## Appendix A

## Fundamentals of Linear Wave

## Theory

Referring to Figure (4.1), we assume inviscid and hence irrotational flow above the seabed. Therefore ther zexists a velocity potential $\Phi(x, y, z, t)$ such that

$$
\begin{equation*}
\nabla^{2} \Phi=0 \tag{A.1}
\end{equation*}
$$

In general, $\nabla^{2} \equiv \frac{\partial^{2}}{\partial x^{2}}+\frac{\partial^{2}}{\partial y^{2}}+\frac{\partial^{2}}{\partial z^{2}}$. The $z$ axis (not shown) is normal to the $x-y$ plane. We assume negligible flow into the seabed, i.e.

$$
\begin{equation*}
\frac{\partial \Phi}{\partial n}=0 \quad \text { on } B C \tag{A.2}
\end{equation*}
$$

We now determine the boundary condition at the water surface. The water surface may be represented by the equation

$$
\begin{equation*}
\zeta(x, y, z, t)=0 \tag{A.3}
\end{equation*}
$$

from which we have

$$
\begin{equation*}
\frac{d \zeta}{d t} \equiv \frac{\partial \zeta}{\partial x} \dot{x}+\frac{\partial \zeta}{\partial y} \dot{y}+\frac{\partial \zeta}{\partial z} \dot{z}+\frac{\partial \zeta}{\partial t}=0 \tag{A.4}
\end{equation*}
$$

on the water surface, where the dots denote differentation with respect to time $t$. This equation may be written

$$
\begin{equation*}
\nabla \zeta \nabla \Phi=-\frac{\partial \zeta}{\partial t} \tag{A.5}
\end{equation*}
$$

on the water surface, since

$$
\nabla \Phi=\left(\begin{array}{c}
\dot{x}  \tag{A.6}\\
\dot{y} \\
\dot{z}
\end{array}\right)
$$

If $\eta(x, z, t)$ denotes the elevation of the water surface above the mean water level $O A$, then the equation of the water surface is given by $y=\eta(x, z, t)$ or

$$
\begin{equation*}
\eta(x, z, t)-y=0 \tag{A.7}
\end{equation*}
$$

From (A.3) and (A.7):

$$
\begin{equation*}
\zeta(x, y, z, t)=r(x, z, t)-y \tag{A.8}
\end{equation*}
$$

from which we get

$$
\begin{align*}
& \frac{\partial \zeta}{\partial t}=\frac{\partial \eta}{\partial t}  \tag{A.9}\\
& \underline{Z} \zeta=\frac{\partial \eta}{\partial x} i+(-\underline{j})+\frac{\partial \eta}{\partial z} k \tag{A.10}
\end{align*}
$$

Here $\dot{i}, \dot{j}, k$ are unit vectors in the $x, y, z$ directions respectively. Substituting (A.9) and (A.10) into (A.5) gives

$$
\begin{equation*}
\frac{\partial \eta}{\partial x} \frac{\partial \Phi}{\partial x}-\frac{\partial \Phi}{\partial y}+\frac{\partial \eta}{\partial z} \frac{\partial \Phi}{\partial z}=-\frac{\partial \eta}{\partial t} \tag{A.11}
\end{equation*}
$$

on the water surface.
In linear wave theory we assume that the water surface elevation $\eta$ is small compared to the wave length so that equation (A.11) may be applied at the mean water level $y=0$. Further, we neglect the non-linear terms in (A.11) which thus becomes

$$
\begin{equation*}
\frac{\partial \Phi}{\partial y}=\frac{\partial \eta}{\partial t} \quad \text { at } y=0 \tag{A.12}
\end{equation*}
$$

Equation (A.12) is known as the linearized kinematic free surface condition. Another condition at the free surface may be deduced from the Bernoulli equation, which is

$$
\begin{equation*}
\frac{\partial \Phi}{\partial t}+\frac{1}{2} \nabla \Phi \nabla \Phi+\frac{p}{\rho_{f}}+g y=f(t) \tag{A.13}
\end{equation*}
$$

where $p$ is water pressure and $\rho_{f}$ is water density; $f(t)$ is an arbitrary function of time $t$. If we consider a purely hydrostatic condition, $\Phi \equiv 0, p=-\rho_{f} g y$ for all time, so that $f(t)$ is indentically zero.

We write equation (A.13) at the water surface and neglect the non-linear term:

$$
\begin{equation*}
\frac{\partial \Phi}{\partial t}+g \eta=0 \quad \text { at } y=\eta \text { i.e. at } y=0 \text { approximately } \tag{A.14}
\end{equation*}
$$

Equation (A.14) is called the linearized dynamic free surface condition.
The linear wave theory model is described by equation (A.1) with boundary conditions (A.2), (A.12) and (A.14). Equations (A.12) and (A.14) may be combined by eliminating $\eta$ :

$$
\begin{equation*}
\frac{\partial^{2} \Phi}{\partial t^{2}}+g \frac{\partial \Phi}{\partial y}=0 \quad \text { at } y=0 \tag{A.15}
\end{equation*}
$$

To derive the velocity potential due to the incident wave, as given in equation (4.14), we consider a flat seabed, constant water depth $h_{0}$, and a wave of surface
elevation

$$
\begin{equation*}
\eta=\eta_{0} e^{i(k x-\omega t)} \tag{A.16}
\end{equation*}
$$

This represents a wave propagating in the positive or negative $x$ direction according as $k>0$ or $k<0$. For the two dimensional problem, $\Phi$ is independent of $z$ and we write

$$
\begin{equation*}
\Phi(x, y)=g(y) e^{i(k x-\omega t)} \tag{A.17}
\end{equation*}
$$

Substituting in (A.1) gives

$$
\frac{\mathrm{d}^{2} g}{\mathrm{~d} y^{2}}-k^{2} g=0
$$

from which we have

$$
\begin{equation*}
g(y)=C \cosh (k y+\alpha) \tag{A.18}
\end{equation*}
$$

Applying (A.2) on the seabed $y=-h_{0}$ implies that $\alpha=k h_{0}$, so that

$$
\begin{equation*}
\Phi(x, y)=C \cosh \left(k y+k h_{0}\right) e^{i(k x-\omega t)} \tag{A.19}
\end{equation*}
$$

Now, using (A.14) we find that

$$
\begin{equation*}
C=-\frac{i g \eta_{0}}{\omega \cosh \left(k h_{0}\right)} \tag{A.20}
\end{equation*}
$$

and hence

$$
\begin{equation*}
\left.\Phi(x, y)=\frac{-i g \eta_{0}}{\omega \cosh \left(k h_{0}\right)} \cosh \left(k y+k h_{0}\right) e^{i(k x-\omega t}\right) \tag{A.21}
\end{equation*}
$$

A dispersion relation is obtained by substituting (A.19) into (A.15), which gives:

$$
\begin{equation*}
\omega^{2}=g k \tanh \left(k h_{0}\right) \tag{A.22}
\end{equation*}
$$

For an incident wave propagating in the negative $x$ direction we set $k=-k_{0}$, the negative real root of (A.22). This gives the incident velocity potential as

$$
\begin{equation*}
\Phi_{I}(x, y)=\frac{-i g \eta_{0}}{\omega \cosh \left(k_{0} h_{0}\right)} \cosh \left(k_{0} y+k_{0} h_{0}\right) e^{-i\left(k_{0} x+\omega t\right)} \tag{A.23}
\end{equation*}
$$

## Appendix B

## The Roots of the Dispersion

## Relation (Linear Wave Theory)

The dispersion relation (4.12) is

$$
\begin{equation*}
\frac{\omega^{2}}{g}=k \tanh \left(k h_{0}\right) \tag{B.1}
\end{equation*}
$$

This is of the form

$$
\begin{equation*}
\frac{a}{x}=\tanh x \tag{B.2}
\end{equation*}
$$

where $x=k h_{0}$ and $a=\frac{w^{2} h_{0}}{g}>0$. The sketch graph Figure (B.1) shows that equation ( B .2 ) has two real roots $x= \pm x_{0}$. Thus equation (B.1) has two real roots $\pm k_{0} \equiv \pm \frac{x_{0}}{h_{0}}$. To obtain the imaginary roots of (B.1) we put $k=i \vec{k}$ where $\bar{k}$ is real. Then (B.1) becomes

$$
\begin{equation*}
-\frac{a}{x}=\tan x \tag{B.3}
\end{equation*}
$$

where $x=\bar{k} h_{0}$, and $a=\frac{w^{2} h_{0}}{g}$.


Fig.(B.1) Real roots of Dispersion Relation


Fig.(B.2) Imaginary Roots of Dispersion Relation

The sketch graph Figure (B.2) shows that (B.3) has an infinity of roots $\pm x_{1}, \pm x_{2}, \ldots$ so that

$$
\begin{equation*}
\bar{k}= \pm \frac{x_{1}}{h_{0}}, \pm \frac{x_{2}}{h_{0}}, \ldots \tag{B.4}
\end{equation*}
$$

i.e. the imaginary roots of (B.1) are given by $\pm i \boldsymbol{k}_{\boldsymbol{i}}$ where

$$
\begin{equation*}
k_{i}=\frac{x_{i}}{h_{0}} \tag{B.5}
\end{equation*}
$$

To obtain accurate values of the roots of (B.3), it is necessary to provide good estimates of these roots. From Figure (B.2) we note that the nth positive root of (B.3) is of the form

$$
\begin{equation*}
x_{n}=n \pi-y \tag{B.6}
\end{equation*}
$$

where $y>0$ and depends on $n$. Substituting (B.6) into (B.3) gives

$$
\begin{equation*}
\tan y-\frac{a}{n \pi-y}=0 \tag{B.7}
\end{equation*}
$$

The value of $y$ is small and becomes smaller as $n$ increases, so we may make the approximation $\tan y \approx y$. This renders equation (B.7) as

$$
y^{2}-n \pi y+a=0
$$

which has two positive roots:

$$
y=\frac{\pi}{2}\left[n \pm \sqrt{n^{2}-\frac{4 a}{\pi^{2}}}\right] \quad \text { for } n>\frac{2}{\pi} \sqrt{a}
$$

The root of interest is the smaller value, i.e.

$$
\begin{equation*}
y=\frac{\pi}{2}\left[n-\sqrt{n^{2}-\frac{4 a}{\pi^{2}}}\right] \quad \text { for } n>\frac{2}{\pi} \sqrt{a} \tag{B.8}
\end{equation*}
$$

We therefore solve (B.7) for values of $n>\frac{2}{\pi} \sqrt{a}$, each time providing an estimate of $y$ from (B.8). The $n$th root of (B.3) is then given by (B.6). For values of
$n<\frac{2}{\pi} \sqrt{a}$ we obtain the roots directly from (B.3) using $x_{n}=n \pi$ as an estimate.
The imaginary roots of(B.1) are then $\pm \frac{i x_{n}}{h_{0}}, n=1,2, \cdots$. The equations (B.2),
( B.3), ( B.7), may be solved using the IMSL routine ZREAL.

## Appendix C

## Evaluation of the Integrals

## (Potential Problem)

We refer to equations (4.33). We consider a typical element $\Gamma_{e}$ as shown in Figure (C.1) with end points $\left(x_{e}, y_{e}\right),\left(x_{e+1}, y_{e+1}\right)$. The direction of integration and the unit outward normal $\underline{n}$ are indicated in the figure. The direction of istegration is such that the interior of the problem domain lies to the left. The element $\Gamma_{e}$ is represented in terms of parameter $\boldsymbol{\xi}$ as follows:

On $\Gamma_{e}$

$$
\begin{array}{r}
x=\frac{1}{2}(1-\xi) x_{e}+\frac{1}{2}(1+\xi) x_{e+1} \\
y=\frac{1}{2}(1-\xi) y_{e}+\frac{1}{2}(1+\xi) y_{e+1} \\
\text { where }-1 \leq \xi \leq 1 \tag{C.1}
\end{array}
$$



Fig.(C.1) Element $\Gamma_{\mathrm{e}}$ (Wave problem)

The line element $d s$ is given by

$$
\left(\frac{d s}{d \xi}\right)^{2}=\left(\frac{d x}{d \xi}\right)^{2}+\left(\frac{d y}{d \xi}\right)^{2}
$$

from which we obtain

$$
\begin{equation*}
d \Gamma_{e}=d s=J_{e} d \xi \tag{C.2}
\end{equation*}
$$

where

$$
\begin{equation*}
J_{e}=\frac{1}{2}\left[\left(x_{e+1}-x_{e}\right)^{2}+\left(y_{e+1}-y_{e}\right)^{2}\right]^{\frac{1}{2}} \tag{C.3}
\end{equation*}
$$

We have defined the distance $r$ as $r=\left|\underline{x}-\underline{x}_{j}\right|$ where $\underline{x}$ is an arbitrary point on $\Gamma_{e}$ and $x_{j}$ is the position vector of node $j$. Since node $j$ lies on the mid-point of element $\Gamma_{j}$ we have the coordinates ( $x_{0}, y_{0}$ ) of node $j$ as:

$$
\begin{aligned}
x_{0} & =\frac{1}{2}\left(x_{j}+x_{j+1}\right) \\
y_{0} & =\frac{1}{2}\left(y_{j}+y_{j+1}\right)
\end{aligned}
$$

Thus

$$
\begin{align*}
r=\left[\left(x-x_{0}\right)^{2}+\left(y-y_{0}\right)^{2}\right]^{\frac{1}{2}} & \text { for } x, y \in \Gamma_{e} \\
\frac{\partial r}{\partial x}=\frac{x-x_{0}}{r} ; & \frac{\partial r}{\partial y}=\frac{y-y_{0}}{r} \tag{C.4}
\end{align*}
$$

The unit normal $n$ is given by

$$
\begin{array}{r}
\underline{n}=\binom{n_{1}}{n_{2}} \text { where } \\
n_{1}=\frac{1}{2 J_{e}}\left(y_{e+1}-y_{e}\right) ; \quad n_{2}=\frac{1}{2 J_{e}}\left(x_{e}-x_{e+1}\right) \tag{C.5}
\end{array}
$$

If $e \neq j$, integration along $\Gamma_{e}$ presents no difficulty. An integral of the form $\int_{\Gamma_{e}} f\left(r_{i} r_{i,}, n_{i}\right) d \Gamma_{e}$ is transformed into one of the form $\int_{-1}^{1} g(\xi) d \xi$ by use of the formulae (C.1) to (C.5) and then straight-forward Gaussian integration is performed.

Integration along $\Gamma_{j}$, however, requires special consideration owing to the logarithmic singularity of the integrand. Putting $e=j$ in (C.1) to (C.5) we find that on $\Gamma_{j}$,

$$
\begin{array}{r}
r=|\xi| J_{j} \\
\frac{\partial r}{\partial x}=\left(\frac{x_{j+1}-x_{j}}{2 J_{j}}\right) \cdot \frac{\xi}{|\xi|} ; \frac{\partial r}{\partial y}=\left(\frac{y_{j+1}-y_{j}}{2 J_{j}}\right) \frac{\xi}{|\xi|} \\
n_{1}=\frac{y_{j+1}-y_{j}}{2 J_{j}} ; n_{2}=\frac{x_{j}-x_{j+1}}{2 J_{j}} \tag{C.8}
\end{array}
$$

Equations ( C.7), ( C.8) imply that

$$
\begin{equation*}
\frac{\partial r}{\partial n}=0 \quad \text { on } \Gamma_{j} \tag{C.9}
\end{equation*}
$$

We consider the integral

$$
\begin{align*}
\int_{\Gamma_{j}} \ln r d \Gamma_{j} & =\int_{-1}^{1} \ln \left[|\xi| J_{j}\right] J_{j} d \xi \\
& =J_{j} \ln J_{j} \int_{-1}^{1} d \xi+J_{j} \int_{-1}^{1} \ln |\xi| d \xi \\
& =2 J_{j} \ln J_{j}+2 J_{j} \int_{0}^{1} \ln \xi d \xi \\
& =2 J_{j}\left(\ln J_{j}-1\right) \tag{C.10}
\end{align*}
$$

We can now evaluate the integrals $A_{e}$ and $B_{e}$. From (4.24)

$$
\begin{aligned}
\phi^{*} & =-\frac{1}{2 \pi} \ln r \\
\Rightarrow \frac{\partial \phi^{*}}{\partial n} & =-\frac{1}{2 \pi r} \frac{\partial r}{\partial n}
\end{aligned}
$$

From (4.33):

$$
A_{e}= \begin{cases}-\frac{1}{2 \pi} \int_{\Gamma_{e}} \frac{1}{r} \frac{\partial r}{\partial n} d \Gamma_{e} & e \neq j \\ -\frac{1}{2 \pi} \int_{\Gamma_{j}} \frac{1}{r} \frac{\partial r}{\partial n} d \Gamma_{j}+\frac{1}{2} & e=j\end{cases}
$$

Using (C.9), this becomes

$$
A_{e}= \begin{cases}-\frac{1}{2 \pi} \int_{\Gamma_{e}} \frac{1}{r} \frac{\partial r}{\partial n} d \Gamma_{e} & e \neq j  \tag{C.11}\\ \frac{1}{2} & e=j\end{cases}
$$

Also, from (4.33) and ( C.10)

$$
B_{e}= \begin{cases}-\frac{1}{2 \pi} \int_{\Gamma_{e}} \ln r d \Gamma_{e} & e \neq j  \tag{C.12}\\ \frac{J_{j}}{\pi}\left(1-\ln J_{j}\right) & e=j\end{cases}
$$

As stated before, the integrals for $e \neq j$ can be simply evaluated using standard Gaussian quadrature.

## Appendix D

## Fundamental Solution (Laplace's

## Equation)

The following derivation is a classical one and can be found in any standard text on the boundary element method, e.g. Brebbia (1984). Referring to equations (4.23) or (5.34) we need to consider an equation of the form

$$
\begin{equation*}
\nabla^{2} \phi+k \Delta\left(\underline{x}, \underline{x}_{p}\right)=0 \tag{D.1}
\end{equation*}
$$

where $k$ is a constant.
We recall the following familiar property of the delta function:

$$
\int_{\Omega} \Delta\left(\underline{x}-x_{p}\right) d \Omega= \begin{cases}1 & x_{p} \in \Omega  \tag{D.2}\\ 0 & x_{p} \in \Omega\end{cases}
$$

where the integration is performed with respect to $x$. We integrate (D.1) over a
circular region $\Omega$, centre $x_{p}$ and radius $\varepsilon$ :

$$
\begin{equation*}
\int_{\Omega} \phi_{, i i} d \Omega+k=0 \tag{D.3}
\end{equation*}
$$

where we have used (D.2). By the divergence theorem,

$$
\begin{equation*}
\int_{\Omega} \phi_{, i i} d \Omega=\int_{\Gamma} \frac{\partial \phi}{\partial n} d \Gamma \tag{D.4}
\end{equation*}
$$

where $\Gamma$ is the boundary of $\Omega$. From (D.3) and (D.4):

$$
\begin{equation*}
\int_{\Gamma} \frac{\partial \phi}{\partial n} d \Gamma=-k \tag{D.5}
\end{equation*}
$$

We define polar coordinates with $x_{p}$ as origin: $r=\left|x-x_{p}\right| ; \theta$ is measured anticlockwise from the $x$-axis through $x_{p}$. On $\Gamma, r=\varepsilon$, and $\frac{\partial \phi}{\partial n}=\frac{\partial \phi}{\partial r}$ so that (D.5) becomes

$$
\begin{equation*}
\int_{0}^{2 \pi}\left(\frac{\partial \phi}{\partial r}\right)_{r=e} \varepsilon d \theta=-k \tag{D.6}
\end{equation*}
$$

Assuming that $\phi$ is independent of $\theta$, we have

$$
\begin{equation*}
\left(\frac{\partial \phi}{\partial r}\right)_{r=e}=\frac{-k}{2 \pi \varepsilon} \tag{D.7}
\end{equation*}
$$

which is satisfied by:

$$
\begin{equation*}
\phi=-\frac{k}{2 \pi} \ln r \tag{D.8}
\end{equation*}
$$

## Appendix E

## A. Solution of Equation (5.108)

The equation is

$$
\begin{equation*}
\nabla^{2} \eta_{i}-a^{2} \eta_{i}=0 \tag{E.1}
\end{equation*}
$$

where $a$ is a real or complex constant. We verify that a solution of (E.1) is

$$
\begin{equation*}
\eta_{i}=r_{, i} K_{1}(a r) \tag{E.2}
\end{equation*}
$$

where $r=|\mathrm{x}-\mathrm{p}|$. We first note that for real or complex $z$, and integer $n$ :

$$
\begin{align*}
K_{n+1}(z) & =K_{n-1}(z)+\frac{2 n}{z} K_{n}(z)  \tag{E.3}\\
\frac{d}{d z}\left[K_{n}(z)\right] & =-\frac{1}{2}\left\lfloor K_{n-1}(z)+K_{n+1}(z)\right] \tag{E.4}
\end{align*}
$$

Also, since $r^{2}=\left(x_{i}-p_{i}\right)\left(x_{i}-p_{i}\right)$,

$$
\begin{align*}
r_{1 j} & =\frac{x_{j}-p_{j}}{r} \\
r_{, i j} & =\frac{\delta_{i j}}{\ddot{ }}-\frac{r_{, i} r_{, j}}{r} \tag{E.5}
\end{align*}
$$

$$
\begin{aligned}
r, j j & =\frac{1}{r} \\
r_{, j} r_{, j} & =1
\end{aligned}
$$

We have from (E.2)

$$
\begin{align*}
\eta_{i, j} & =r, i\left[-K_{0}(a r)-\frac{1}{a r} K_{1}(a r)\right] a r_{, j}+r_{, i j} K_{1}(a r) \\
& =-a r_{, i} r_{, j}\left[K_{0}(a r)+\frac{2}{a r} K_{1}(a r)\right]+\frac{K_{1}(a r)}{r} \delta_{i j} \\
& =-a r_{, i} r_{, j} K_{2}(a r)+\frac{K_{1}(a r)}{r} \delta_{i j} \tag{E.6}
\end{align*}
$$

Also, using ( E.5) it is easy to show that

$$
\begin{equation*}
\left(r_{, i} r_{, j, j}\right)_{, j}=\frac{r_{, i}}{r} \tag{E.7}
\end{equation*}
$$

Differentiating (E.6) with respect to $x_{j}$ gives

$$
\begin{aligned}
\eta_{i, j j}= & -a\left\{r_{, i} r_{, j}\left[-K_{1}(a r)-\frac{2}{a r} K_{2}(a r)\right] a r_{, j}+K_{2}(a r)\left(r_{, i} r_{, j}\right), j\right\} \\
& +\delta_{i j}\left\{\frac{1}{r}\left[-K_{0}(a r)-\frac{1}{a r} K_{1}(a r)\right] a r_{, j}-\frac{r, j}{r^{2}} K_{1}(a r)\right\} \\
= & a\left\{a r_{, i}\left[K_{1}(a r)+\frac{2}{a r} K_{2}(a r)\right]-\frac{r_{, i}}{r} K_{2}(a r)\right\} \\
& +\delta_{i j}\left\{-\frac{a}{r} r_{, j} K_{0}(a r)-\frac{2 r_{, j},}{r^{2}} K_{1}(a r)\right\} \\
= & a\left\{a r_{, i} K_{1}(a r)+\frac{2 r,, i}{r} K_{2}(a r)-\frac{r_{, i}}{r} K_{2}(a r)-\frac{r_{, i}}{r} K_{0}(a r)-\frac{2 r_{, i}}{a r^{2}} K_{1}(a r)\right\} \\
= & a\left\{a r_{, i} K_{1}(a r)+\frac{r_{i,}}{r}\left[K_{2}(a r)-K_{0}(a r)-\frac{2}{a r} K_{1}(a r)\right]\right\} \\
= & a^{2} r_{, i} K_{1}(a r)=a^{2} \eta_{i}
\end{aligned}
$$

i.e.

$$
\nabla^{2}\left[r, i K_{1}(a r)\right]=a^{2}\left[r_{, i} K_{1}(a r)\right]
$$

## Appendix F

## Functions Required for Kelvin

## Function Approximations

These functions are defined in equations (6.21) to (6.27) and (6.28).

$$
\begin{aligned}
& u_{0}(x)=\left\{\begin{array}{cc}
-\left[1+\alpha_{1}\left(\frac{x}{8}\right)^{4}+\alpha_{2}\left(\frac{x}{8}\right)^{8}+\cdots+\alpha_{7}\left(\frac{x}{8}\right)^{28}\right] & 0<x \leq 8 \\
0 & x>8
\end{array}\right. \\
& v_{0}(x)=\left\{\begin{array}{cc}
\frac{\pi}{4}\left(\frac{x}{8}\right)^{2}\left[\beta_{1}+\beta_{2}\left(\frac{x}{8}\right)^{4}+\beta_{3}\left(\frac{x}{8}\right)^{8}+\cdots+\beta_{7}\left(\frac{x}{8}\right)^{24}\right] & 0<x \leq 8 \\
+\left[a_{0}+a_{1}\left(\frac{x}{8}\right)^{4}+a_{2}\left(\frac{x}{8}\right)^{8}+\cdots+a_{7}\left(\frac{x}{8}\right)^{28}\right] & \\
\sqrt{\frac{\pi}{2 x}} e^{\nu(x)} \cos \lambda(x) & x>8
\end{array}\right.
\end{aligned}
$$

$$
\begin{aligned}
& u_{0}^{\prime}(x)=\left\{\begin{array}{cc}
\left(\frac{x}{8}\right)^{2}\left[\beta_{1}+\beta_{2}\left(\frac{x}{8}\right)^{4}+\beta_{3}\left(\frac{x}{8}\right)^{8}+\cdots+\beta_{7}\left(\frac{x}{8}\right)^{24}\right] & 0<x \leq 8 \\
0 & x>8
\end{array}\right. \\
& \int \frac{\pi}{4}\left[1+\alpha_{1}\left(\frac{x}{8}\right)^{4}+\alpha_{2}\left(\frac{x}{8}\right)^{8}+\cdots+\alpha_{7}\left(\frac{x}{8}\right)^{28}\right] \quad 0<x \leq 8 \\
& v_{0}^{\prime}(x)=\left\{-\left(\frac{x}{8}\right)^{2}\left[b_{1}+b_{2}\left(\frac{x}{8}\right)^{4}+b_{3}\left(\frac{x}{8}\right)^{8}+\cdots+b_{7}\left(\frac{x}{8}\right)^{24}\right]\right. \\
& -\sqrt{\frac{\pi}{2 x}} e^{\nu(x)} \sin \lambda(x) \quad x>8 \\
& u_{1}(x)=\left\{\begin{array}{cl}
\left(\frac{1}{2}\right)+\beta_{1}^{\prime}\left(\frac{x}{8}\right)^{4}+\beta_{2}^{\prime}\left(\frac{x}{8}\right)^{8}+\cdots+\beta_{6}^{\prime}\left(\frac{x}{8}\right)^{24} & 0<x \leq 8 \\
0 & x>8
\end{array}\right. \\
& v_{1}(x)=\left\{\begin{array}{cc}
\frac{1}{8^{2}}\left[\beta_{1}+\beta_{2}\left(\frac{x}{8}\right)^{4}+\beta_{3}\left(\frac{x}{8}\right)^{8}+\cdots+\beta_{7}\left(\frac{x}{8}\right)^{24}\right] & 0<x \leq 8 \\
+\frac{\pi}{4}\left[\alpha_{1}^{\prime}\left(\frac{x}{8}\right)^{2}+\alpha_{2}^{\prime}\left(\frac{x}{8}\right)^{6}+\cdots+\alpha_{7}^{\prime}\left(\frac{x}{8}\right)^{26}\right]-g_{0}^{\prime}(x) & \\
\frac{\sqrt{\pi}}{(2 x)^{\frac{3}{2}}} e^{\nu(x)}\left\{s_{\phi}(x)[\sin \lambda(x)+\cos \lambda(x)]+t_{\phi}(x)[\sin \lambda(x)-\cos \lambda(x)]\right\} & x>8
\end{array}\right.
\end{aligned}
$$

$$
\begin{gathered}
u_{1}^{\prime}(x)=\left\{\begin{array}{cl}
\alpha_{1}^{\prime}\left(\frac{x}{8}\right)^{2}+\alpha_{2}^{\prime}\left(\frac{x}{8}\right)^{0}+\cdots+\alpha_{7}^{\prime}\left(\frac{x}{8}\right)^{26} & 0<x \leq 8 \\
0 & x>8
\end{array}\right. \\
v_{1}^{\prime}(x)=\left\{\begin{array}{cc}
\frac{1}{8^{2}}\left[\alpha_{1}\left(\frac{x}{8}\right)^{2}+\alpha_{2}\left(\frac{x}{8}\right)^{6}+\cdots+\alpha_{7}\left(\frac{x}{8}\right)^{26}\right] \\
-\frac{\pi}{4}\left[\frac{1}{2}+\beta_{1}^{\prime}\left(\frac{x}{8}\right)^{4}+\beta_{2}^{\prime}\left(\frac{x}{8}\right)^{8}+\cdots+\beta_{6}^{\prime}\left(\frac{x}{8}\right)^{24}\right]-f_{0}^{\prime}(x) & 0<x \leq 8 \\
\frac{\sqrt{\pi}}{(2 x)^{\frac{3}{2}}} e^{\nu(x)}\left\{s_{\phi}(x)[\cos \lambda(x)-\sin \lambda(x)]+t_{\phi}(x)[\cos \lambda(x)+\sin \lambda(x)]\right\} & \\
-\frac{1}{x^{2}} & x>8
\end{array}\right.
\end{gathered}
$$

$$
\begin{gathered}
u_{2}(x)=u_{0}(x)+2 u_{1}(x) \\
u_{2}^{\prime}(x)=u_{0}^{\prime}(x)+2 u_{1}^{\prime}(x) \\
v_{2}(x)=v_{0}(x)+2 v_{1}(x) \\
v_{2}^{\prime}(x)=v_{0}^{\prime}(x)+2 v_{1}^{\prime}(x) \\
u_{3}(x)=\left\{\begin{array}{cc}
\frac{1}{8 \sqrt{2}}\left[\left(-\beta_{1}-2 \alpha_{1}^{\prime}\right)\left(\frac{x}{8}\right)+\left(-\alpha_{1}+2 \beta_{1}^{\prime}\right)\left(\frac{x}{8}\right)^{3}\right. \\
+\left(-\beta_{2}-2 \alpha_{2}^{\prime}\right)\left(\frac{x}{8}\right)^{5}+\left(-\alpha_{2}+2 \beta_{2}^{\prime}\right)\left(\frac{x}{8}\right)^{7} \\
\left.+\cdots+\left(-\alpha_{7}+2 \beta_{7}^{\prime}\right)\left(\frac{x}{8}\right)^{27}\right] & 0<x \leq 8 \\
0 & x>8
\end{array}\right.
\end{gathered}
$$

$$
\begin{aligned}
& w_{3}(x)=\left\{\begin{array}{c}
\frac{1}{8 \sqrt{2}}\left[\frac { \pi } { 4 } \left\{\left(\beta_{1}+2 \alpha_{1}^{\prime}\right)\left(\frac{x}{8}\right)+\left(-\alpha_{1}+2 \beta_{1}^{\prime}\right)\left(\frac{x}{8}\right)^{3}\right.\right. \\
\left(\beta_{2}+2 \alpha_{2}^{\prime}\right)\left(\frac{x}{8}\right)^{5}+\left(-\alpha_{2}+2 \beta_{2}^{\prime}\right)\left(\frac{x}{8}\right)^{7} \\
\left.+\cdots+\left(-\alpha_{7}+2 \beta_{7}^{\prime}\right)\left(\frac{x}{8}\right)^{27}\right\} \\
+\left(b_{1}+2 a_{1}^{\prime}\right)\left(\frac{x}{8}\right)^{2}+\left(a_{1}-2 b_{1}^{\prime}\right)\left(\frac{x}{8}\right)^{3} \\
+\left(b_{2}+2 a_{2}^{\prime}\right)\left(\frac{x}{8}\right)^{5}+\left(a_{2}-2 b_{2}^{\prime}\right)\left(\frac{x}{8}\right)^{7} \\
+\cdots+\left(a_{7}-2 b_{7}^{\prime}\right)\left(\frac{x}{8}\right)^{27} \\
+\frac{2}{8^{2}}\left\{-\alpha_{1}\left(\frac{x}{8}\right)+\beta_{2}\left(\frac{x}{8}\right)^{3}-\alpha_{2}\left(\frac{x}{8}\right)^{5}+\beta_{3}\left(\frac{x}{8}\right)^{7}\right. \\
\left.\left.+\cdots+\beta_{7}\left(\frac{x}{8}\right)^{23}-\alpha_{7}\left(\frac{x}{8}\right)^{25}\right\}\right] \\
+\quad 0<x \leq 8 \\
\frac{\sqrt{\pi}}{2 x^{\frac{3}{2}}} e^{\nu(x)}[\cos \lambda(x)+\sin \lambda(x)+ \\
\left.\frac{2}{x}\left\{s_{\phi}(x) \sin \lambda(x)-t_{\phi}(x) \cos \lambda(x)\right\}\right]+\frac{\sqrt{2}}{x^{3}}
\end{array} \quad x>8\right. \\
& z_{3}(x)=\left\{\begin{array}{cl}
\frac{1}{\sqrt{2}}\left(a_{0}-2 b_{0}^{\prime}+\frac{1}{2}\right) & 0<x \leq 8 \\
0 & x>8
\end{array}\right.
\end{aligned}
$$

$$
\begin{aligned}
& u_{3}^{\prime}(x)=\left\{\begin{array}{c}
\frac{1}{8 \sqrt{2}}\left[\left(\beta_{1}+2 \alpha_{1}^{\prime}\right)\left(\frac{x}{8}\right)+\left(-\alpha_{1}+2 \beta_{1}^{\prime}\right)\left(\frac{x}{8}\right)^{3}\right. \\
+\left(\beta_{2}+2 \alpha_{2}^{\prime}\right)\left(\frac{x}{8}\right)^{5}+\left(-\alpha_{2}+2 \beta_{2}^{\prime}\right)\left(\frac{x}{8}\right)^{7} \\
\left.+\cdots+\left(-\alpha_{7}+2 \beta_{7}^{\prime}\right)\left(\frac{x}{8}\right)^{27}\right] \quad . \quad 0<x \leq 8
\end{array}\right. \\
& 0 \quad x>8 \\
& \frac{1}{8 \sqrt{2}}\left[\frac { \pi } { 4 } \left\{\left(\beta_{1}+2 \alpha_{1}^{\prime}\right)\left(\frac{x}{8}\right)+\left(\alpha_{1}-2 \beta_{1}^{\prime}\right)\left(\frac{x}{8}\right)^{3}\right.\right. \\
& +\left(\beta_{2}+2 \alpha_{2}^{\prime}\right)\left(\frac{x}{8}\right)^{5}+\left(\alpha_{2}-2 \beta_{2}^{\prime}\right)\left(\frac{x}{8}\right)^{7} \\
& \left.+\cdots+\left(\alpha_{7}-2 \beta_{7}^{\prime}\right)\left(\frac{x}{8}\right)^{2 ?}\right\} \\
& +\left(-b_{1}-2 a_{1}^{\prime}\right)\left(\frac{x}{8}\right)+\left(a_{1}-2 b_{1}^{\prime}\right)\left(\frac{x}{8}\right)^{3} \\
& +\left(-b_{2}-2 a_{2}^{\prime}\right)\left(\frac{x}{8}\right)^{5}+\left(a_{2}-2 b_{2}^{\prime}\right)\left(\frac{x}{8}\right)^{7} \quad 0<x \leq 8 \\
& \begin{array}{l}
+\cdots+\left(a_{7}-2 b_{7}^{\prime}\right)\left(\frac{x}{8}\right)^{27} \\
\left(\frac{x}{8}\right)+\beta_{2}\left(\frac{x}{8}\right)^{=}+\alpha_{2}\left(\frac{x}{8}\right)^{5}+\beta_{3}\left(\frac{x}{8}\right)^{7}
\end{array} \\
& \left.\left.+\cdots+\beta_{7}\left(\frac{x}{8}\right)^{23}+\alpha_{7}\left(\frac{x}{8}\right)^{25}\right\}\right] \\
& \frac{\sqrt{\pi}}{2 x^{\frac{3}{2}}} e^{\nu(x)}[\cos \lambda(x)+\sin \lambda(x)+ \\
& \left.\frac{2}{x}\left\{s_{\phi}(x) \cos \lambda(x)+t_{\phi}(x) \sin \lambda(x)\right\}\right]-\frac{\sqrt{2}}{x^{3}} \quad x>8
\end{aligned}
$$

$$
\left.\begin{array}{c}
z_{3}^{\prime}(x)=\left\{\begin{array}{cc}
\frac{a_{0}-2 b_{0}^{\prime}+\frac{1}{2}}{\sqrt{2}} & 0<x \leq 8 \\
0 & x>8
\end{array}\right. \\
u_{4}(x)=u_{3}(x)+\frac{x}{\sqrt{2}}\left[u_{1}(x)+u_{1}^{\prime}(x)\right] \\
w_{4}(x)=w_{3}(x)+\frac{x}{\sqrt{2}}\left[v_{1}(x)+v_{1}^{\prime}(x)\right] \\
z_{4}(x)=z_{3}(x)+\frac{1}{\sqrt{2}} \\
u_{4}^{\prime}(x)=u_{3}^{\prime}(x)-\frac{x}{\sqrt{2}}\left[u_{1}(x)-u_{1}^{\prime}(x)\right]
\end{array}\right\} \begin{aligned}
& w_{5}^{\prime}(x)=w_{3}^{\prime}(x)-\frac{x}{\sqrt{2}}\left[v_{1}(x)-v_{1}^{\prime}(x)\right] \\
& u_{5}(x)=v_{0}^{\prime}(x)+v_{1}(x)
\end{aligned}
$$

$$
\begin{gathered}
u_{5}^{\prime}(x)=u_{0}^{\prime}(x)+u_{1}^{\prime}(x) \\
v_{5}^{\prime}(x)=v_{0}^{\prime}(x)+v_{1}^{\prime}(x) \\
u_{6}(x)=4 u_{3}(x)+\frac{x}{\sqrt{2}}\left[u_{1}(x)+u_{1}^{\prime}(x)\right] \\
w_{6}(x)=4 w_{3}(x)+\frac{x}{\sqrt{2}}\left[v_{1}(x)+v_{1}^{\prime}(x)\right] \\
z_{8}(x)=4 z_{3}(x)+\frac{1}{\sqrt{2}} \\
u_{6}^{\prime}(x)=4 u_{3}^{\prime}(x)-\frac{x}{\sqrt{2}}\left[u_{1}(x)-u_{1}^{\prime}(x)\right] \\
w_{8}^{\prime}(x)=4 w_{3}^{\prime}(x)-\frac{x}{\sqrt{2}}\left[v_{1}(x)-v_{1}^{\prime}(x)\right] \\
u_{8}(x)= \begin{cases}1 & x>x \leq 8 \\
0 & x\end{cases} \\
z_{8}^{\prime}(x)=4 z_{3}^{\prime}(x)+\frac{1}{\sqrt{2}}
\end{gathered}
$$

Functions needed for the above are as follows:
For $0<x \leq 8$ :

$$
\begin{gathered}
f_{0}(x)=a_{0}+a_{1}\left(\frac{x}{8}\right)^{4}+a_{2}\left(\frac{x}{8}\right)^{8}+\cdots+a_{7}\left(\frac{x}{8}\right)^{28} \\
g_{0}(x)=b_{1}\left(\frac{x}{8}\right)^{2}+b_{2}\left(\frac{x}{8}\right)^{6}+\cdots+b_{7}\left(\frac{x}{8}\right)^{28} \\
f_{0}^{\prime}(x)=a_{1}^{\prime}\left(\frac{x}{8}\right)^{2}+a_{2}^{\prime}\left(\frac{x}{8}\right)^{6}+\cdots+a_{7}^{\prime}\left(\frac{x}{8}\right)^{28} \\
g_{0}^{\prime}(x)=b_{0}^{\prime}+b_{1}^{\prime}\left(\frac{x}{8}\right)^{4}+b_{2}^{\prime}\left(\frac{x}{8}\right)^{8}+\cdots+b_{6}^{\prime}\left(\frac{x}{8}\right)^{24}
\end{gathered}
$$

For $x>8$ :

$$
\begin{gathered}
\nu(x)=c_{0}-c_{1}\left(\frac{8}{x}\right)+c_{2}\left(\frac{8}{x}\right)^{2}-c_{3}\left(\frac{8}{x}\right)^{3}+\cdots+c_{6}\left(\frac{8}{x}\right)^{6}-\frac{x}{\sqrt{2}} \\
\lambda(x)=d_{0}-d_{1}\left(\frac{8}{x}\right)+d_{2}\left(\frac{8}{x}\right)^{2}-d_{3}\left(\frac{8}{x}\right)^{3}+\cdots+d_{6}\left(\frac{8}{x}\right)^{6}-\frac{x}{\sqrt{2}} \\
s_{\phi}(x)=\left(s_{0}+t_{0}\right)-\left(s_{1}+t_{1}\right)\left(\frac{8}{x}\right)+\left(s_{2}+t_{2}\right)\left(\frac{8}{x}\right)^{2}-\left(s_{3}+t_{3}\right)\left(\frac{8}{x}\right)^{3}+\cdots+\left(s_{6}+t_{6}\right)\left(\frac{8}{x}\right)^{6} \\
t_{\phi}(x)=\left(s_{0}-t_{0}\right)-\left(s_{1}-t_{1}\right)\left(\frac{8}{x}\right)+\left(s_{2}-t_{2}\right)\left(\frac{8}{x}\right)^{2}-\left(s_{3}-t_{3}\right)\left(\frac{8}{x}\right)^{3}+\cdots+\left(s_{6}-t_{6}\right)\left(\frac{8}{x}\right)^{6}
\end{gathered}
$$

## List of Coefficients

$$
\begin{array}{ll}
\alpha_{1}=-64 & \alpha_{2}=113.77777774 \\
\alpha_{3}=-32.36345652 & \alpha_{4}=2.64191397 \\
\alpha_{5}=-0.08349609 & \alpha_{6}=0.00122552 \\
\alpha_{7}=-0.00000901 & \\
\beta_{1}=16 & \beta_{2}=-113.77777774 \\
\beta_{3}=72.81777742 & \beta_{4}=-10.56765779 \\
\beta_{5}=0.52185615 & \beta_{6}=-0.01103667 \\
\beta_{7}=0.00011346 & \alpha_{2}^{\prime}=14.22222222 \\
\alpha_{1}^{\prime}=-4 & \alpha_{4}^{\prime}=0.66047849 \\
\alpha_{3}^{\prime}=-6.06814810 & \beta_{8}^{\prime}=0.14677204 \\
\alpha_{5}^{\prime}=-0.02609253 & \alpha_{6}^{\prime}=0.00045957 \\
\alpha_{7}^{\prime}=-0.00000394 & \\
\beta_{5}^{\prime}=-0.00379386 & \beta_{6}^{\prime}=0.00004609 \\
\beta_{1}^{\prime}=-10.66666666 & \beta_{2}^{\prime}=11.37777772 \\
\beta_{7}^{\prime}=-2.31167514 & \beta_{4}^{\prime} \\
& \\
\hline
\end{array}
$$

$$
\begin{array}{ll}
a_{0}=-0.57721566 & a_{1}=-59.05819744 \\
a_{2}=171.36272133 & a_{3}=-60.60977451 \\
a_{4}=5.65539121 & a_{5}=-0.19636347 \\
a_{6}=0.00309699 & a_{7}=-0.00002458 \\
& \\
b_{1}=6.76454936 & b_{2}=-142.91827687 \\
b_{3}=124.23569650 & b_{4}=-21.30060904 \\
b_{5}=1.17509064 & b_{6}=-0.02695875 \\
b_{7}=0.00029532 & \\
a_{1}^{\prime}=-3.69113734 & a_{2}^{\prime}=-21.42034017 \\
a_{3}^{\prime}=-11.36433272 & a_{4}^{\prime}=1.41384780 \\
a_{5}^{\prime}=-0.06136358 & a_{6}^{\prime}=-0.00116137 \\
a_{7}^{\prime}=-0.00001075 & \\
b_{0}^{\prime}=0.21139217 & b_{1}^{\prime}=-13.39858846 \\
b_{2}^{\prime}=19.41182758 & b_{3}^{\prime}=-4.65950823 \\
b_{4}^{\prime}=0.33049424 & b_{5}^{\prime}=-0.00926707 \\
b_{6}^{\prime}=0.00011997 & b_{7}^{\prime}=0 \\
& \\
& \\
& \\
\hline
\end{array}
$$

$$
\begin{array}{ll}
c_{0}=0 & c_{1}=0.0110486 \\
c_{2}=0 & c_{3}=-0.0000906 \\
c_{4}=-0.0000252 & c_{5}=-0.0000034 \\
c_{0}=0.0000006 & \\
& \\
d_{0}=-0.3926991 & d_{1}=-0.0110485 \\
d_{2}=-0.0009765 & d_{3}=-0.0000901 \\
d_{4}=0 & d_{5}=0.0000051 \\
d_{6}=0.0000019 & \\
s_{0}=0.7071068 & s_{1}=-0.0625001 \\
s_{2}=-0.0013813 & s_{3}=0.0000005 \\
s_{4}=0.0000346 & s_{5}=0.0000117 \\
s_{8}=0.0000016 & \\
t_{0}=0.7071068 & t_{1}=-0.0000001 \\
t_{2}=0.0013811 & t_{3}=0.0002452 \\
t_{4}=0.0000338 & t_{5}=-0.0000024 \\
t_{6}=-0.0000032 &
\end{array}
$$

## Appendix G

## Real and Imaginary Parts of the

## Poroelastic Fundamental

## Solutions

The real and imaginary parts of the fundamental solutions are defined in equation (6.30). The functions $u_{n}, v_{n}, w_{n}$ and $z_{n}$ used in the following take argument $c r$ where the parameter $c$ is defined from (6.16) as

$$
c^{2}=\frac{\omega f k H}{R(\lambda+2 \mu)}
$$

The parameters $R, k$ and $H$ are given by (3.24), (5.118) and (5.121) respectively. The following dimensionless groups are useful in describing the fundamental
solutions

$$
\begin{gathered}
\Pi_{a}=\frac{\mu}{H}=\frac{1}{\frac{2(1-\nu)}{1-2 \nu}+\frac{K_{f}}{\mu f}} \\
\Pi_{b}=\frac{K_{f}}{H}=\frac{\frac{K_{f}}{\mu}}{\frac{2(1-\nu)}{1-2 \nu}+\frac{K_{f}}{\mu f}} \\
\Pi_{c}=\frac{1-2 \nu}{2(1-\nu)}+\frac{\mu f}{K_{f}}
\end{gathered}
$$

Thus in defining $c, \Pi_{a}, \Pi_{b}$ and $\Pi_{c}$ is necessary to know the following material constants: porosity $f$, permeability $k_{0}\left(m s^{-1}\right)$, shear modulus of soil skeleton $\mu\left(N m^{-2}\right)$, Poisson's ratio of soil skeleton $\nu$, density of pore water $\rho_{f}\left(\mathrm{~kg} \mathrm{~m} \mathrm{~m}^{-3}\right)$, bulk modulus of pore water $K_{f}\left(\mathrm{Nm}^{-2}\right)$.

The following dimensionless groups occur in the fundamental solutions. The symbol $r_{0}$ reprezents an arbitrary length.

$$
\begin{aligned}
\Pi_{1} & =\frac{H-\mu}{2 \tilde{\Pi}}=\frac{1}{2}\left(1-\Pi_{a}\right) \\
\Pi_{2} & =\frac{H+\mu}{2 H}=\frac{1}{2}\left(1+\Pi_{a}\right) \\
\Pi_{3} & =\frac{i k_{0} K_{f} a^{2} \mu}{\omega f^{2} \rho_{f} g H^{2}}=\frac{\Pi_{b}^{2} \Pi_{c}}{f^{2}} \\
\Pi_{4} & =\frac{K_{f} a^{2} r_{0}}{f H} \equiv-i \Pi_{4}^{\prime}
\end{aligned}
$$

where

$$
\begin{gathered}
\Pi_{4}^{\prime}=\frac{\Pi_{b} c^{2} r_{0}}{f} \\
\Pi_{5}=\frac{i k_{0} a^{2} K_{f} \mu}{\omega f \rho_{f} g H}=\frac{\Pi_{b} \Pi_{c}}{f} \\
\Pi_{6}=a r_{0} \equiv\left(\frac{1-i}{\sqrt{2}}\right) \Pi_{6}^{\prime}
\end{gathered}
$$

where

$$
\begin{gathered}
\Pi_{6}^{\prime}=c r_{0} \\
\Pi_{7}=\frac{2(H-\mu)}{H}=2\left(1-\Pi_{a}\right) \\
\Pi_{8}=\frac{\mu}{H}=\Pi_{a} \\
\Pi_{G}=\frac{i 2 \mu K_{f}^{2} k_{0} a^{3} r_{0}}{\omega f^{2} \rho_{f} g H^{2}} \equiv(1-i) \Pi_{9}^{\prime}
\end{gathered}
$$

where

$$
\begin{gathered}
\Pi_{9}^{\prime}=\frac{\sqrt{2} \Pi_{b}^{2} \Pi_{c} c r_{0}}{f^{2}} \\
\Pi_{10}=\frac{K_{f} a^{2} r_{0}^{2}}{f H} \equiv-i \Pi_{10}^{\prime}
\end{gathered}
$$

where

$$
\begin{gathered}
\Pi_{10}^{\prime}=\frac{\Pi_{b} c^{2} r_{0}^{2}}{f} \\
\Pi_{11}=\frac{K_{f} a^{2} r_{0}}{f H}=-i \Pi_{4}^{\prime} \\
\Pi_{12}=-i \Pi_{12}^{\prime}
\end{gathered}
$$

where

$$
\Pi_{12}^{\prime}=\frac{\rho_{g} g \omega r_{0}^{2}}{\mu k_{0}}
$$

The functions defined in equation (6.30) are given below.

## Function $H_{i k}$

For $i=1,2$
For $k=1,2$

$$
H_{i k}=-\Pi_{2} \delta_{i k} u_{\mathrm{B}}+\Pi_{3}\left(u_{2} r_{, i} r_{, k}-u_{1} \delta_{i k}\right)
$$

For $k=3$

$$
H_{i k}=\Pi_{4}^{\prime} r r_{, i} u_{1}^{\prime}
$$

## For $i=3$

For $k=1,2$

$$
H_{i k}=\Pi_{5}\left(-u_{2} r_{, k} \frac{\partial r}{\partial n}+u_{1} n_{k}\right)
$$

For $k=3$

$$
H_{i k}=-c^{2} r_{0} r \frac{\partial r}{\partial n} u_{1}^{\prime}
$$

Function $H_{i k}^{\prime}$
For $i=1,2$
For $k=1,2$

$$
H_{i k}^{\prime}=\Pi_{3}\left(u_{2}^{\prime} r_{, i} r_{, k}-u_{1}^{\prime} \delta_{i k}\right)
$$

For $k=3$

$$
H_{i k}^{\prime}=-\Pi_{4}^{\prime} r r_{, i} u_{1}
$$

For $i=3$
For $k=1,2$

$$
H_{i k}^{\prime}=\Pi_{5}\left(-u_{2}^{\prime} r_{, k} \frac{\partial r}{\partial n}+u_{1}^{\prime} n_{k}\right)
$$

For $k=3$

$$
H_{i k}^{\prime}=c^{2} r_{0} r \frac{\partial r}{\partial n} u_{1}
$$

## Function $G_{\text {ik }}$

For $i=1,2$
For $k=1,2$

$$
G_{i k}=\Pi_{1} r_{, i} r_{, k}+\Pi_{2} \delta_{i k}\left[\ln \left(\frac{c}{2}\right)-\left(1-u_{8}\right) \ln \left(\frac{c r}{2}\right)\right]+\Pi_{3}\left(v_{2} r_{, i} r_{, k}-v_{1} \delta_{i k}\right)
$$

For $k=3$

$$
G_{i k}=\Pi_{4}^{\prime} r r_{, i} v_{1}^{\prime}
$$

## For $i=3$

For $k=1,2$

$$
G_{i k}=\Pi_{5}\left(-v_{2} r_{, k} \frac{\partial r}{\partial n}+v_{1} n_{k}\right)
$$

For $k=3$

$$
G_{i k}=-c^{2} r_{0} r \frac{\partial r}{\partial n} v_{1}^{\prime}
$$

## Function $G_{i k}^{\prime}$

For $i=1,2$
For $k=1,2$

$$
G_{i k}^{\prime}=\Pi_{3}\left(v_{2}^{\prime} r_{;} r_{, k}-v_{1}^{\prime} \delta_{i k}\right)
$$

For $k=3$

$$
G_{i k}^{\prime}=-\Pi_{4}^{\prime} r r_{. i} v_{1}
$$

## For $i=3$

For $k=1,2$

$$
G_{i k}^{\prime}=\Pi_{5}\left(-v_{2}^{\prime} r_{, k} \frac{\partial r}{\partial n}+v_{1}^{\prime} n_{k}\right)
$$

For $k=3$

$$
G_{i k}^{\prime}=-c^{2} r_{0} r \frac{\partial r}{\partial n} v_{1}
$$

## Function $P_{\text {ik }}$

For $i=1,2$
For $k=1,2$

$$
P_{i k}=\Pi_{9}^{\prime}\left\{\left(\frac{\partial r}{\partial n} \delta_{i k}+r_{, i} n_{k}\right)\left(u_{3}+u_{3}^{\prime}\right)-\frac{\partial r}{\partial n} r_{, i} r_{, k}\left(u_{6}+u_{6}^{\prime}\right)+r_{, k} n_{i}\left(u_{4}+u_{4}^{\prime}\right\}\right.
$$

For $k=3$

$$
P_{i k}=\Pi_{10}^{\prime}\left(-u_{2}^{\prime}\left(2 r_{, i} \frac{\partial r}{\partial n}-n_{i}\right)+u_{0}^{\prime} n_{i}\right\}
$$

## For $i=3$

For $k=1,2$

$$
P_{i k}=\Pi_{4}^{\prime} r_{, k} r u_{1}^{\prime}
$$

For $k=3$

$$
P_{i k}=\Pi_{12}^{\prime} u_{0}^{\prime}
$$

## Function $P_{i k}^{\prime}$

For $i=1,2$
For $k=1,2$
$P_{i k}^{\prime}=\Pi_{9}^{\prime}\left\{-\left(\frac{\partial r}{\partial n} \delta_{i k}+r_{, i} n_{k}\right)\left(u_{3}-u_{3}^{\prime}\right)+\frac{\partial r}{\partial n} r_{i,} r_{, k}\left(u_{6}-u_{6}^{\prime}\right)-r_{, k} n_{i}\left(u_{4}-u_{4}^{\prime}\right)\right\}$
For $k=3$

$$
P_{i k}^{\prime}=\Pi_{10}^{\prime}\left\{u_{2}\left(2 r_{i} \frac{\partial r}{\partial n}-n_{i}\right)+u_{0} n_{i}\right\}
$$

## For $i=3$

For $k=1,2$

$$
P_{i k}^{\prime}=-\Pi_{4}^{\prime} r_{, k} r u_{1}
$$

For $k=3$

$$
P_{i k}^{\prime}=-\Pi_{12}^{\prime} u_{0}
$$

Function $Q_{i k}$
For $i=1,2$
For $k=1,2$
$Q_{i k}=\Pi_{9}^{\prime}\left\{\left(\frac{\partial r}{\partial n} \delta_{i k}+r_{i,} n_{k}\right)\left(w_{3}+w_{3}^{\prime}\right)-\frac{\partial r}{\partial n} r_{, i} r_{, k}\left(w_{6}+w_{6}^{\prime}\right)+r_{, k} n_{i}\left(w_{4}+w_{4}^{\prime}\right)\right\}$
For $k=3$

$$
Q_{i k}=\Pi_{10}^{\prime}\left\{-v_{2}^{\prime}\left(2 r_{, i} \frac{\partial r}{\partial n}-n_{i}\right)+v_{0}^{\prime} n_{i}\right\}
$$

## For $i=3$

For $k=1,2$

$$
Q_{i k}=\Pi_{4}^{\prime} r, k r v_{1}^{\prime}
$$

For $k=3$

$$
Q_{i k}=\Pi_{12}^{\prime} v_{0}^{\prime}
$$

## Function $Q_{i k}^{\prime}$

For $i=1,2$
For $k=1,2$
$Q_{i k}^{:}=\Pi_{9}^{\prime}\left\{-\left(\frac{\partial r}{\partial n} \delta_{i k}+r_{, i} n_{k}\right)\left(u_{3}-w_{3}^{\prime}\right)+\frac{\partial r}{\partial n} r_{, i} r_{, k}\left(w_{6}-w_{6}^{\prime}\right)-r_{, k} n_{i}\left(w_{4}-w_{4}^{\prime}\right)\right\}$
For $k=3$

$$
Q_{i k}^{\prime}=\Pi_{10}^{\prime}\left\{v_{2}\left(2 r_{i} \frac{\partial r}{\partial n}-n_{i}\right)-v_{0} n_{i}\right\}
$$

## For $i=3$

For $k=1,2$

$$
Q_{i k}^{\prime}=-\Pi_{4}^{\prime} r_{, k} r v_{1}
$$

For $k=3$

$$
Q_{i k}^{\prime}=-\Pi_{12}^{\prime} v_{0}
$$

## Function $Z_{\text {ik }}$

For $i=1,2$
For $k=1,2$

$$
\begin{aligned}
Z_{i k}= & -\Pi_{7} c r_{0} \frac{\partial r}{\partial n} r_{, i} r_{, k}-\Pi_{8} c r_{0}\left(\frac{\partial r}{\partial n} \delta_{i k}+r_{, i} n_{k}-r_{, k} n_{i}\right) \\
& +\Pi_{9}^{\prime}\left[\left(\frac{\partial r}{\partial n} \delta_{i k}+r_{, i} \pi_{k}\right)\left(z_{3}+z_{3}^{\prime}\right)-\frac{\partial r}{\partial n} r_{, i} r_{, k}\left(z_{8}+z_{8}^{\prime}\right)\right. \\
& \left.+r_{, k} n_{i}\left(z_{4}+z_{4}^{\prime}\right)\right]
\end{aligned}
$$

For $k=3$

$$
Z_{i k}=0
$$

## For $i=3$

$$
Z_{i k}=0 \text { for } k=1,2,3
$$

Function $Z_{i k}^{\prime}$
For $i=1,2$
For $k=1,2$

$$
z_{i k}^{\prime}=\Pi_{9}^{\prime}\left[-\left(\frac{\partial r}{\partial n} \delta_{i k}+r_{, i} n_{k}\right)\left(z_{3}-z_{3}^{\prime}\right)+\frac{\partial r}{\partial n} r_{i} r_{, k}\left(z_{6}-i_{6}^{\prime}\right)-r_{, k} n_{i}^{\prime}\left(x_{4}-z_{4}^{\prime}\right)\right]
$$

For $k=3$

$$
z_{i k}^{\prime}=0
$$

## For $i=3$

$$
Z_{i k}^{\prime}=0 \text { for } k=1,2,3
$$

## Appendix H

## Functions Required for the

## Evaluation of Interior Effective

## Stresses and Pore Pressure

These functions are the integrands occurring in equations (6.60)and (6.61). In the following,

$$
\begin{aligned}
& \eta \equiv \frac{2 \nu}{1-2 \nu} \text { where } \nu \text { is Poisson's ratio of the solid skeleton } \\
& \qquad a^{2}=-i c^{2}, \quad a=e^{-\frac{i \pi}{4}} c=\frac{c}{\sqrt{2}}(1-i)
\end{aligned}
$$

$K_{0}$ and $K_{1}$ are the modified Bessel functions of the second hind of orders 0 and 1 respectively.

$$
K_{0}(a r)=k e r(c r)-i k e i(c r)
$$

$$
K_{1}(a r)=-\frac{1}{\sqrt{2}}\left[k e r^{\prime}(c r)+k e i^{\prime}(c r)\right]+\frac{i}{\sqrt{2}}\left[k e i^{\prime}(c r)-k e r^{\prime}(c r)\right]
$$

where $k e r, k e i, k e r^{\prime}, k e i^{\prime}$ are the Kelvin functions.

$$
\begin{aligned}
F_{1}(a r)= & \frac{2}{a^{2} r^{2}}-K_{0}(a r)-\frac{2}{a r} K_{1}(a r) \\
= & -k e r(c r)+\frac{2}{c r} k e i^{\prime}(c r) \\
& +i\left[\frac{2}{c^{2} r^{2}}+k e i(c r)+\frac{2}{c r} k e r^{\prime}(c r)\right]
\end{aligned}
$$

The functions $D_{\alpha k j}, S_{\alpha k j}, \tilde{\bar{u}}_{\alpha}^{* 3}, \tilde{T}_{\alpha}^{* 3}$ are defined below.

Function $D_{\alpha k j}$
For $\alpha=1,2$
For $k=1,2$ and $j=1,2$

$$
\begin{aligned}
D_{\alpha k j}= & -\Pi_{1} \frac{1}{r}\left(2 r_{, \alpha} \delta_{k j}+r_{, k} \delta_{\alpha j}+r_{, j} \delta_{\alpha k}-4 r_{, \alpha} r_{, j} r_{, k}\right) \\
& +\Pi_{2} \frac{1}{r}\left(r_{, j} \delta_{\alpha k}+r_{, k} \delta_{\alpha j}\right) \\
& +\Pi_{3}\left\{\frac{2}{r} F_{1}(a r)\left[r_{, \alpha} \delta_{k j}+r_{, k} \delta_{\alpha j}+r_{, j} \delta_{\alpha k}-4 r_{, \alpha} r_{, j} r_{, k}\right]\right. \\
& \left.+2 a K_{1}(a r) r_{, \alpha} r_{, j} r_{, k}\right\} \\
& -\eta r_{, \alpha} \delta_{k j}\left[\frac{1}{r}\left(\Pi_{1}-\Pi_{2}\right)-\Pi_{3} a K_{1}(a r)\right]
\end{aligned}
$$

For $\alpha=3$
For $k=1,2$ and $j=1,2$

$$
\begin{aligned}
& D_{\alpha k j}=-\Pi_{5}\left\{\frac{2}{r} F_{1}(a r)\left[\frac{\partial r}{\partial n}\left(\delta_{k j}-4 r_{, k} r_{, j}\right)+r_{, k} n_{j}+r_{, j} n_{k}\right]\right. \\
&\left.+a \frac{\partial r}{\partial n} K_{1}(a r)\left[2 r_{, j} r_{, k}+\eta \delta_{k j}\right]\right\}
\end{aligned}
$$

## Function $S_{\alpha k j}$

For $\alpha=1,2$
For $k=1,2$ and $j=1,2$

$$
\begin{aligned}
S_{\alpha k j}= & r_{0} \Pi_{7} \frac{1}{r^{2}}\left\{\frac{\partial r}{\partial n}\left(2 r_{, \alpha} \delta_{k j}+r_{, k} \delta_{\alpha j}+r_{, j} \delta_{\alpha k}-8 r_{, \alpha} r_{, j} r_{, k}\right)\right. \\
& \left.+r_{, \alpha}\left(r_{, k} n_{j}+r_{, j} n_{k}\right)\right\} \\
& +r_{0} \Pi_{8} \frac{1}{r^{2}}\left\{2\left(n_{j} \delta_{\alpha k}+n_{k} \delta_{\alpha j}\right)-2 r_{, j}\left(\frac{\partial r}{\partial n} \delta_{\alpha k}+r_{, \alpha} n_{k}\right)\right. \\
& \left.-2 r_{, k}\left(\frac{\partial r}{\partial n} \delta_{\alpha j}+r_{, \alpha} n_{j}\right)-2\left(\delta_{k j}-2 r_{, k} r_{, j}\right) n_{\alpha}\right\} \\
& +2(1-i) \Pi_{9}^{\prime}\left\{\frac { F _ { 1 } ( a r ) } { a r ^ { 2 } } \left[4 \frac{\partial r}{\partial n}\left(6 r_{, \alpha} r_{, k} r_{, j}-r_{, \alpha} \delta_{k j}-r_{, k} \delta_{\alpha j}-r_{, j} \delta_{\alpha k}\right)\right.\right. \\
& +n_{\alpha} \delta_{k j}+n_{k} \delta_{\alpha j}+n_{j} \delta_{\alpha k} \\
& \left.-4\left(r_{, \alpha} r_{, k} n_{j}+r_{, \alpha} r_{, j} n_{k}+r_{, k} r_{, j} n_{\alpha}\right)\right] \\
& +\frac{K_{1}(a r)}{r}\left[\frac{\partial r}{\partial n}\left(r_{, j} \delta_{\alpha k}+r_{, \alpha} \delta_{k j}+r_{, k} \delta_{\alpha j}-8 r_{, \alpha} r_{, k} r_{, j}\right)\right. \\
& \left.+r_{, \alpha}\left(r_{, k} n_{j}+r_{, j} n_{k}\right)+n_{\alpha}\left(3 r_{, k} r_{, j}-\delta_{k j}\right)\right] \\
& \left.+a K_{0}(a r)\left[-\frac{\partial r}{\partial n} r_{, \alpha} r_{, k} r_{, j}+r_{, k} r_{, j} n_{\alpha}\right]\right\} \\
& +\eta \delta_{k j}\left\{2 r_{0} \Pi_{8} \frac{1}{r^{2}}\left(n_{\alpha}-2 r_{, \alpha} \frac{\partial r}{\partial n}\right)\right. \\
& +(1-i) \Pi_{9}^{\prime}\left[\frac{K_{1}(a r)}{r}\left(n_{\alpha}-2 r_{, \alpha} \frac{\partial r}{\partial n}\right)\right. \\
& \left.\left.+a K_{0}(a r)\left(n_{\alpha}-r_{, \alpha} \frac{\partial r}{\partial n}\right)\right]\right\}
\end{aligned}
$$

For $\alpha=3$
For $k=1,2$ and $j=1,2$

$$
S_{\alpha k j}=-i \Pi_{4}^{\prime}\left\{(1+\eta) \delta_{k j} K_{0}(a r)+F_{1}(a r)\left(\delta_{k j}-2 r_{, k} r_{, j}\right)\right\}
$$

$\underline{\underline{\text { Function }} \tilde{\tilde{u}}_{\alpha}^{* 3}}$
For $\alpha=1,2$

$$
\tilde{\tilde{u}}_{\alpha}^{* 3}=\frac{1}{2} i \Pi_{4}^{\prime} r r_{, \alpha}\left[F_{1}(a r)+K_{0}(a r)\right]
$$

For $\alpha=3$

$$
\tilde{\bar{u}}_{\alpha}^{* 3}=-\frac{(1-i)}{\sqrt{2}} \Pi_{6}^{\prime} \frac{\partial r}{\partial n} K_{1}(a r)
$$

$\underline{\text { Function } \tilde{\bar{T}}_{\alpha}^{* 3}}$
For $\alpha=1,2$

$$
\tilde{T}_{\alpha}^{* 3}=-i \Pi_{10}^{\prime}\left\{\left(2 r_{, \alpha} \frac{\partial r}{\partial n}-n_{\alpha}\right) F_{1}(a r)+K_{0}(a r) n_{\alpha}\right\}
$$

For $\alpha=3$

$$
\tilde{\bar{T}}_{\alpha}^{* 3}=-i \Pi_{12}^{\prime} K_{0}(a r)
$$

## Appendix I

## Wave loading of a Flat

## Homogeneous Isotropic

## Poroelastic Seabed - Analytic

## Solution

For the case of a flat seabed, it is possible to determine by analytic means the wave induced effective stresses and pore pressures. We present here a summary of the technique used by the author in his M.Eng thesis (Raman-Nair, 1985). It was demonstrated that for sand beds the acceleration terms may be deleted from the governing equations. We thus obtain from (3.17) and (3.18), for no body forces,

$$
\begin{equation*}
\tau_{i j, j}=0 \tag{I.1}
\end{equation*}
$$

$$
\begin{equation*}
-p_{, i}=\frac{\rho_{f} g}{k_{0}} \dot{w}_{i} \tag{I.2}
\end{equation*}
$$

where

$$
w_{i}=f\left(U_{i}-u_{i}\right)
$$

We also recall from (3.5) and (3.26) the constitutive laws in the form

$$
\begin{align*}
& \tau_{i j}=2 \mu e_{i j}+\delta_{i j} \lambda e-\delta_{i j} p  \tag{I.3}\\
& -p=\frac{(1-f)}{f} K_{f} e+K_{f} \varepsilon \tag{I.4}
\end{align*}
$$

where we have used the fact that $K_{r} \rightarrow \infty$,so that $\alpha \approx 1$. Substituting (I.3) into ( I.1) gives

$$
\begin{equation*}
\mu u_{i, j j}+(\lambda+\mu) u_{j, j i}=p_{, i} \tag{1.5}
\end{equation*}
$$

We shall rewrite equation (I.2) in terms of the soil displacement vector $u_{i}$ (rather than $w_{i}$ ). Differentiating (I.2) with respect to $x_{i}$ gives

$$
\begin{equation*}
-p_{, i i}=\frac{\rho_{f} g f}{k_{0}} \frac{\partial}{\partial t}(\varepsilon-e) \tag{I.6}
\end{equation*}
$$

From ( I.4):

$$
K_{f}(\varepsilon-e)=-p-\frac{K_{f}}{f} e
$$

Substituting this into ( I.6) gives:

$$
\begin{equation*}
\frac{k_{0}}{\rho_{f} g} \nabla^{2} p=\frac{f}{K_{f}} \frac{\partial p}{\partial t}+\frac{\partial e}{\partial t} \tag{I.7}
\end{equation*}
$$

This is the so-called storage equation derived in a different way by Verruijt (1969) and Biot (1941).

The system of equations (I.5) and (I.7) must be solved subject to appropriate boundary conditions. We take the $x$-axis on the flat seabed and parallel to the
direction of propagation of the wave. The $z$-axis is vertically downward into the soil. The problem domain is two dimensional in the $x-z$ plane and is illustrated in Figure (I.1). The boundary conditions are:

$$
\begin{aligned}
& \text { (a) } \tau_{x x}^{\prime}=\tau_{x x}^{\prime}=0 \text { at } z=0 \\
& \text { (b) } p(\text { in bed })=P_{0} e^{i(k x-\omega t)}
\end{aligned}
$$

where $P_{0}$ is the amplitude of the wave induced pressure on the seabed, $k$ is the wave number and $\omega$ is the circular wave frequency.

$$
\begin{array}{ll}
\text { (c) } u_{1}=u_{2}=0 & \text { at } z=z_{0}
\end{array}
$$

where $u_{1}$ and $u_{2}$ denote the $x$ and $z$ components of the soil displacement vector $\underline{y}$.

$$
\text { (d) } \frac{\partial p}{\partial z}=0 \quad \text { at } z=z_{0}
$$

The value of $P_{0}$ is determined from the wave velocity potential $\Phi$ which is given by:

$$
\begin{equation*}
\Phi(x, z, t)=\frac{i \eta_{0} g}{\omega}\left[\cosh k(z+h)-\frac{\omega^{2}}{k h} \sinh k(z+h)\right] e^{i(k x-\omega t)} \tag{I.8}
\end{equation*}
$$

where $\eta_{0}$ is the wave amplitude. Then from Bernoulli's equation,

$$
P_{0} e^{i(k x-\omega t)}=-\rho_{f} \frac{\partial \Phi}{\partial t} \quad \text { at } z=0
$$

so that

$$
\begin{equation*}
P_{0}=\rho_{j} \eta_{0} g\left[\cosh (k h)-\frac{\omega^{2}}{g k} \sinh (k h)\right] \tag{I.9}
\end{equation*}
$$

In view of the dispersion relation

$$
\omega^{2}=g k \tanh (k h)
$$

$$
\text { Wave Form: } \eta=\eta_{0} e^{i(k z-\omega t)}
$$



FIG.(I.1) Wave Loading of Flat Seabed
we can write ( I.9) more compactly as

$$
\begin{equation*}
P_{0}=\frac{\rho_{f} \eta_{0} g}{\cosh (k h)} \tag{I.10}
\end{equation*}
$$

Equation (I.5) is satisfied by the function

$$
\begin{equation*}
u_{i}=\left(\phi+x_{j} \psi_{j}\right)_{i}-4(1-\nu) \psi_{i} \tag{I.11}
\end{equation*}
$$

where

$$
\nu=\frac{\lambda}{2(\lambda+\mu)} \text { (Poisson's ratio) }
$$

provided that

$$
\begin{equation*}
\frac{2 \mu(1-\nu)}{1-2 \nu}\left(\nabla^{2} \phi+x_{j} \nabla^{2} \psi_{j}\right)_{, i}-4 \mu(1-\nu) \nabla^{2} \psi_{i}=p_{, i} \tag{I.12}
\end{equation*}
$$

Equation ( 1.11 ) is the well known Papkovich-Neuber solution of the equilibrium equations of theory of elasticity, $\phi$ and $\psi_{i}$ being functions of $x, z$, and $t$. Any one of the functions $\phi, \psi_{1}, \psi_{2}$ may be taken to be zero without loss of completeness provided that the coordinate system is chosen in an appropriate way and $4 \nu$ is not a positive integer. We choose

$$
\psi \equiv 0 \text { and } \psi_{2} \equiv \psi
$$

Then equations ( I.12) become

$$
\begin{gather*}
2 \mu \beta \frac{\partial}{\partial x}\left(\nabla^{2} \phi+z \nabla^{2} \psi\right)=\frac{\partial p}{\partial x}  \tag{I.13}\\
2 \mu \beta \frac{\partial}{\partial z}\left(\nabla^{2} \phi+z \nabla^{2} \psi\right)-4 \mu(1-\nu) \nabla^{2} \psi=\frac{\partial p}{\partial z} \tag{I.14}
\end{gather*}
$$

where

$$
\begin{equation*}
\beta=\frac{1-\nu}{1-2 \nu} \tag{I.15}
\end{equation*}
$$

Equations ( 1.13 ) and ( I.14) represent equation ( I.5) in terms of the functions $\phi$ and $\psi$. Using (I.11), we express (I.7) in terms of $\phi$ and $\psi$ :

$$
\begin{equation*}
\frac{k_{0}}{\rho_{f} g} \nabla^{2} p-\frac{f}{K_{f}} \frac{\partial p}{\partial t}=\frac{\partial}{\partial t}\left[\nabla^{2} \phi+z \nabla^{2} \psi-2(1-2 \nu) \frac{\partial \psi}{\dot{\partial} z}\right] \tag{I.16}
\end{equation*}
$$

It can be shown (Raman-Nir, 1985) that the solution of (I.13), (I.14) and (I.16) is of the form

$$
\begin{align*}
& \phi(x, z, t)=\left[\left(a_{1}+a_{2} z\right) e^{-k z}+\left(a_{3}+a_{4} z\right) e^{k z}+a_{5} e^{-k^{\prime} x}+a_{6} e^{k^{\prime} x}\right] e^{i(k x-\omega t)}(\mathrm{I} .17) \\
& \psi(x, z, t)=\left(B e^{-k z}+D e^{k z}\right) e^{i(k x-\omega t)} \tag{I.18}
\end{align*}
$$

where

$$
\begin{aligned}
B & =\frac{-4 \mu \beta a_{2}}{\alpha_{0}} \\
D & =\frac{-4 \mu \beta a_{4}}{\alpha_{0}} \\
\alpha_{0} & =-2(1-\nu)\left[\frac{f}{K_{\rho}}+\frac{1}{2 \mu \beta}\right]^{-1}
\end{aligned}
$$

and

$$
\begin{equation*}
p(x, z, t)=\left(A_{1} e^{-k z}+A_{2} e^{k x}+A_{3} e^{-k^{\prime} x}+A_{4} e^{k^{\prime} z}\right) e^{i(k z-\omega t)} \tag{I.19}
\end{equation*}
$$

where

$$
\begin{aligned}
A_{1} & =-4 \mu \beta k a_{2} \\
A_{2} & =4 \mu \beta k a_{4} \\
A_{3} & =2 \mu \beta\left[\left(k^{\prime}\right)^{2}-k^{2}\right] a_{5} \\
A_{1} & =2 \mu \beta\left[\left(k^{\prime}\right)^{2}-k^{2}\right] a_{6} \\
\left(k^{\prime}\right)^{2} & =k^{2}-\frac{i \omega}{c} \\
c & =\frac{k_{0}}{\rho_{f} g}\left[\frac{f}{K_{f}}+\frac{1}{2 \mu \beta}\right]^{-1}
\end{aligned}
$$

Using the constitutive laws the effective stress components $\tau_{i j}^{\prime}$ are written in terms of $\phi$ and $\psi$ :

$$
\begin{align*}
\tau_{i j}^{\prime}= & 2 \mu\left[\phi_{, i j}+z \psi_{, i j}-(1-2 \nu)\left(\psi_{i, j}+\psi_{j, i}\right)\right] \\
& +\delta_{i j}\left(\frac{2 \mu \nu}{1-2 \nu}\right)\left[\nabla^{2} \phi+z \nabla^{2} \psi-2(1-2 \nu) \frac{\partial \psi}{\partial z}\right] \tag{I.20}
\end{align*}
$$

Hence the effective stress components and pore pressure can be expressed in terms of the six constants $a_{1}, a_{2}, \cdots a_{6}$. The six boundary conditions lead to an equation of the form

$$
\begin{equation*}
[A] \underline{x}=\underline{\mathbf{b}} \tag{I.21}
\end{equation*}
$$

where

$$
\begin{aligned}
& \underline{x}=\left(a_{1}, a_{2}, a_{3}, a_{4}, a_{5}, a_{6}\right)^{T} \\
& \underline{b}=\left(0,0, \frac{P_{0}}{2 \mu \beta}, 0,0,0\right)^{T}
\end{aligned}
$$

The elements $a_{i j}$ of the matrix [ $A$ ] are given by

$$
\begin{aligned}
& a_{11}=k^{2} \\
& a_{12}=-2 k \beta\left[1+\frac{4(1-\nu) \mu}{\alpha_{0}}\right] \\
& a_{13}=k^{2} \\
& a_{14}=-a_{12} \\
& a_{15}=k^{2}+\beta\left[\left(k^{\prime}\right)^{2}-k^{2}\right] \\
& a_{16}=a_{15} \\
& a_{21}=-k \\
& a_{22}=1+\frac{4(1-\nu) \mu}{\alpha_{0}} \\
& a_{23}=k
\end{aligned}
$$

$$
\begin{aligned}
& a_{24}=a_{32} \\
& a_{25}=-k^{\prime} \\
& a_{26}=k^{\prime} \\
& a_{31}=0 \\
& a_{32}=-2 k \\
& a_{33}=0 \\
& a_{34}=2 k \\
& a_{35}=\left(k^{\prime}\right)^{2}-k^{2} \\
& a_{36}=a_{35} \\
& a_{41}=e^{-k z_{0}} \\
& a_{42}=z_{0} e^{-k z_{0}}\left(1-\frac{4 \mu \beta}{\alpha_{0}}\right) \\
& a_{43}=e^{k z_{0}} \\
& a_{44}=z_{0} e^{k z_{0}}\left(1-\frac{4 \mu \beta}{\alpha_{0}}\right) \\
& a_{45}=e^{-k^{\prime} z_{0}} \\
& a_{46}=e^{k^{\prime} x_{0}} \\
& a_{51}=-k e^{-k z_{0}} \\
& a_{52}=\left[-k z_{0}+1+\left(k z_{0}+3-4 \nu\right) \frac{4 \mu \beta}{\alpha_{0}}\right] e^{-k z_{0}} \\
& a_{53}=k e^{k z_{0}} \\
& a_{54}=\left[k z_{0}+1+\left(-k z_{0}+3-4 \nu\right) \frac{4 \mu \beta}{\alpha_{0}}\right] e^{k x_{0}} \\
& a_{55}=-k^{\prime} e^{-k^{\prime} x_{0}} \\
& a_{56}=k^{\prime} e^{k^{\prime} z_{0}} \\
& a_{61}=0 \\
& a_{0}
\end{aligned}
$$

$$
\begin{aligned}
& a_{62}=2 k^{2} e^{-k x_{0}} \\
& a_{63}=0 \\
& a_{64}=2 k^{2} e^{k z_{0}} \\
& a_{65}=-k^{\prime}\left[\left(k^{\prime}\right)^{2}-k^{2}\right] e^{-k^{\prime} z_{0}} \\
& a_{66}=k^{\prime}\left[\left(k^{\prime}\right)^{2}-k^{2}\right] e^{k^{\prime} z_{0}}
\end{aligned}
$$

## Appendix J

Stress Angle Contours


Figure J.1: Stress Angle Contours(degrees): Fine Sand, $K=0.5, L=300 \mathrm{~m}, \beta=20^{\circ}$

$$
l_{0}=14 \mathrm{~m}, l_{0}^{\prime}=34 \mathrm{~m}
$$

Horizontal Distance (m.)


## Horizontal Distance (m.)



Horizoñtal Distance (m.)


Horizontal Distance (m.)


Figure J.2: Stress Angle Contours(degrees): Fine Sand, $h=0.5, L=300 \mathrm{~m}, \beta=12^{\prime \prime}$

## Horizontal Distance (m.)



## Horizontal Distance (m.)



## Horizontal Distance (m.)



## Horizontal Distance (m.)



Figure J.3: Stress Angle Contours(degrees):Fine Sand, $K^{\prime}=0.5, L=300 \mathrm{~m}, \beta=5^{\circ}$


Figure J.4: Stress Angle Contours(degrees):Fine Sand, $\kappa=0.5, L=300 \mathrm{~m}, \beta=2^{\circ}$

Horizontal Distance (m.)


Horizontal Distance (m.)


## Horizontal Distance (m.)



Horizontal Distance (m.)


Figure J.5: Stress Angle Contours(degrees):Fine Sand, $K^{\prime}=0.5, L=225 \mathrm{~m}, \beta=12^{\circ}$

Horizontal Distance (m.)


Horizontal Distance (m.)


Horizontal Distance (m.)


Horizontal Distance (m.)


Figure J.6: Stress Angle Contours(degrees):Fine Sand, $K^{\prime}=0.5, L=225 \mathrm{~m}, \beta=5^{\circ}$


Figure J.7: Stress Angle Contours(degrees):Fine Sand, $K=0.5, L=225 \mathrm{~m}, \beta=2^{\circ}$


Figure J.8: Stress Angle Contours(degrees):Fine Sand, $K^{\prime}=0.5, L=150 \mathrm{~m}, \beta=20^{\circ}$


Figure J.9: Stress Angle Contours(degrees):Fine Sand, $K^{\prime}=0.5, L=150 \mathrm{~m}, \beta=12^{\circ}$


Figure J.10: Stress Angle Contours(degrees):Fine Sand, $K^{\prime}=0.5, L=150 \mathrm{~m}, \beta=5^{\circ}$


Figure J.11: Stress Angle Contours(degrees):Fine Sand, $K^{\prime}=0.5, L=150 \mathrm{~m}, \beta=2^{\circ}$


Figure J.12: Stress Angle Contours(degrees): Coarse Sand, $K^{\prime}=0.5, L=300 \mathrm{~m}, \beta=12^{\circ}$


Figure J.13: Stress Angle Contours(degrees): Coarse Sand, $K^{\prime}=0.5, L=225 \mathrm{~m}, \beta=12^{\circ}$

Horizontal Distance (m.)


Horizontal Distance (m.)


Horizontal Distance (m.)


Horizontal Distance (m.)


Figure J.14: Stress Angle Contours(degrees):Coarse Sand, $K^{\prime}=0.5, L=225 \mathrm{~m}, \beta=5^{\circ}$


Figure J.15: Stress Angle Contours(degrees):Coarse Sand, $K=0.5, L=225 \mathrm{~m}, \beta=2^{\circ}$

## Horizontal Distance (m.)



Horizontal Distance (m.)


## Horizontal Distance (m.)



Horizontal Distance (m.)


Figure J.16: Stress Angle Contours(degrees):Coarse Sand, $h^{\prime}=0.5, L=150 \mathrm{~m}, \beta=20^{\circ}$

Horizontal Distance (m.)


Horizontal Distance (m.)


Horizontal Distance (m.)


Horizontal Distance (m.)


Figure J.17: Stress Angle Contours(degrees):Coarse Sand, $K^{\prime}=0.5, L=150 \mathrm{~m}, \beta=5^{\circ}$


Figure J.18: Stress Angle Contours(degrees):Coarse Sand, $K=0.5, L=150 \mathrm{~m}, \beta=2^{\circ}$


Figure J. 19: Stress Angle Contours(degrees):Coarse Sand, $L=225 \mathrm{~m}, \omega t=\frac{\pi}{2}, K=0.7$

Horizontal Distance (m.)


Horizontal Distance (m.)


## Horizontal Distance (m.)



Horizontal Distance (m.)


Figure J.20: Stress Angle Contours(degrees):Coarse Sand, $L=225 \mathrm{~m}, \omega t=\frac{\pi}{2}, K=1.0$


Figure J. 21 : Stress Angle Contours(degrees):Coarse Sand, $L=150 \mathrm{~m}, \omega t=\frac{\pi}{2}, K=0.7$


Figure J. 22: Stress Angle Contours(degrees):Coarse Sand, $L=150 \mathrm{~m}, \omega t=\frac{\pi}{2}, \boldsymbol{h}=1.0$
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